



Advances in Space Weather using GNSS in Asia over the past decade

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OUTLINE

Introduction : ionospheric effects on GNSS signal

Regular plasma irregularities using S_4 and ROTI index

Climatology of Ionosphere using TEC data

Effects of Geomagnetic storms

- TEC
- ROTI index

THE IONOSPHERE

The ionosphere is a ionized layer around the Earth (from ~ 50 km up to 800 km).

The ionosphere is the largest source of perturbations for GNSS

Ionospheric electric currents are at the origin of variations of the Earth's magnetic field and Ground Induced Electric Currents (GIC)

Regular and irregular variations

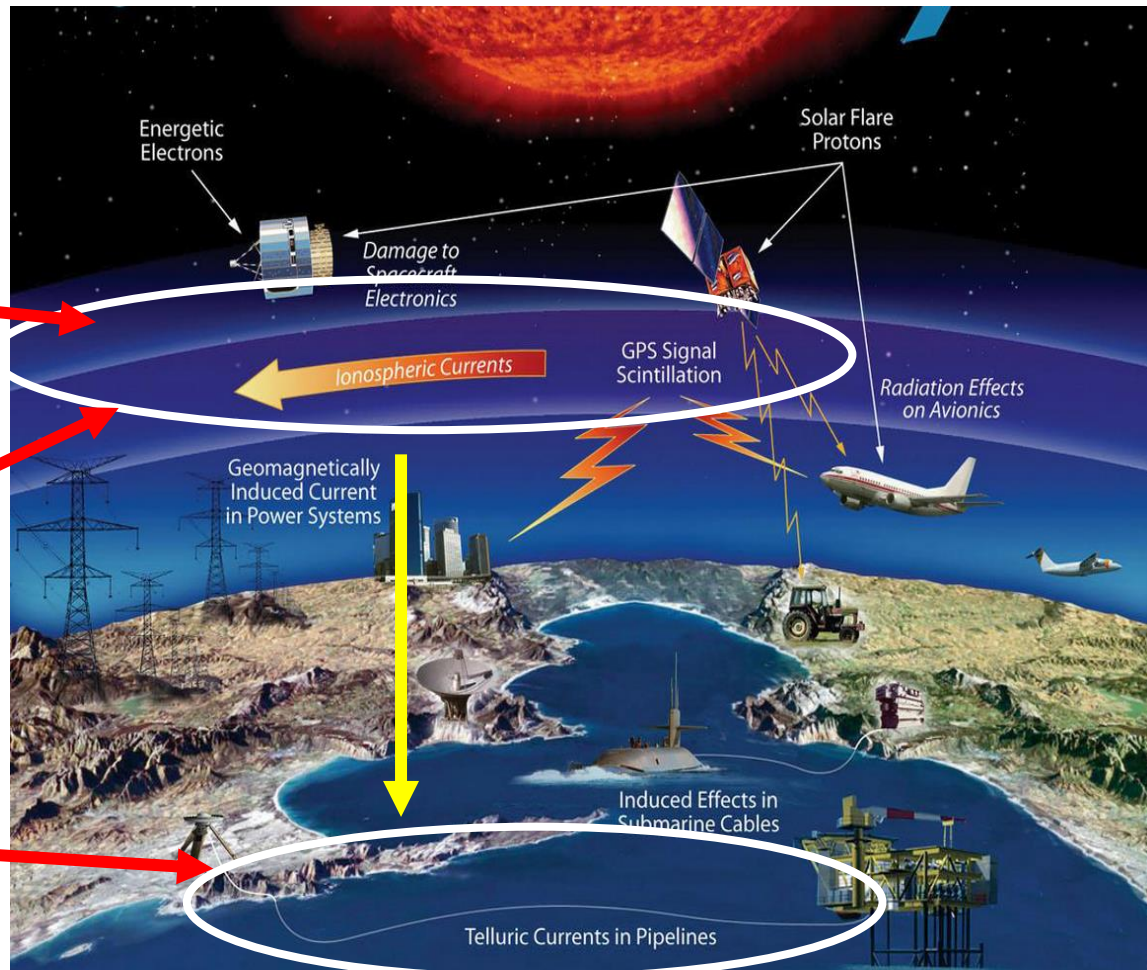
1) Ionization

Propagation of electromagnetic Waves

2) Ionospheric Electric current

3) Variations of the Earth's magnetic field and GIC

Nasa website



Effects of the ionosphere on propagation (TEC)

–Reduction of the phase path length
(with respect to propagation in vacuum)

$$\Delta P_\phi = P_\phi - L = \int_L (n - 1) ds$$
$$n = 1 - a \frac{N_e}{f^2}$$

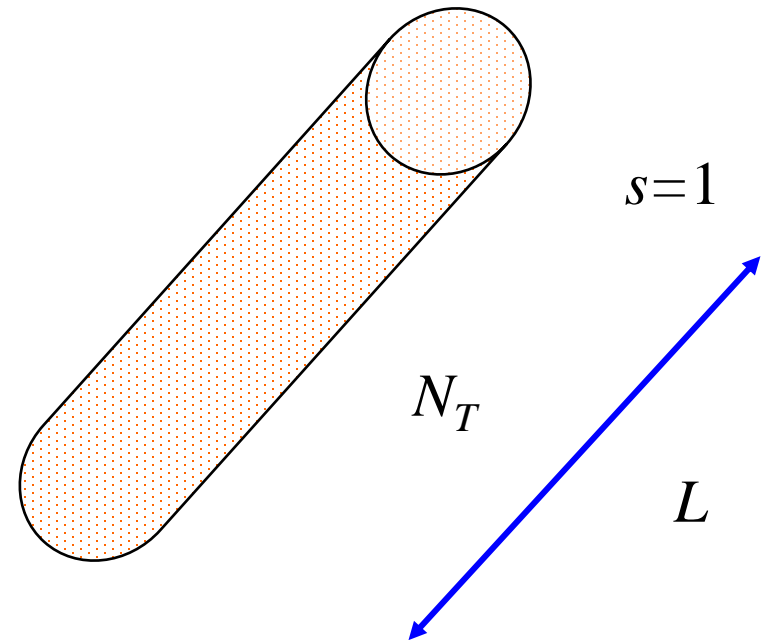
$$\Delta P_\phi = -\frac{a}{f^2} \int_L N_e ds$$

$$\Delta P_\phi = -a \frac{N_T}{f^2}$$

Phase path length : Distance that a wave needs to propagate in a vacuum to have the same total phase shift (ϕ)

– Total Electron Content (TEC)

$$N_T = \int_L N_e dl$$

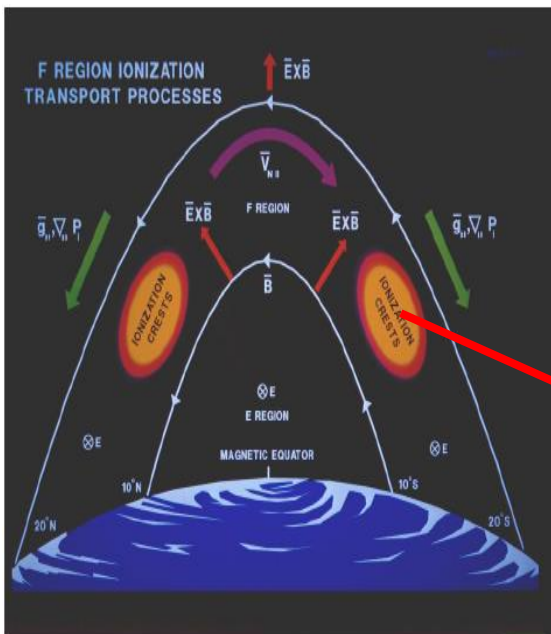


$$1 \text{ TECU} = 10^{16} \text{ electron/m}^2$$

Scintillations a regular phenomenon

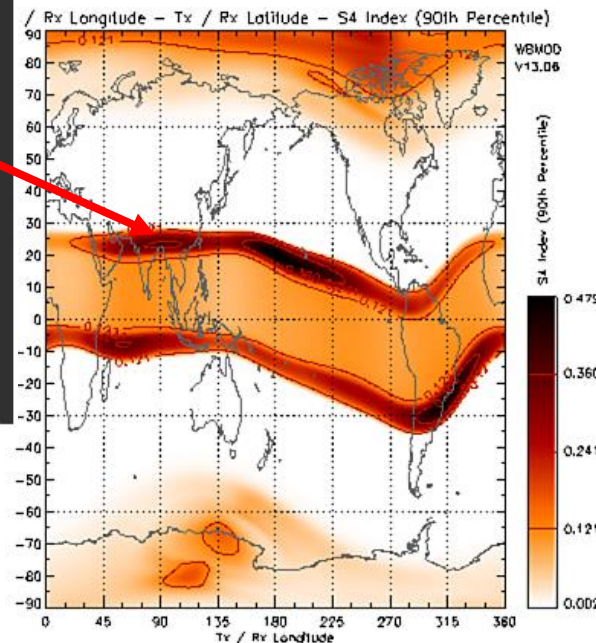
Ionospheric scintillation is the rapid modification of radio waves caused by small scale structures in the ionosphere **Physical Process : Instabilities in Plasma**

At the magnetic equator the Earth's magnetic field is horizontal



**Equatorial Fountain
EIA**

$$S_4 = \sqrt{\frac{\langle I^2 \rangle - \langle I \rangle^2}{\langle I \rangle^2}}$$



Scintillation index at GPS L1 (1575.42 MHz) assuming constant local time 23.00 at all longitudes (from <http://www.sws.bom.gov.au>)

During the daytime the east–west electric field and the north–south geomagnetic field produce the lift of plasma in E ionospheric region by **vertical E X B drift**. At higher altitudes in F region, the plasma diffuses downward along the geomagnetic field lines into both hemispheres under the influence of gravity and pressure gradients, this produces the EIA which is characterized by an electron density trough at the magnetic equator, and two crests of enhanced electron density at about $\pm 15^\circ$ magnetic latitude

PLASMA IRREGULARITIES

S4 -> fluctuations of the GPS power signal

$$I = \frac{A^2}{2}$$

$$s4 = \sqrt{\frac{\langle I^2 \rangle - \langle I \rangle^2}{\langle I \rangle^2}}$$

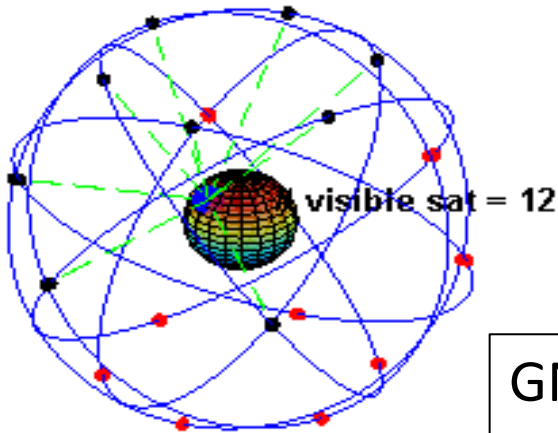
ROTI Index

$$\text{rot} = \frac{STEC_{k+1} - STEC_k}{time_{k+1} - time_k} * 60$$

$$\text{roti} = \sqrt{\langle \text{rot}^2 \rangle - \langle \text{rot} \rangle^2}$$

Training on daily global positioning system GPS data , Coordinates a monthly magazine on positioning, navigation and beyond, <http://www.mycoordinates.org>, **Volume XIII, Issue 03, March 2017** (Amory-Mazaudier, Rolland Fleury, Sharafat Gadimova, Abderrahmane Touzani)
Software of Rolland Fleury on www.girgea.org

During Solar Disturbances the propagation of electromagnetic waves is disturbed by changes in ionization => GNSS is affected
Global Navigation Satellite System (GPS, GLONASS, GALILEO, IRNSS, BIEDOU etc...)



Daily life



GNSS is used for many applications of the daily life: car guidance, civil aviation, agriculture geography, epidemiology etc

GNSS is a valuable tool for various research concerning the Earth's environment: Space Weather, plate tectonics, atmosphere etc...

We focus on Space Weather and mainly on TEC, S4 and ROTI parameters, concerning ionosphere.

We will only present some effects of the sun on the ionosphere.

the GNSS technique makes it possible to approach many new fields of research in different countries

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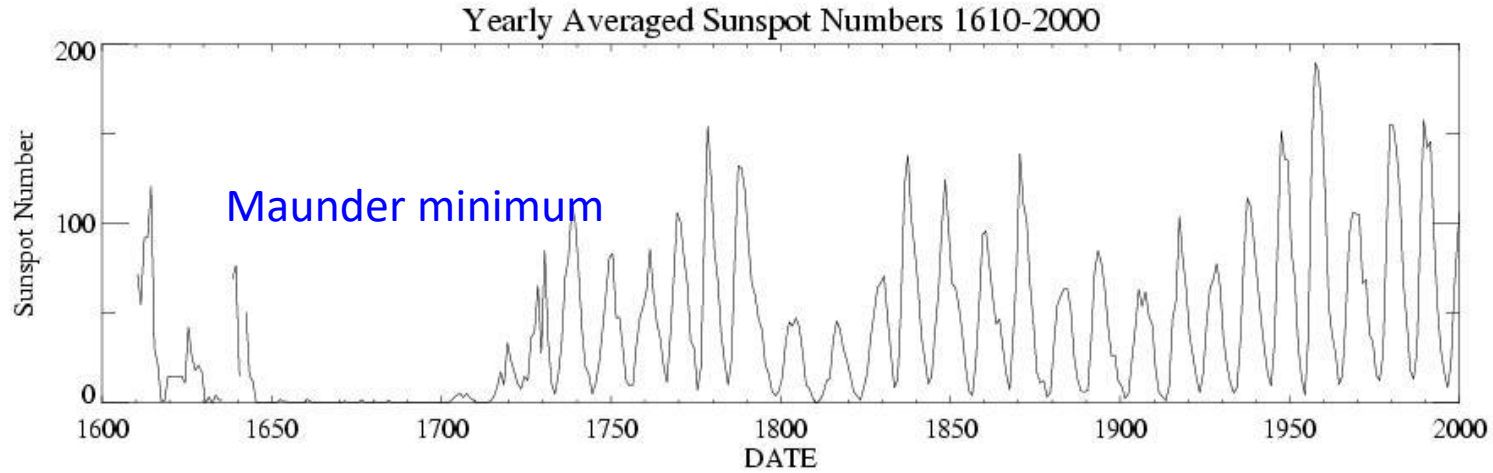
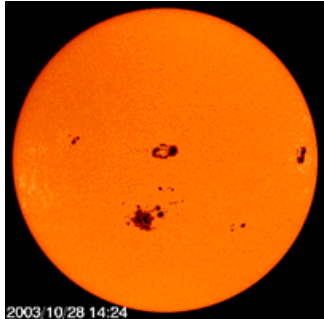
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PHYSICS

Solar electromagnetic emissions/ Sunspot cycles



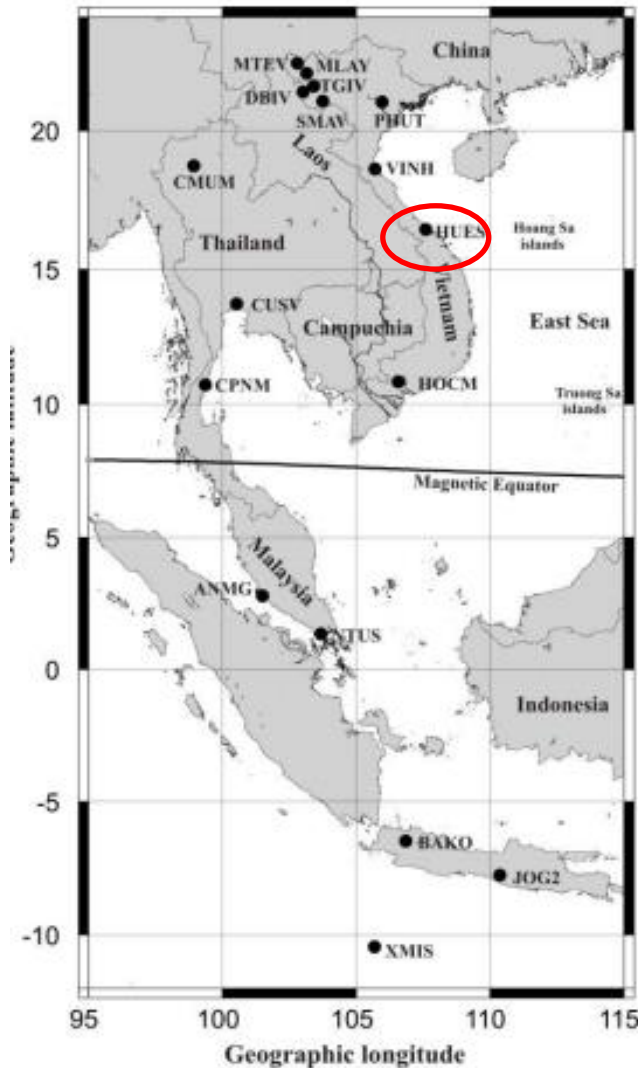
Solar cycle 24



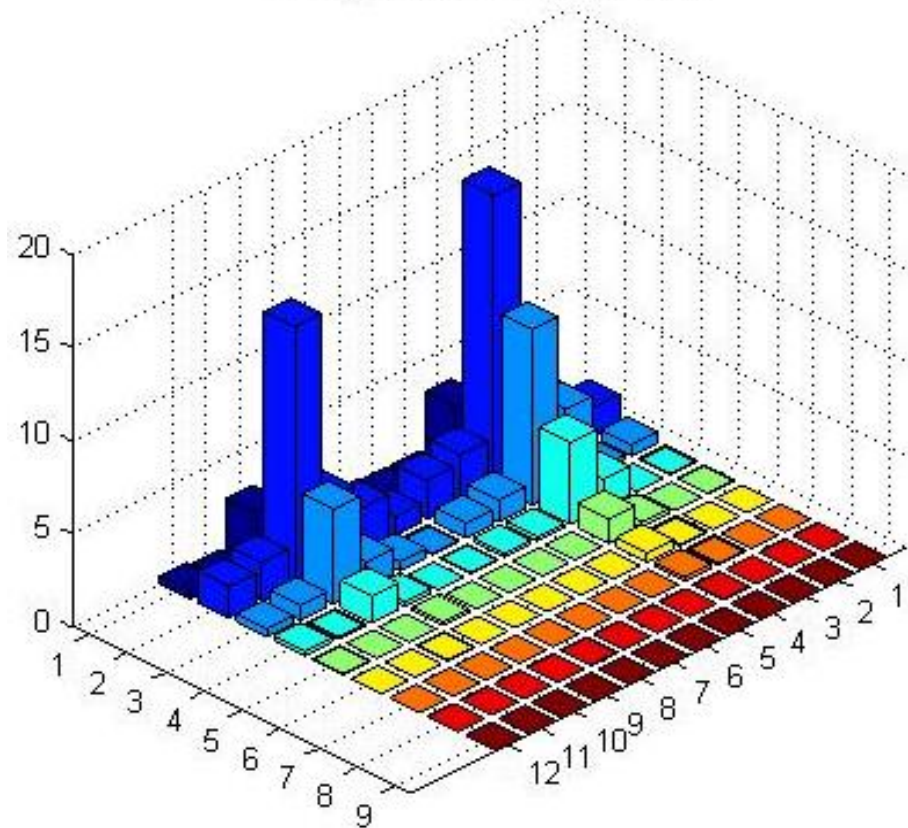
Scintillation index S4 observed at Hue (Vietnam) from 2006 to 2008 -> fluctuations of the GPS power signal

$$I = \frac{A^2}{2} \quad s4 = \sqrt{\frac{\langle I^2 \rangle - \langle I \rangle^2}{\langle I \rangle^2}}$$

Histogram of s4 vs month - Hue



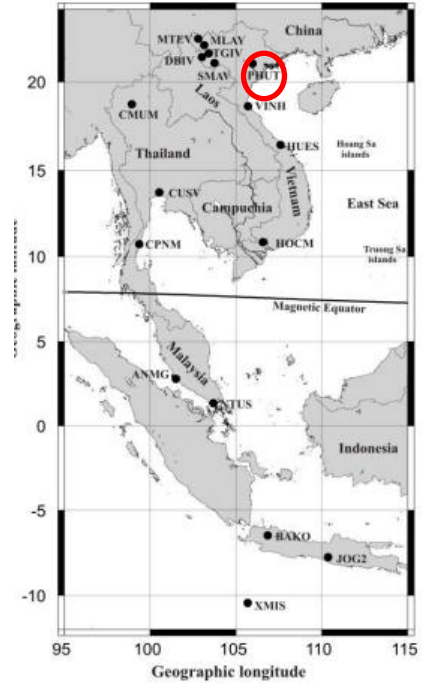
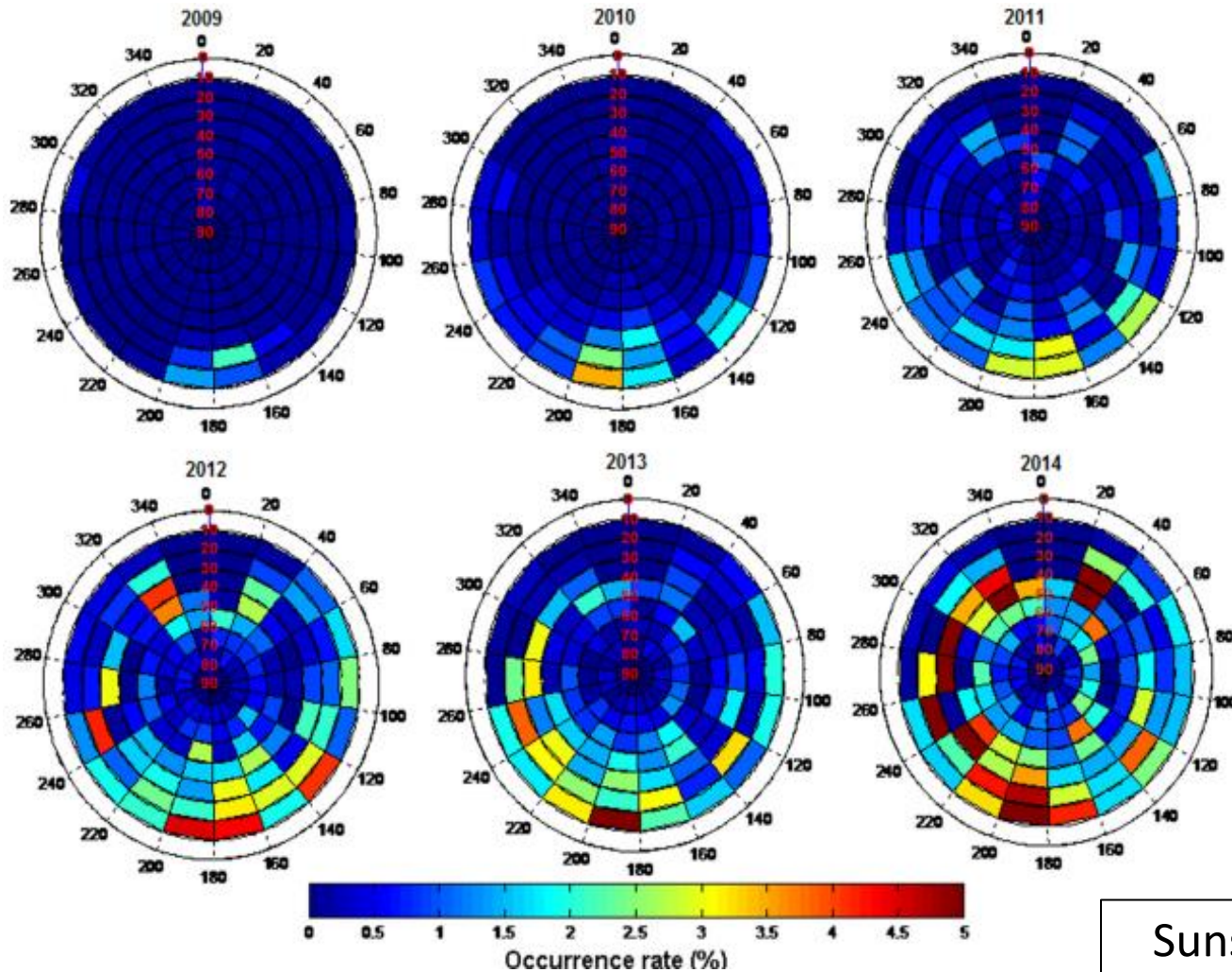
Distribution of GPS receivers in Vietnam
And adjacent region



Seasonal variations: equinox maximum

The directional distribution of scintillations observed from PHUT station during 2009–2014.

VIETNAM

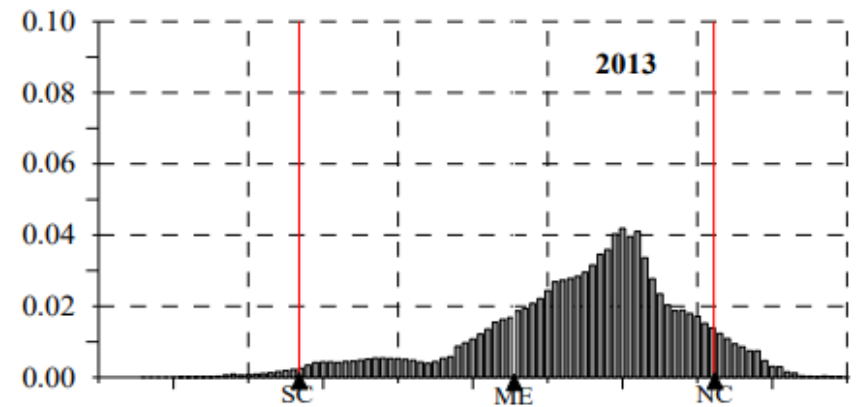
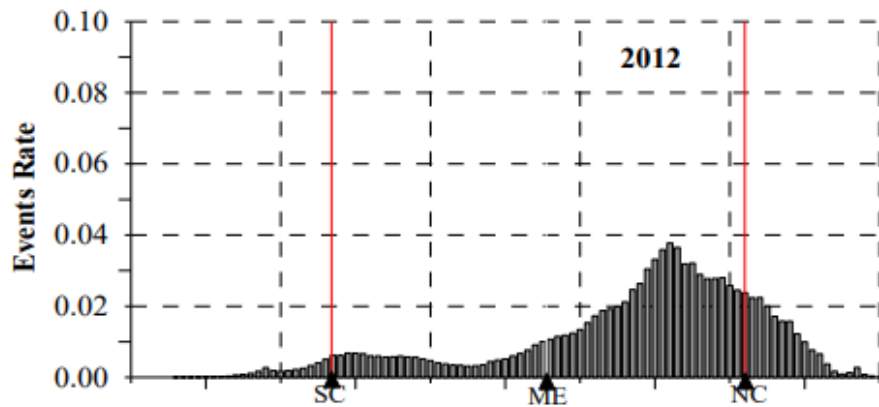
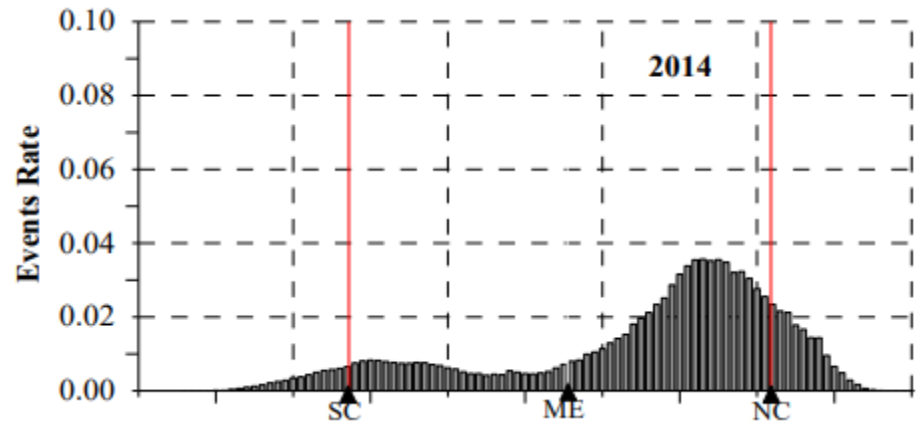
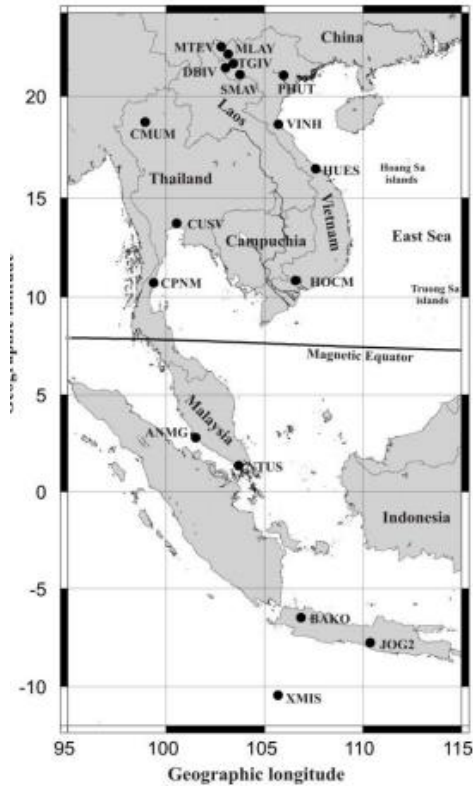


Sunspot cycle Maximum
2012,2013,2014

Statistics of ionospheric irregularities ($ROTI \geq 0.5$) along geographic latitude during the period 2008-2018.

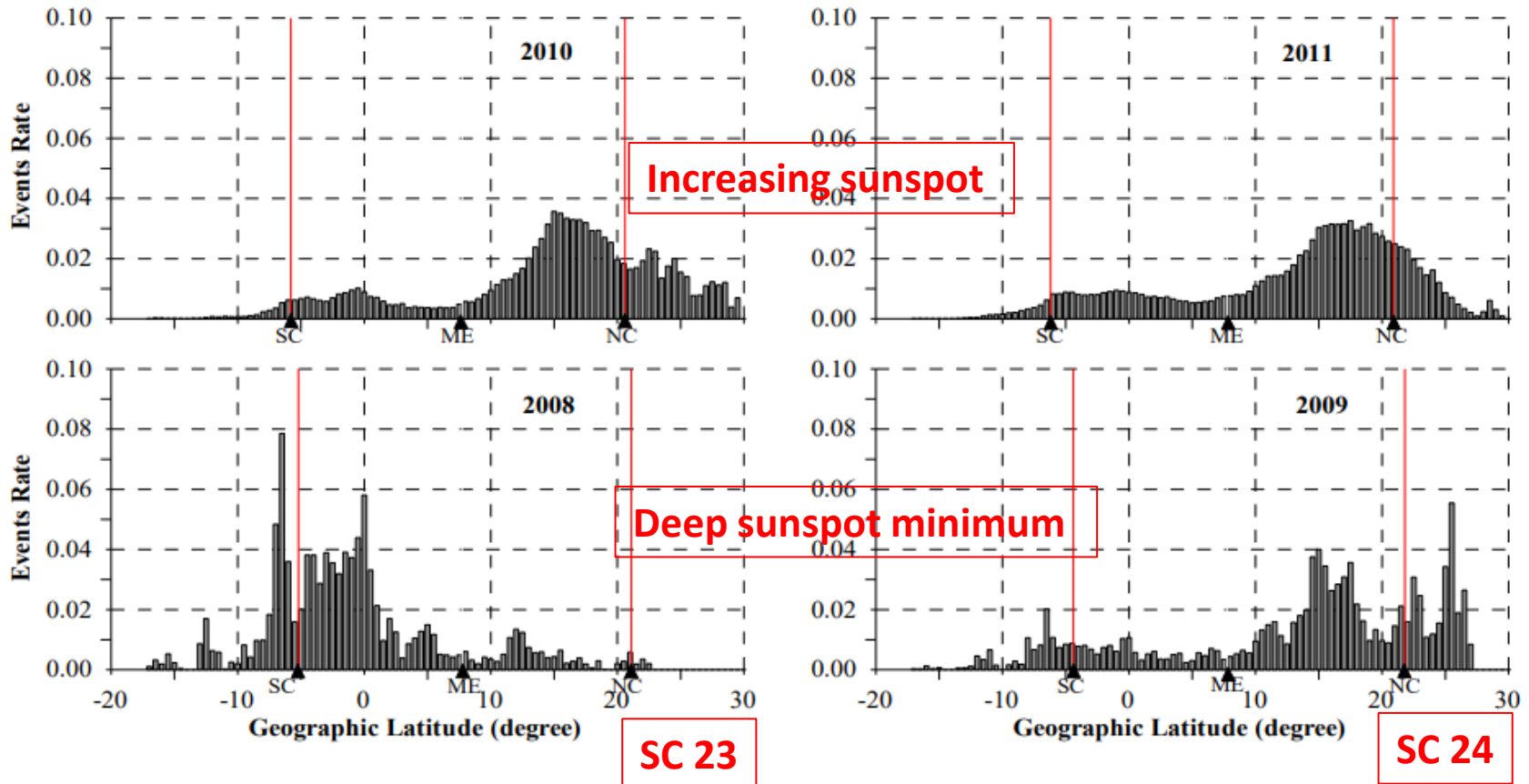
Phase maximum of sunspot cycle 24

ASIAN SECTOR



Statistics of ionospheric irregularities ($ROTI \geq 0.5$) along geographic latitude during the period 2008-2018. ME: magnetic equatorial; SC: southern crest; NC: northern crest

ASIAN SECTOR

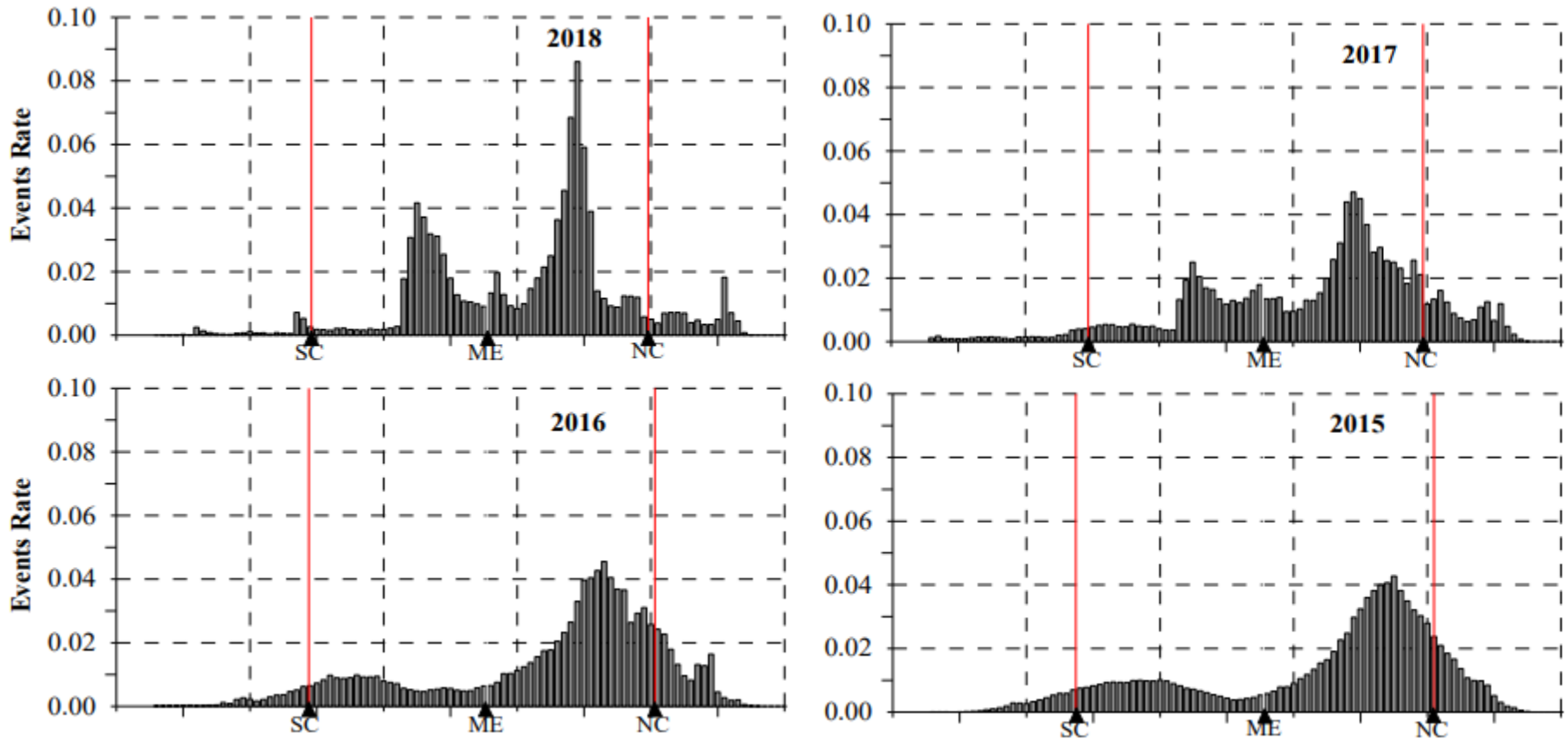


Dung Nguyen Thanh, Minh Le Huy, Christine Amory-Mazaudier, Rolland Fleury, Susumu Saito, Thang Nguyen Chien, Hong Pham Thi Thu, Thanh Le Truong, Mai Nguyen Thi, Characterization of ionospheric irregularities over Vietnam and adjacent region for the 2008-2018 period, Vietnam Journal of Earth Sciences, 1-20, <https://doi.org/10.15625/2615-9783/16502>

Statistics of ionospheric irregularities ($\text{ROTI} \geq 0.5$) along geographic latitude during the period 2008-2018. ME: magnetic equatorial; SC: southern crest; NC: northern crest

ASIAN SECTOR

Decreasing phase of sunspot cycle 24



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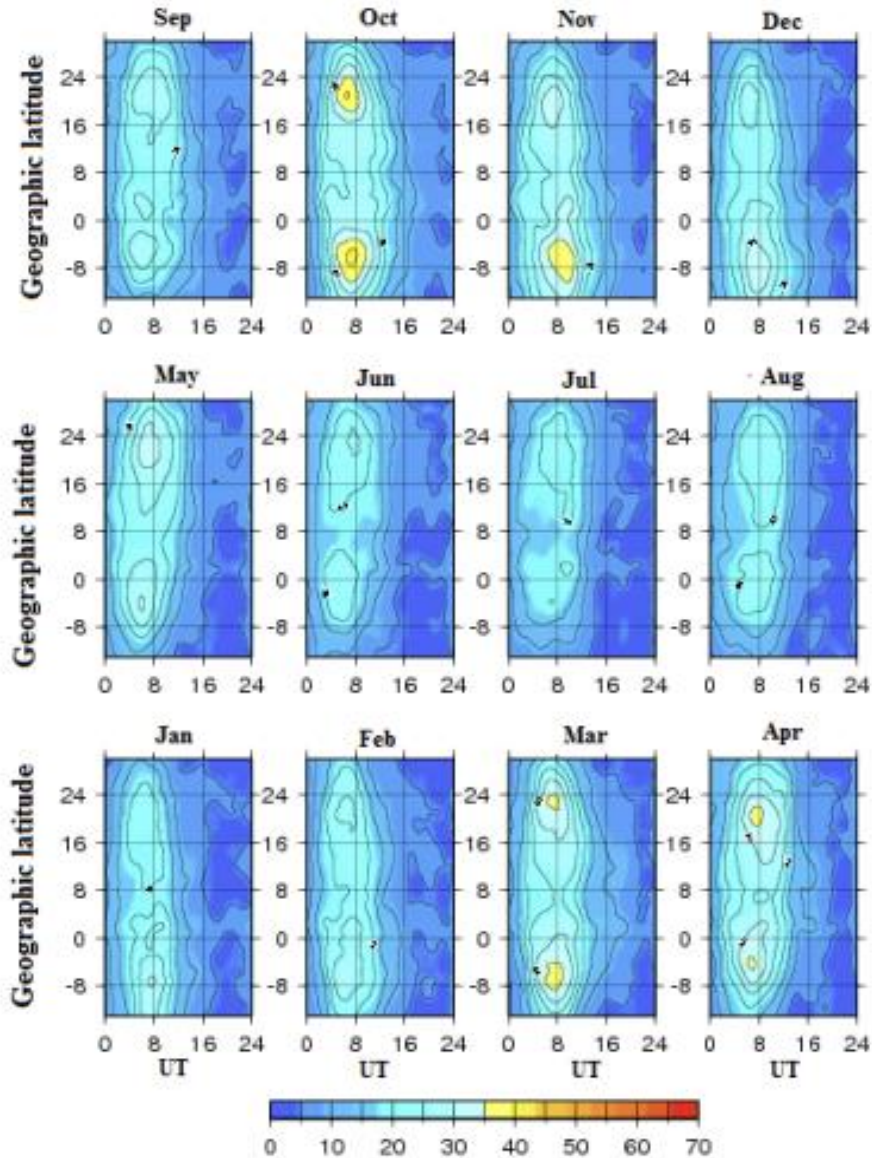
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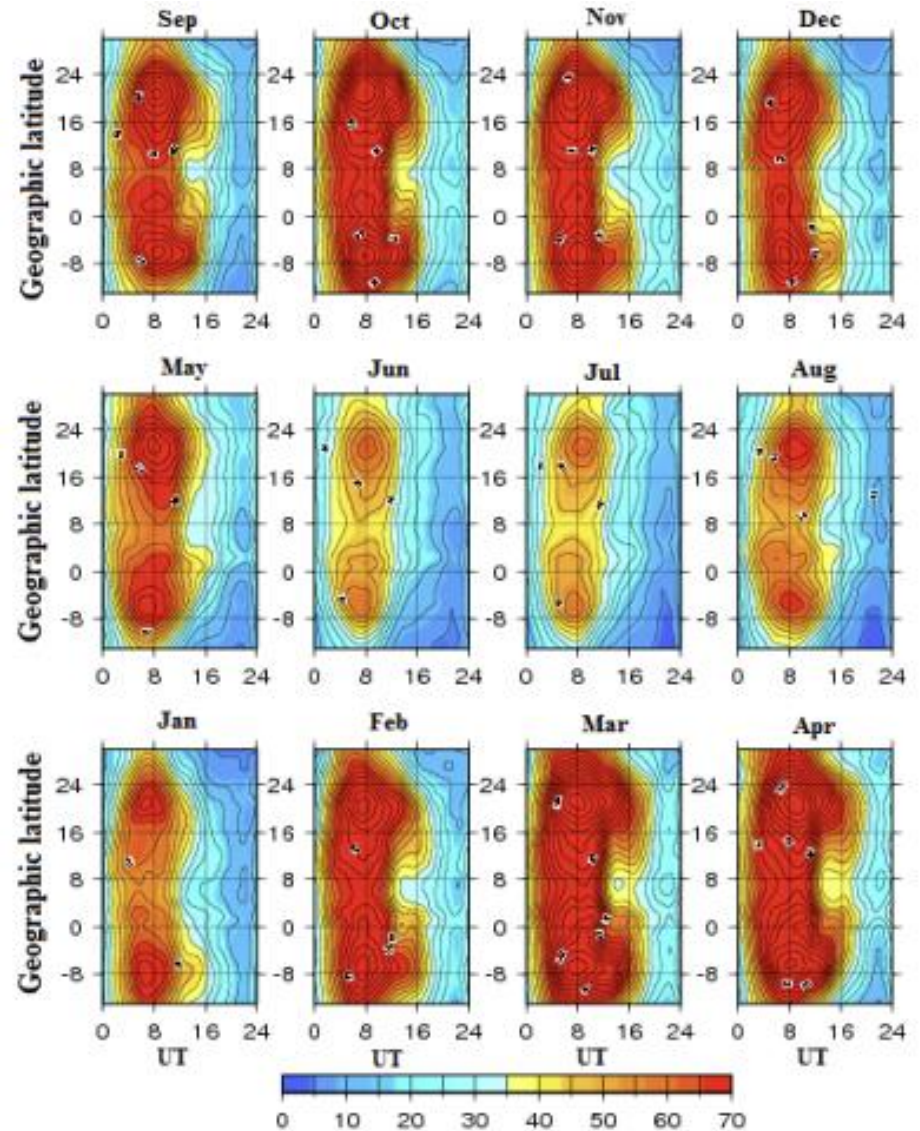
- TEC
- ROTI index

MAPS of Total Content Electron

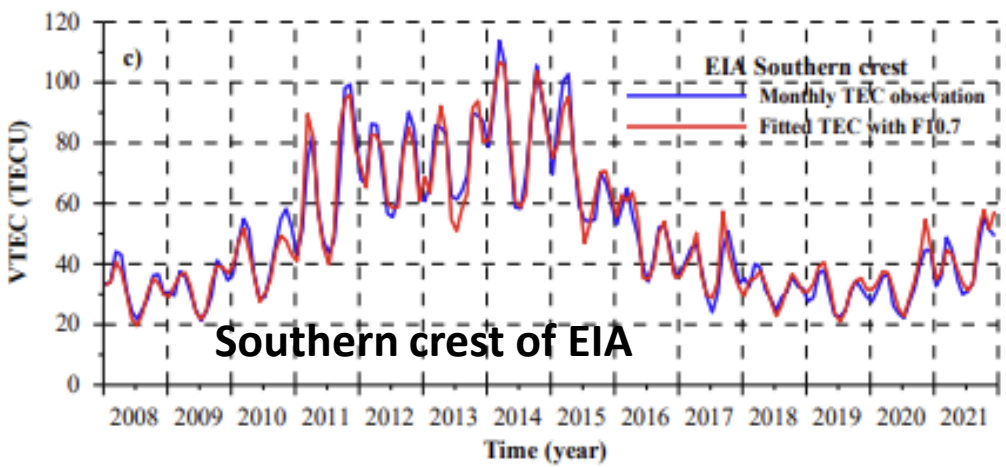
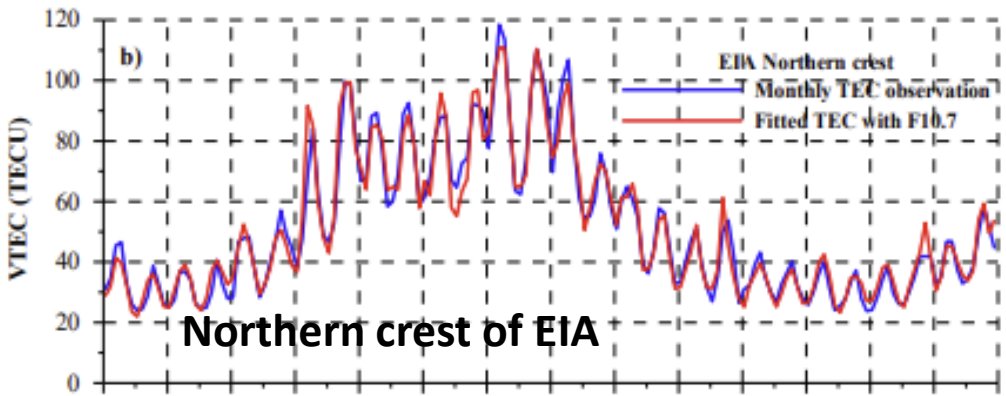
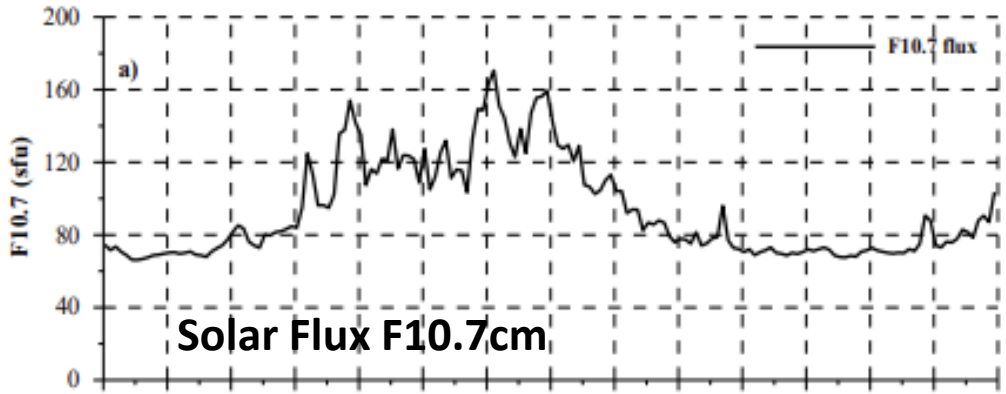
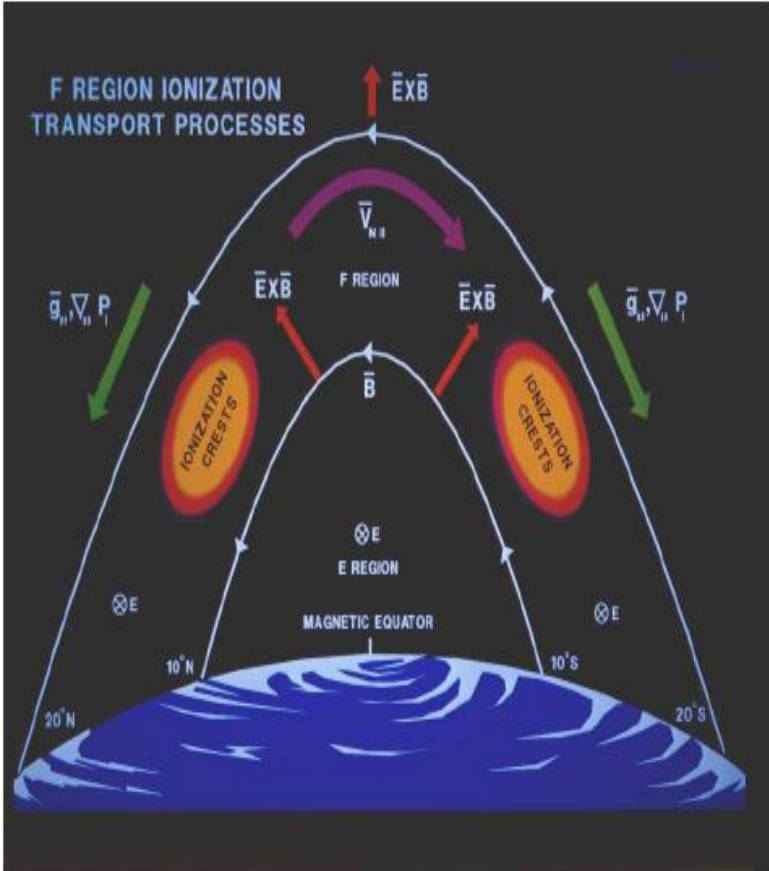
Minimum sunspot cycle -2009



Maximum sunspot cycle-2014

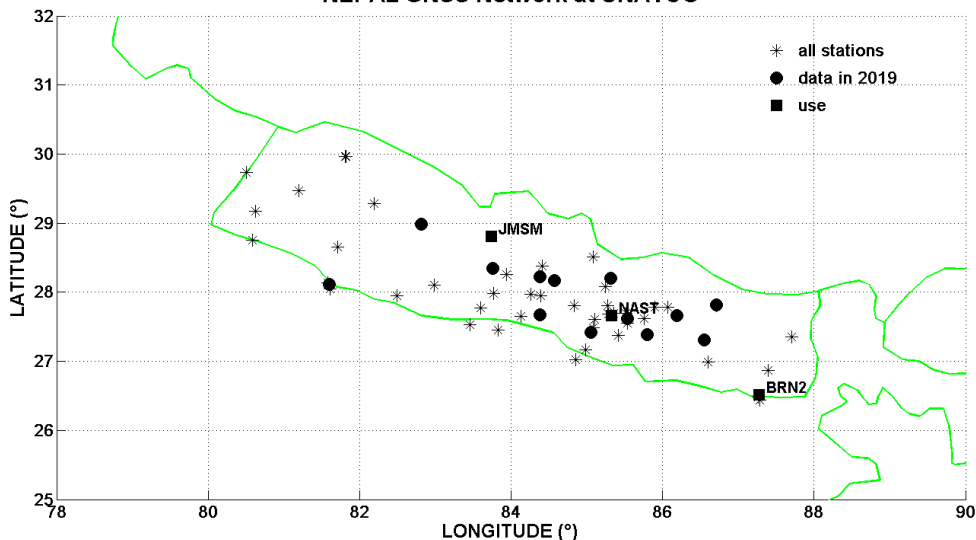


Equatorial Ionization Anomaly Equatorial Fountain



NEPAL / CAPACITY BUILDING/SPACE WEATHER

NEPAL GNSS Network at UNAVCO



ICG workshop, December 2016



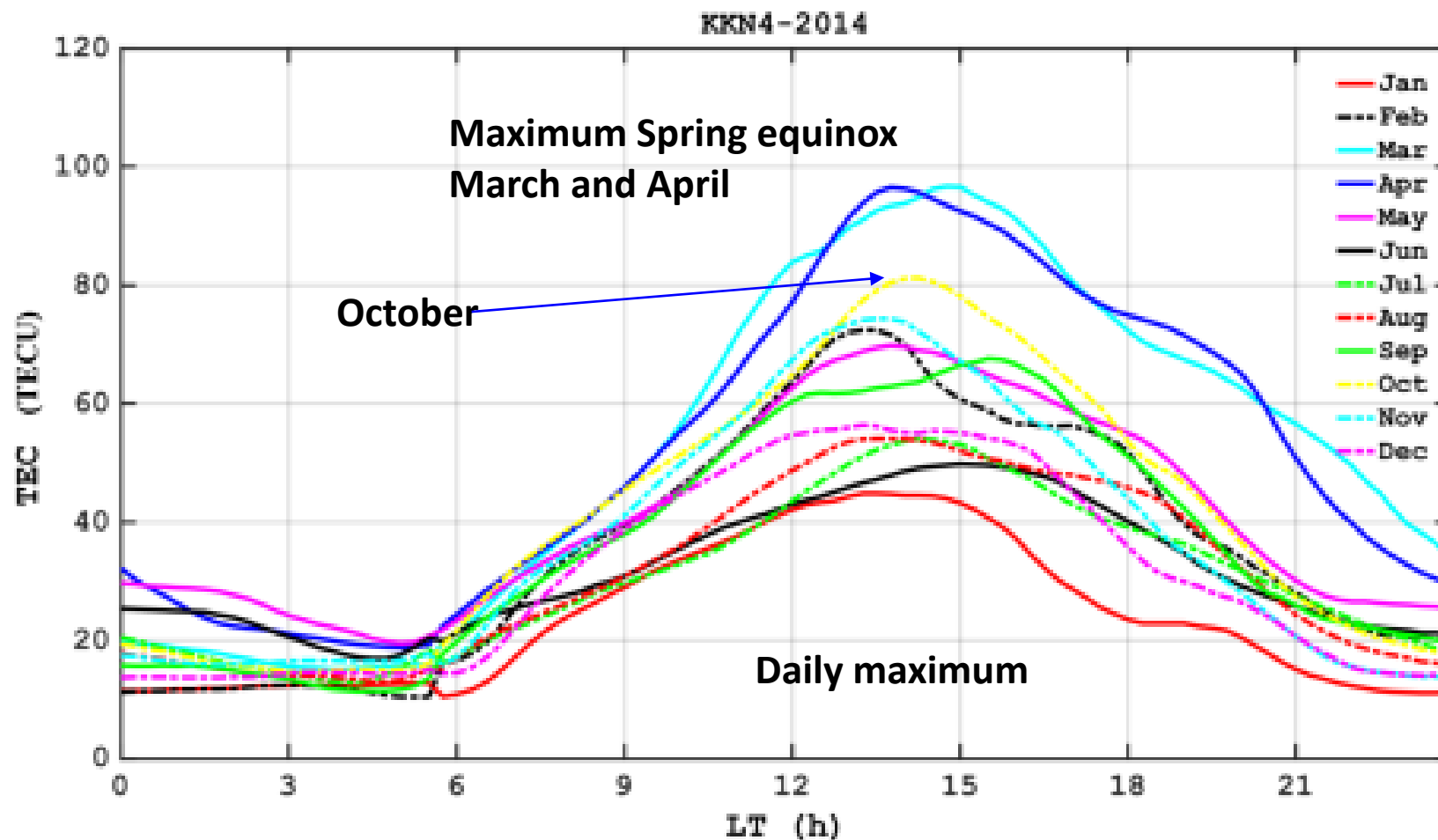
School organized by ICTP- September 2019

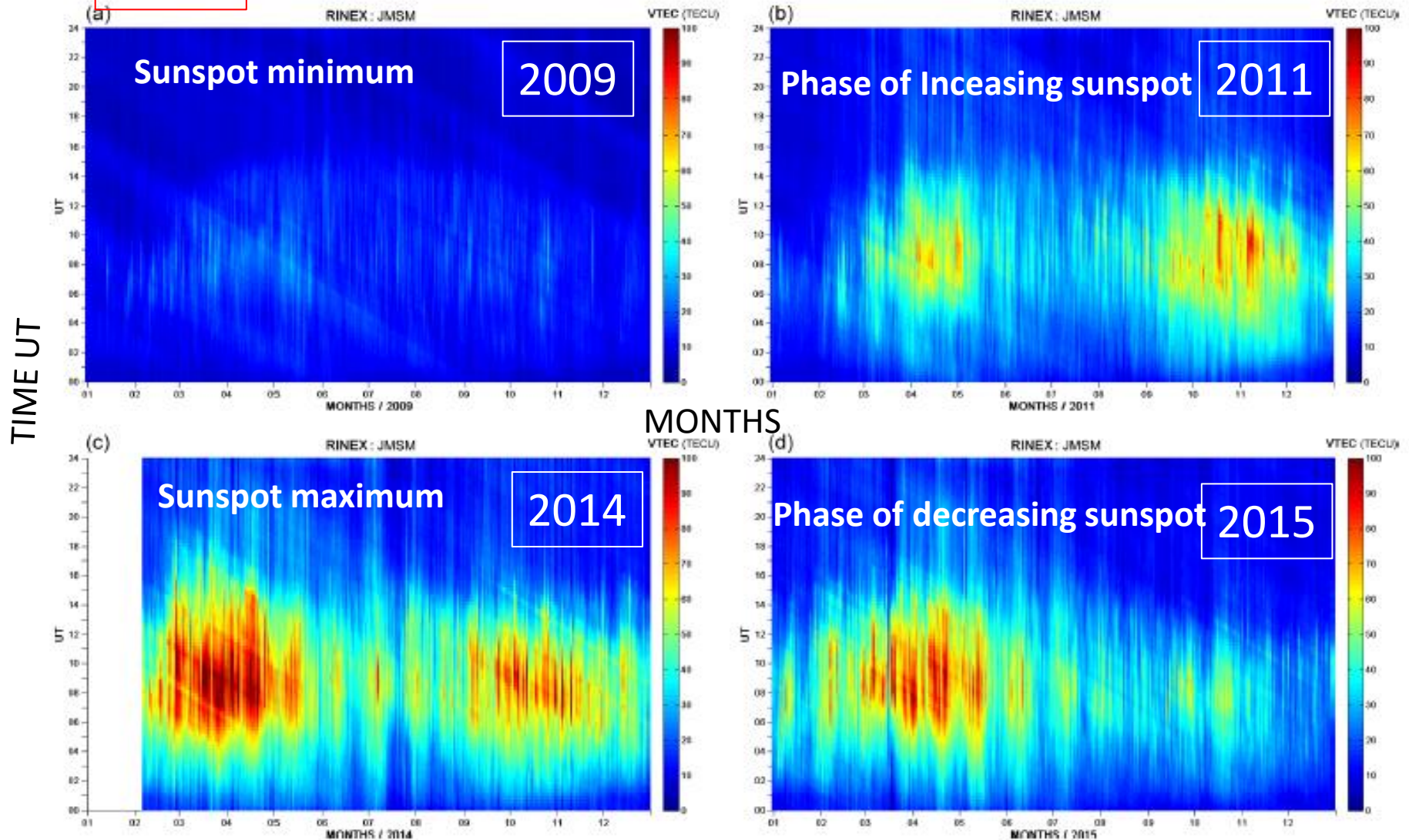


PhD of D. Pandit on September 23, 2022

MONTHLY Variations

Monthly variation in vertical TEC in LT for 2014 at KKN4 station.

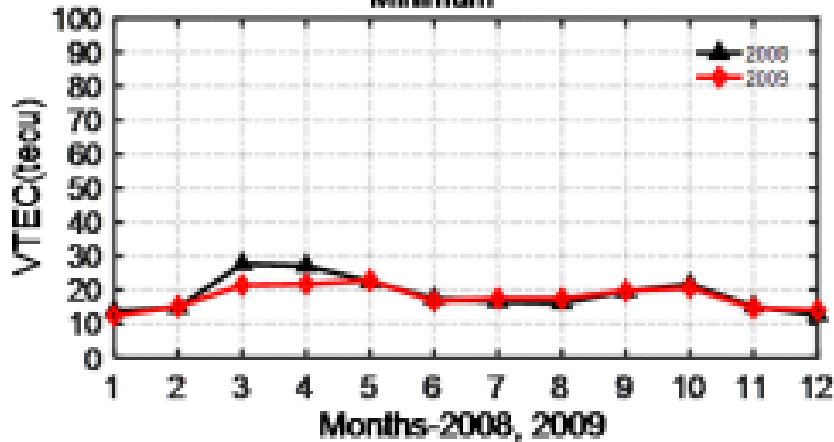




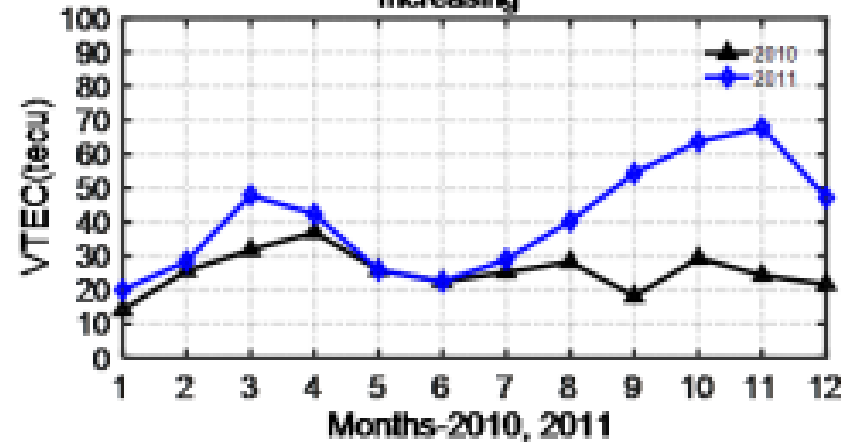
(a–d) A two-dimensional (2D) variation in vertical TEC according to UT at the JMSM station for one of the years of the minimum (2009), ascending (2011), maximum (2014) and descending (2015) phases of solar cycle 24. Pandit, D et al. <https://doi.org/10.5194/angeo-39-743-2021>

PHASES OF SUNSPOT CYCLE 24

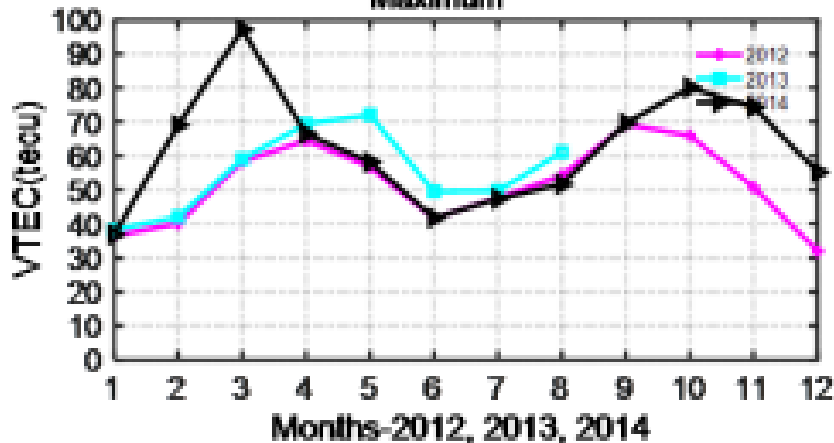
Minimum



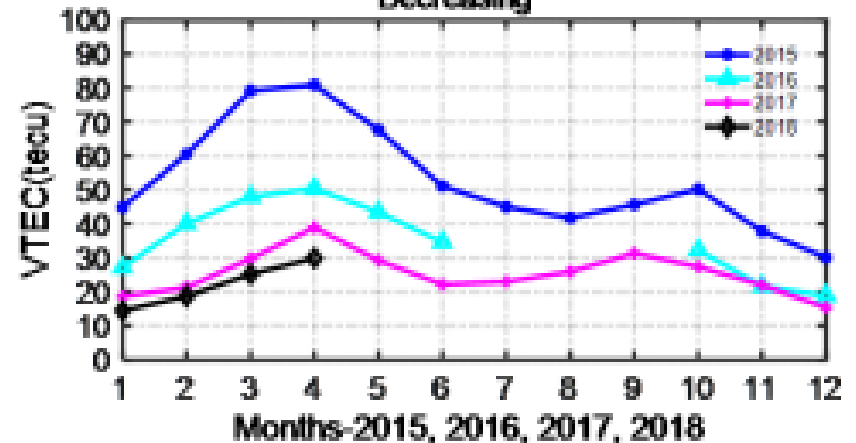
Increasing



Maximum

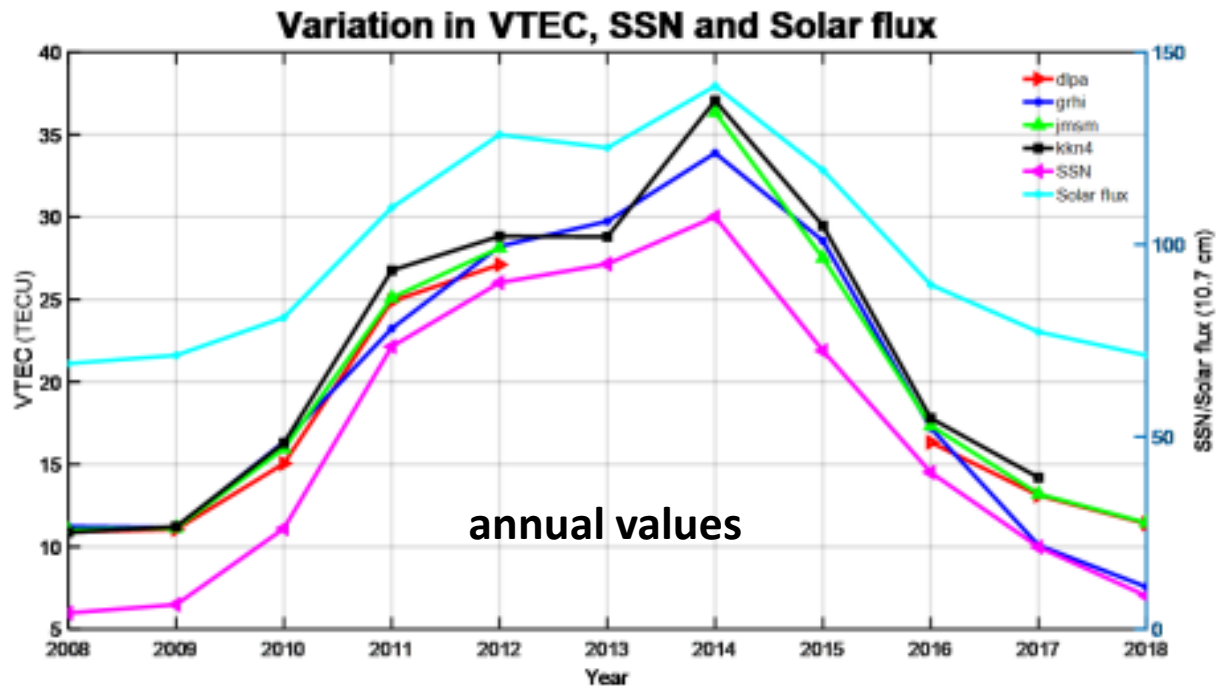
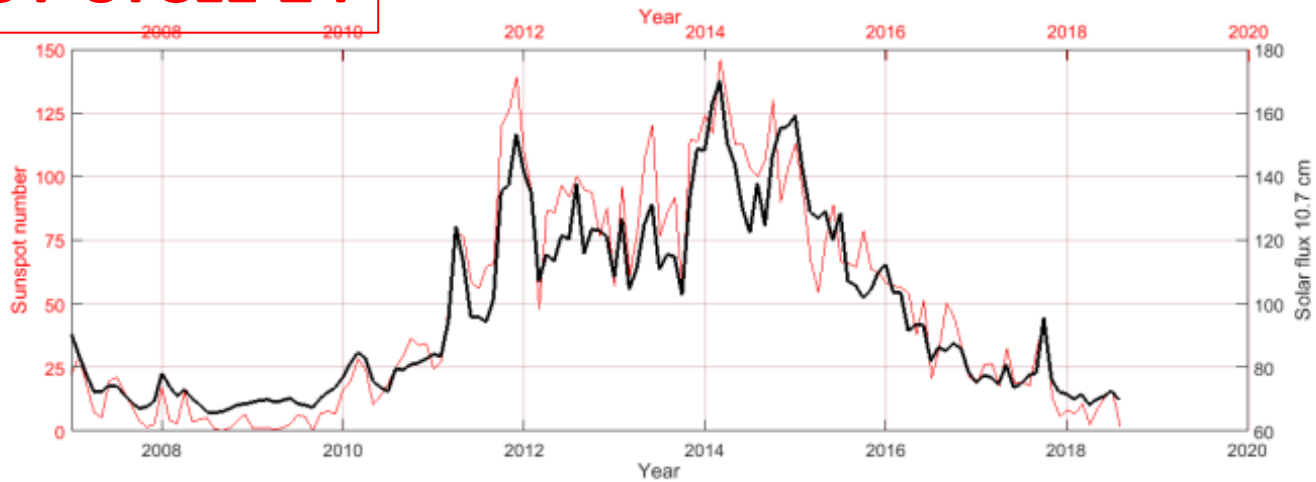


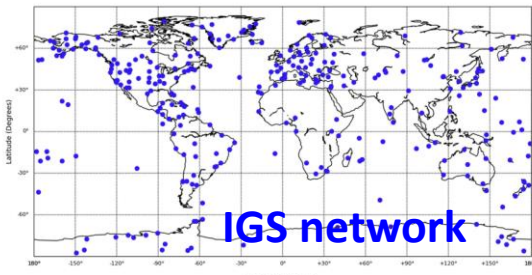
Decreasing



VTEC variability in GRHI station during minimum, increasing, maximum and decreasing phases of solar cycle 24

SUNSPOT CYCLE 24

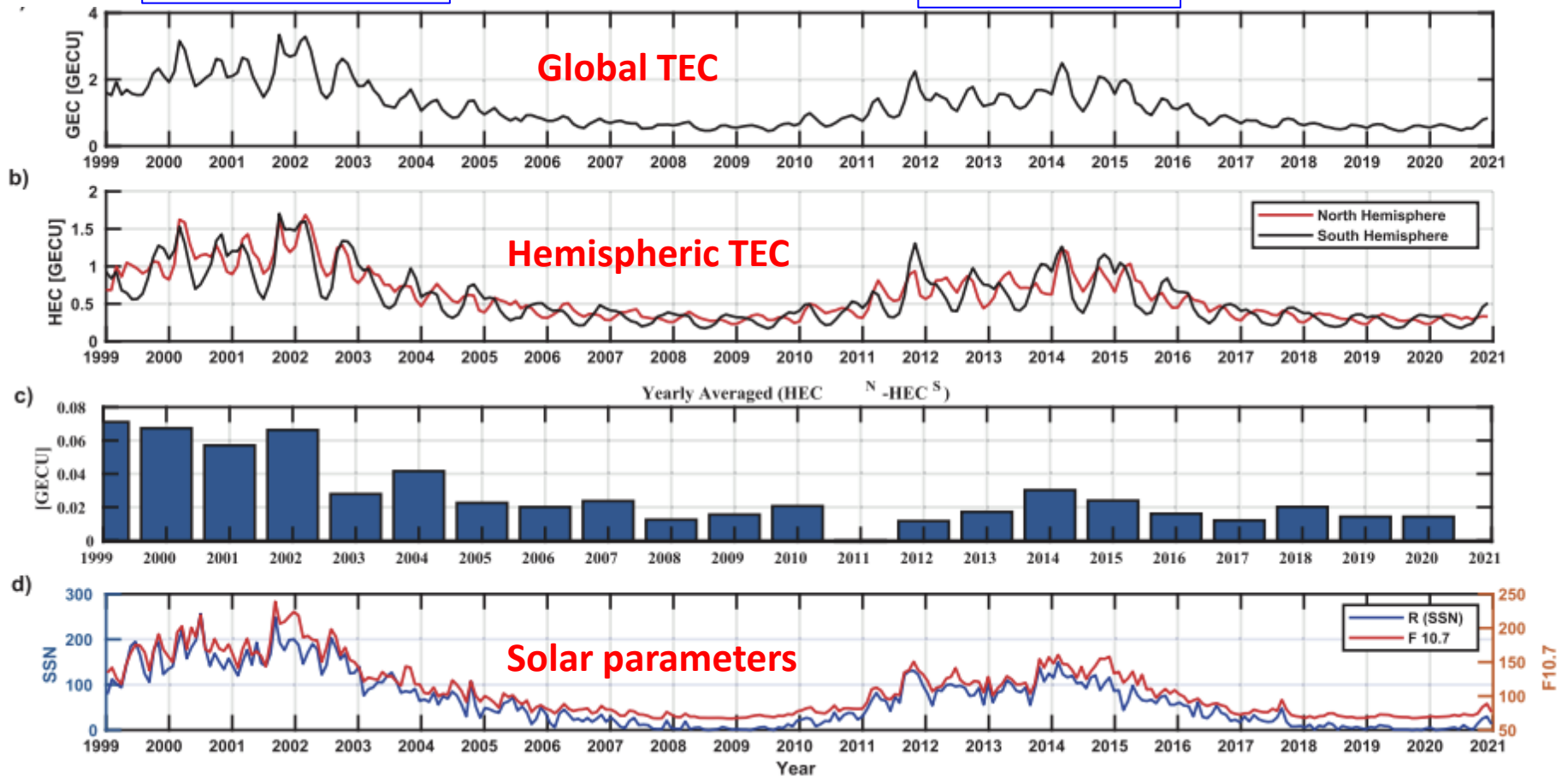




Climatology of global, hemispheric and regional electron content variations during the solar cycles 23 and 24

Solar cycle 23

solar cycle 24



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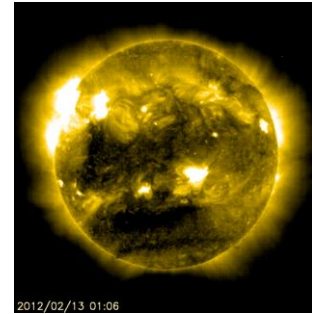
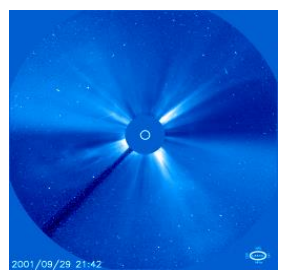
Effects of Geomagnetic storms

- **TEC**

- **ROTI index**

SOLAR EVENTS

Coronal Mass Ejection and Magnetic cloud
High Speed Solar Wind HSSW and CIR



If the Interplanetary Magnetic Field , IMF field is opposite to the terrestrial magnetic field, i.e directed toward the South, there is reconnection between the IMF and the Earth's magnetic field and **there is a magnetic storm**

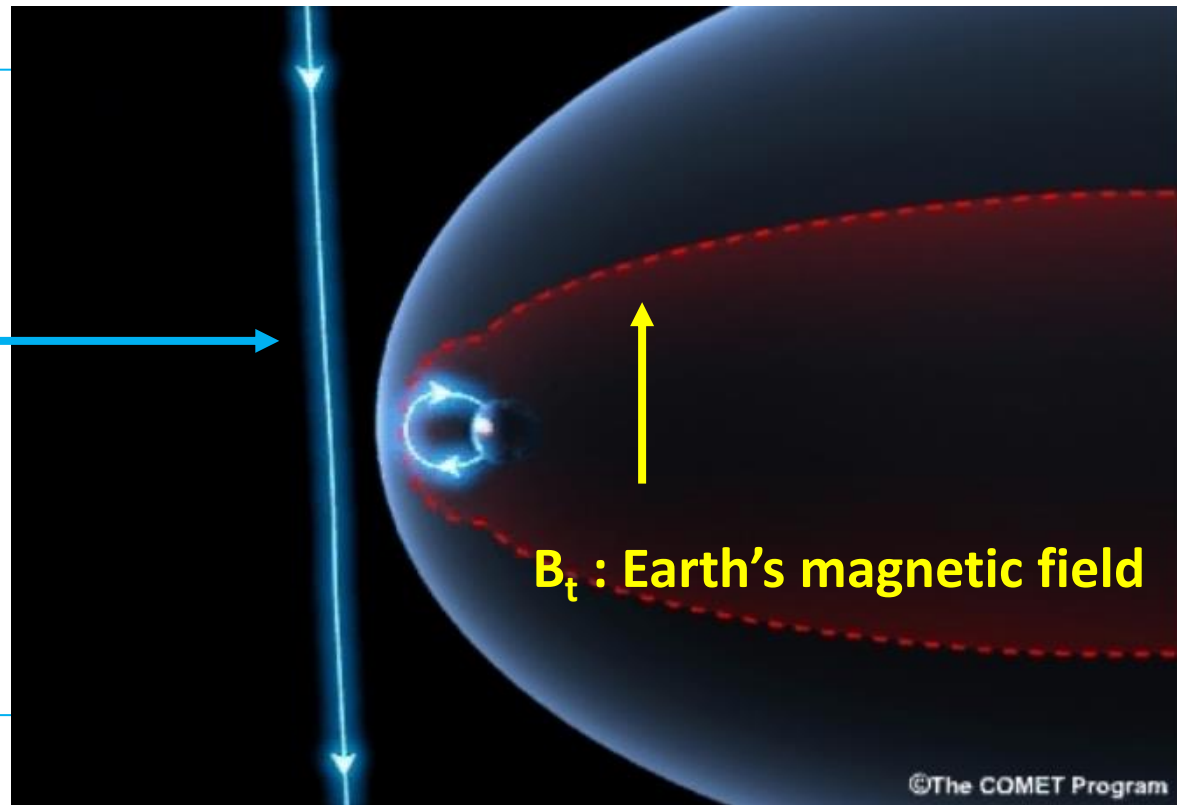
Key parameters for Space Weather

B_z IMF

V_s : solar wind speed

$$E_y = -V_x \cdot B_z$$

Solar wind is a flow of particules emitted from the sun continuously



Solar wind – Magnetosphere Dynamo : $E=V_s \times B$ movement is converted into electrical energy

Coupling between high and low latitudes

There are many physical processes acting in aurora region

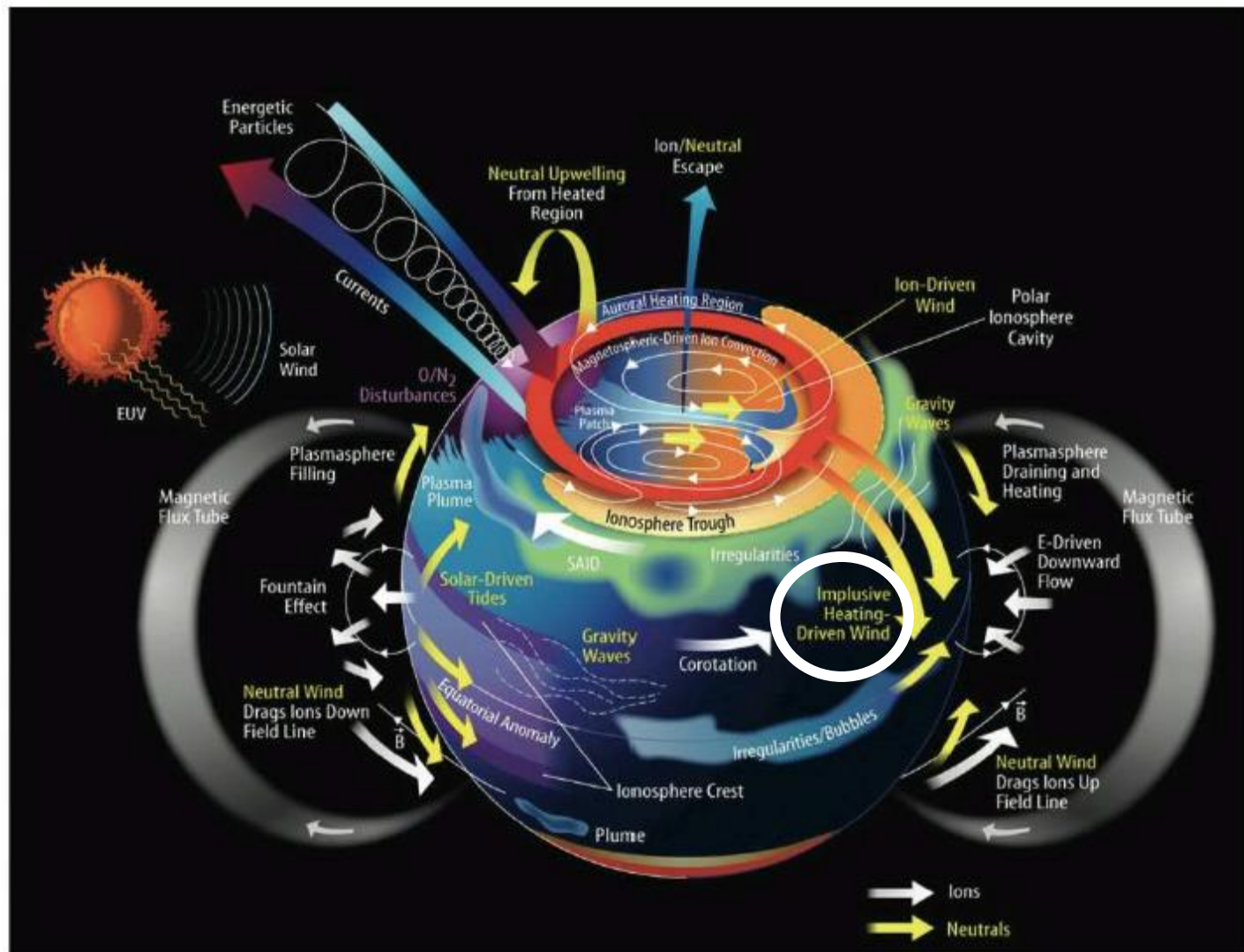


Figure 1. Known processes in the coupled MIT system. Credits: NASA's Scientific Visualization [tps://svs.gsfc.nasa.gov/4641](https://svs.gsfc.nasa.gov/4641)

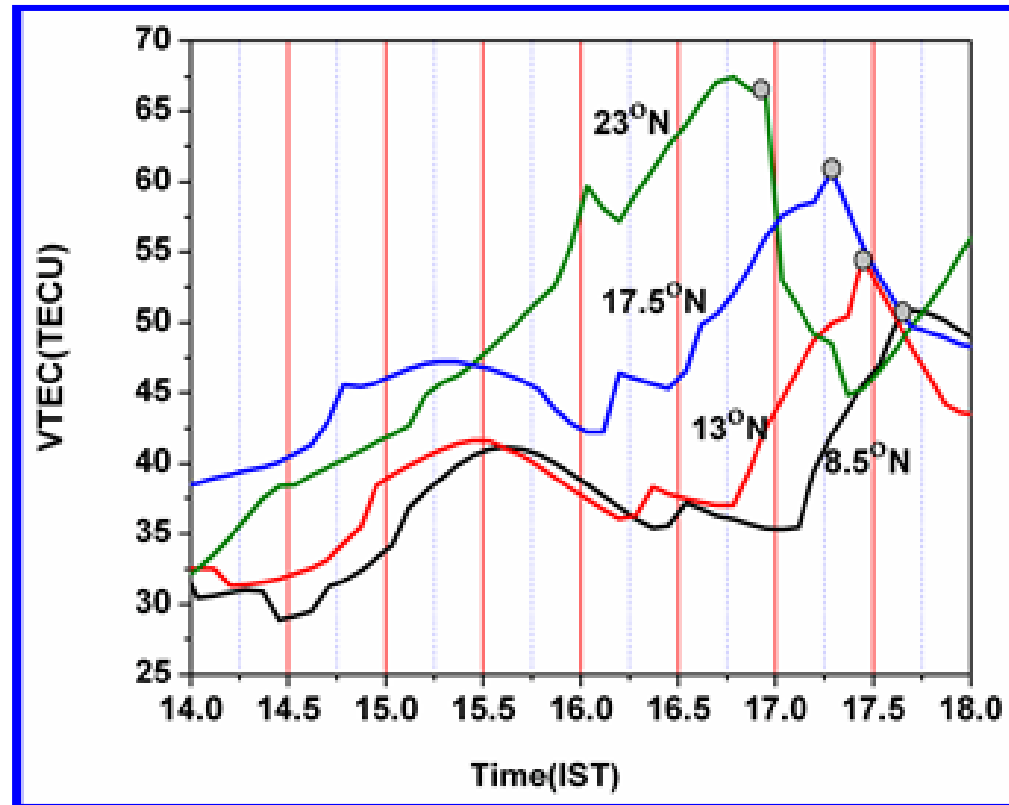
Thermal expansion of the atmosphere: Travelling Atmospheric disturbance (TAD's) => disturbed TEC [Theory Fuller Rowell et al., (1994), (1996)]

24/08/2005

SSC : 13.00 UT
Main Phase : 16 00 UT

INDIA
77-78°E meridian

V~750m/s

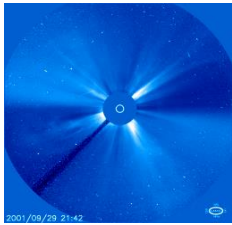


A time delay in the VTEC variations over the different latitudes indicates a propagation of TAD's Velocity 750m/s

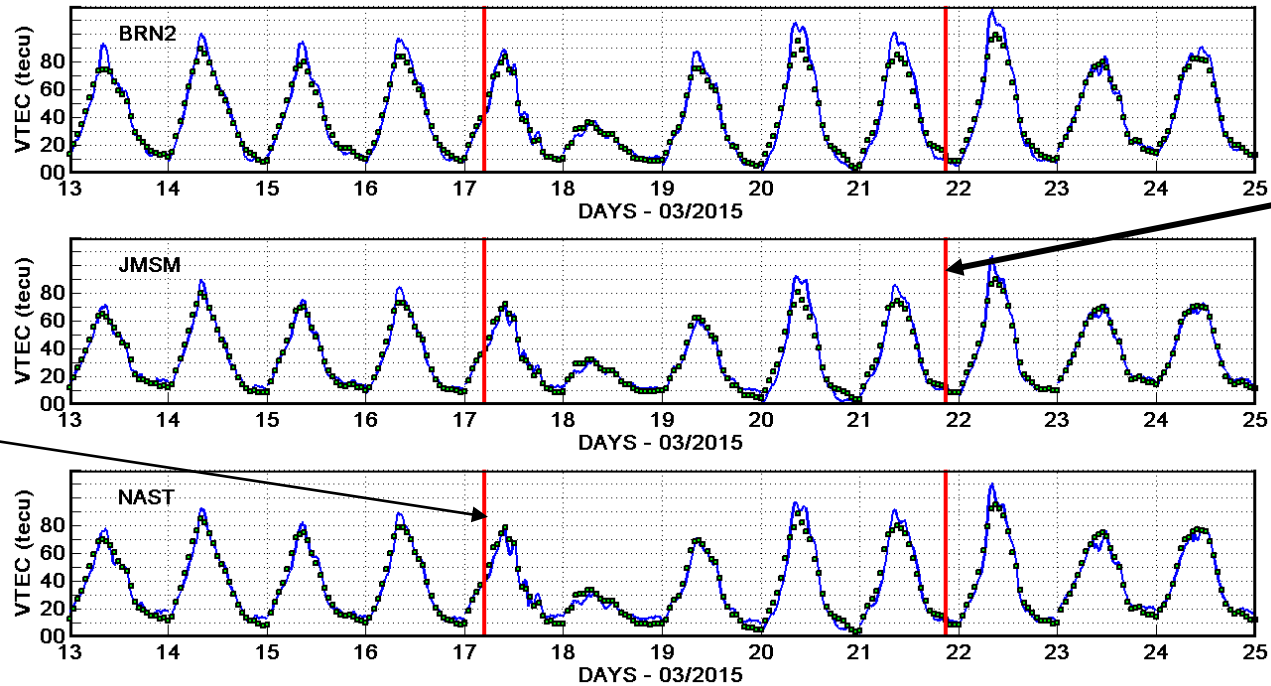
Sreeja et al., JGR vol 114, A12307, 2009

Variations of VTEC over NEPAL, during the period 14–24 March 2015, plotted using CODG TEC respectively.

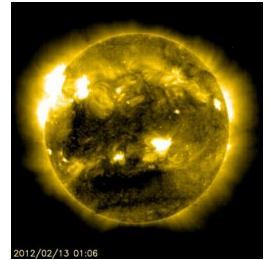
CME



STORM SSC
March 17,
04.45UT



HSSW



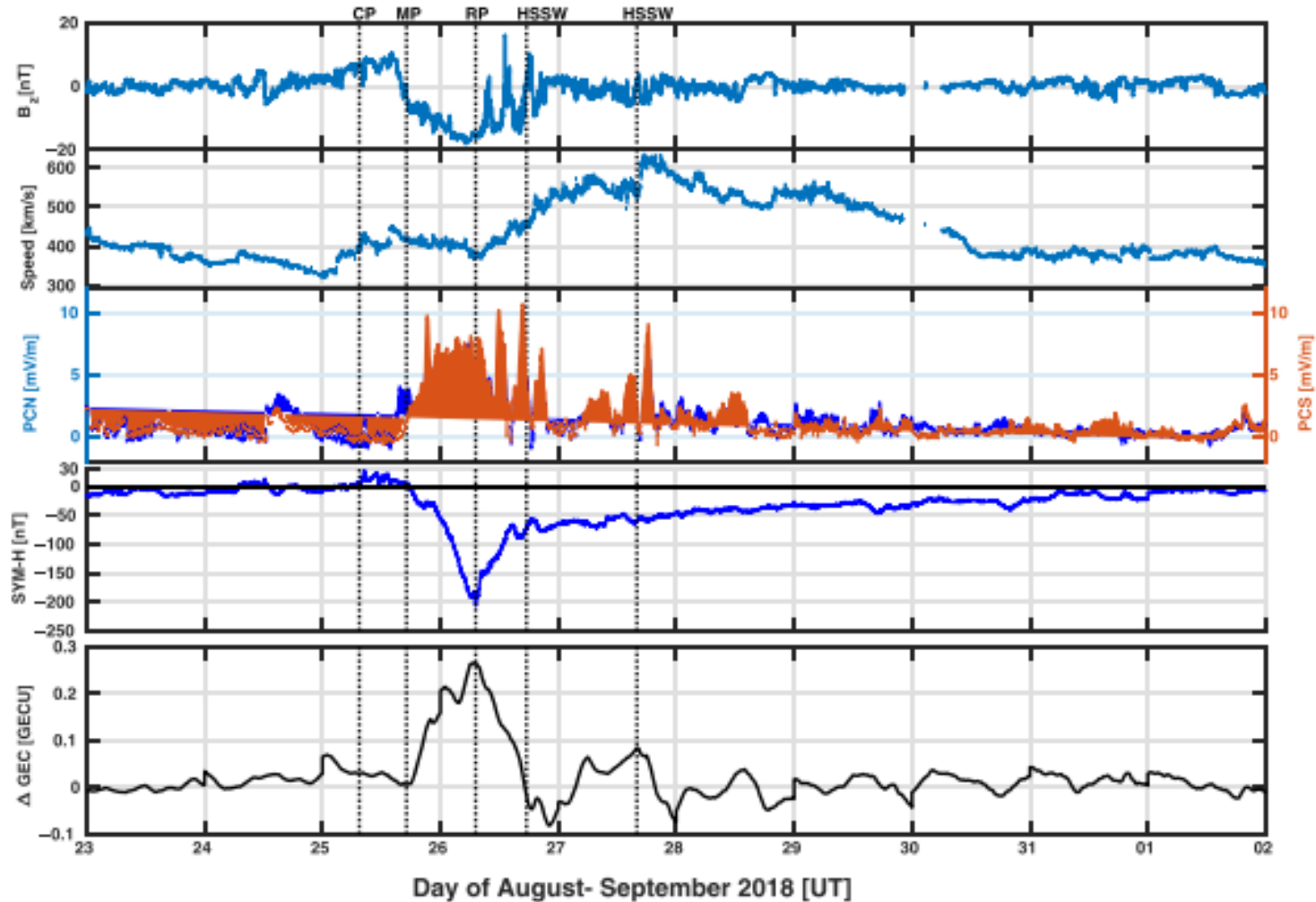
Arrival HSSW
March 21,
20.54UT

| Days | BRN2 | | JMSM | | NAST | |
|------|------------|------------|------------|------------|------------|------------|
| | RINEX-CODG | RINEX-IGSG | RINEX-CODG | RINEX-IGSG | RINEX-CODG | RINEX-IGSG |
| 14 | 10 | 10 | 9 | 10 | 6 | 9 |
| 15 | 11 | 13 | 11 | 7 | 13 | 13 |
| 16 | 13 | 16 | 16 | 16 | 10 | 11 |
| 17 | 4 | 5 | 2 | 2 | -4 | -4 |
| 18 | -3 | -2 | 0 | 1 | -3 | -3 |
| 19 | -1 | -5 | -1 | 0 | -3 | -4 |
| 20 | 12 | 19 | 10 | 21 | 7 | 19 |
| 21 | 3 | 8 | 3 | 8 | 12 | 13 |
| 22 | 19 | 19 | 16 | 17 | 10 | 14 |
| 23 | 7 | 8 | -3 | -2 | 1 | 1 |
| 24 | -10 | -7 | 1 | 1 | -3 | -4 |

Comparison with GIM maps provided by CODE:
CODG <http://aiuws.unibe.ch/ionosphere/>
and IGSG.

Pandit, D. et al., VTEC Observations of Intense Geomagnetic Storms above Nepal: Comparison with Satellite Data CODE and IGSG Models in Indian J Phys <https://doi.org/10.1007/s12648-022-02441-w>

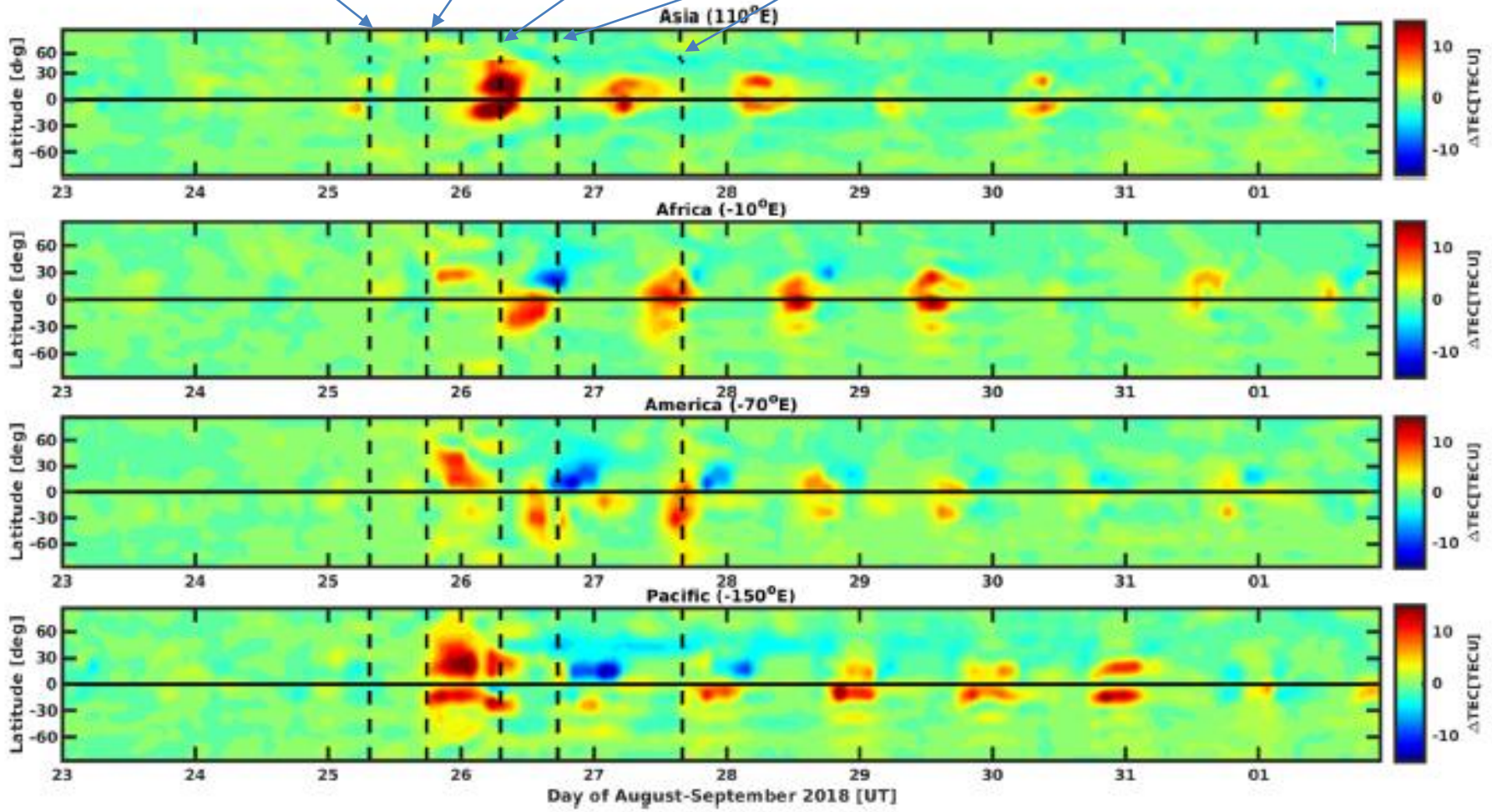
Global parameters, from 23 August to 1 September: (from top to bottom) the Bz component of IMF in nanotesla, the solar wind speed in km/s, the SYM-H index in nanotesla, polar cap indices in mV/m, and GEC in GECU **[PAKISTAN]**



Younas, W. C., C. Amory-Mazaudier, M. Khan, R. Fleury, Ionospheric and Magnetic signatures of a Space Weather event on 25-29 August 2018 : CME and HSSWs, , *Journal of Geophysical Research: Space Physics*, 125, e2020JA027981.

<https://doi.org/10.1029/2020JA027981>

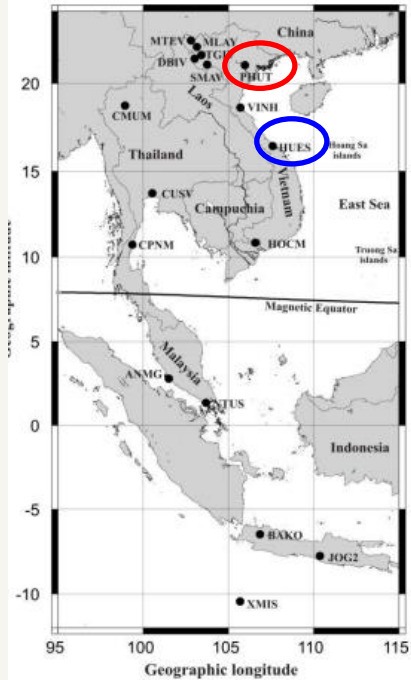
Compression Phase- SSC main phase Recovery phase HSSW MAGNETIC STORM



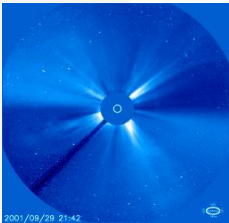
VTEC is influenced by PPEF, DDEF and Thermal expansion of atmosphere

S₄ and ROTI index in Vietnam

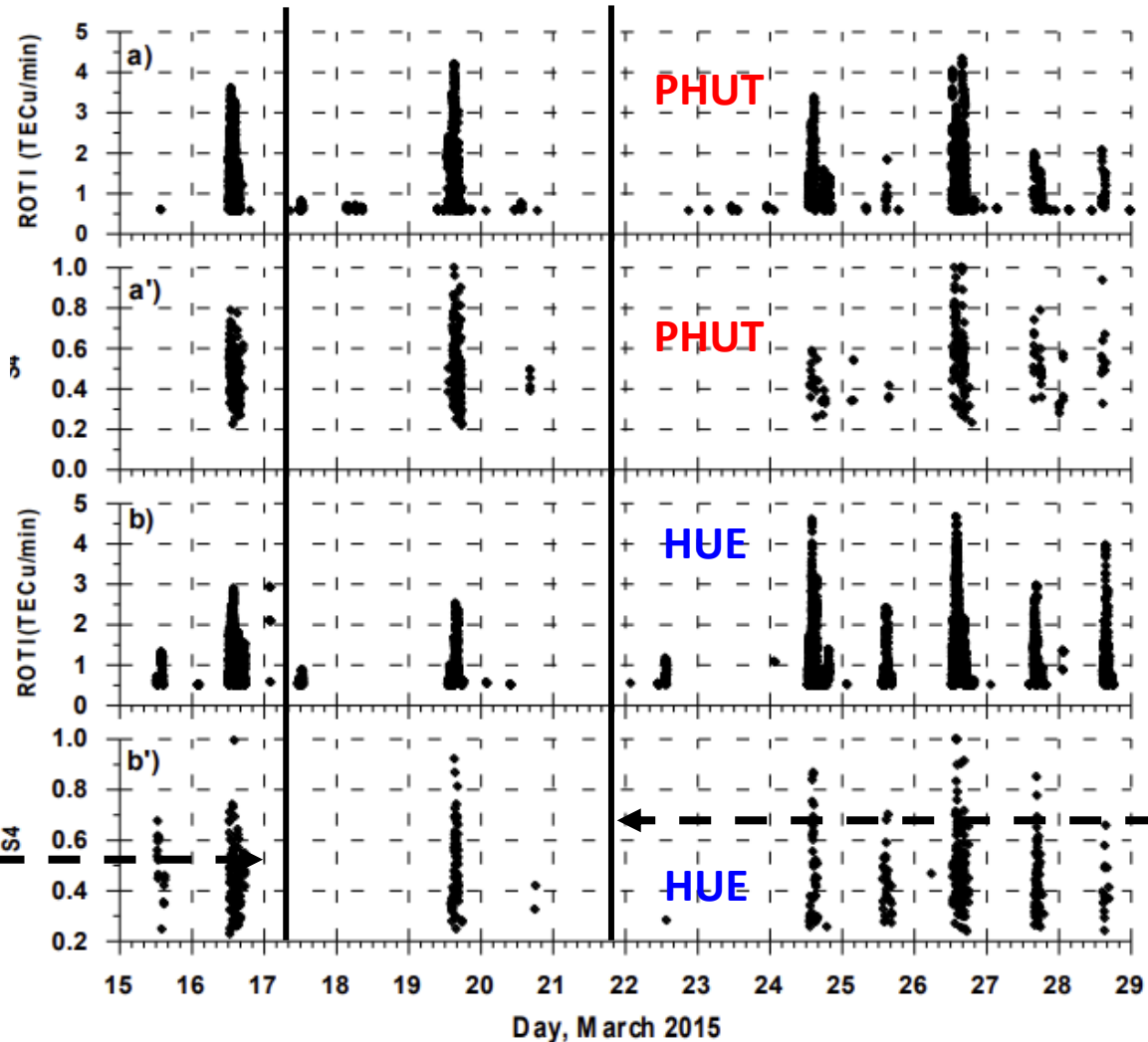
Vietnam Journal of Earth Sciences Vol 38 (3) 267-284



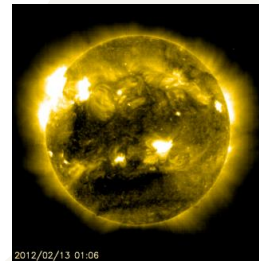
CME



STORM SSC
March 17, 04.45UT



HSSW



Arrival HSSW
March 21,
20.54UT

CONCLUSION

The use of the GNSS technique has allowed the development of studies on the ionosphere in countries where the ionosphere was not studied for lack of scientific tools.

These studies carried out within the framework of the IHY 2007-2009) and ISWI (2010-2012) projects integrating a systemic approach of the Sun-Earth system have enabled the emergence of pioneers in the discipline of Space Weather in many countries.

These students had new data that led them to publish in the best journals, to have a position in their country and to be recognized internationally.

The strength of the GNSS technique is that it works continuously and it can capture all the variations of ionospheric ionization due to different physical phenomena and therefore study their impacts on ionosphere (solar flare, eclipse, earthquake, SW, QBO, hurricane, etc...)