# **Italian Contribution to Space Weather**

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UN COPUOS 53rd Session STSC 18 February 2016.

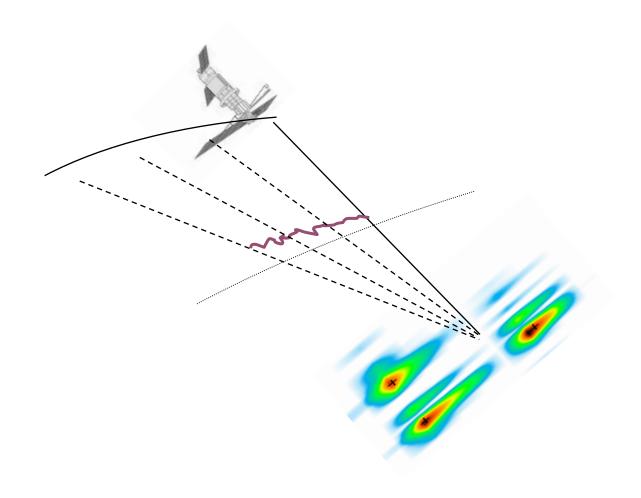
**Themis** 

Solar Orbiter

## Outline



# Italian Space Weather strategic initiatives







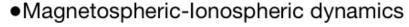
## Observational, theorethical studies and modeling



 Solar physics from photosphere to corona and sola irradiance

Interplanetary medium physics: structures,
 turbulence and propagation of CMEs and SEPs

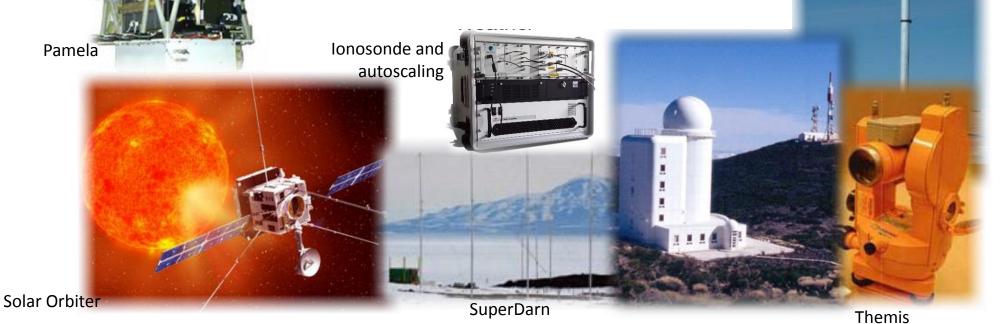
Solar wind-magnetosphere coupling and interaction



Ground based magnetic field variations

Forecasting and nowcasting modelling

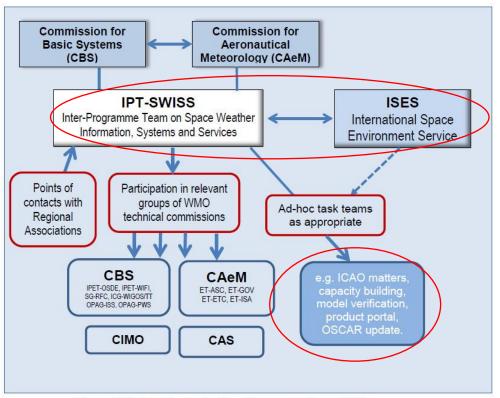
Planetary Space Weather



## **World Meteorological Organization Congress**

Resolution 38 (Cg-17) — "Four-year Plan for WMO Coordination of Space Weather Activities".

Since 2012 Italy joined the WMO Space Weather initiative <a href="Inter-programme Coordination Team on SW">Inter-programme Coordination Team on SW</a>, ITAF – INAF - INGV









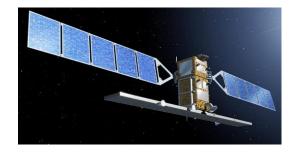


# Space Weather italian initiative for operations

SW nowcasting and safety support

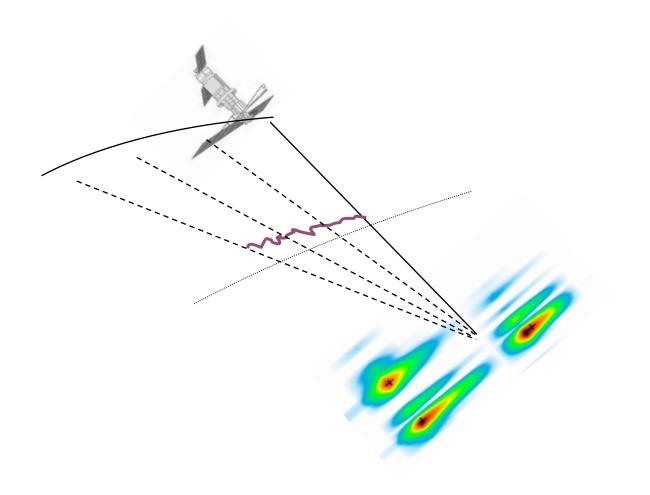






Space Weather knowledge is not only for safety but also for capacity augmentation, as weather.

# Solar physics to Space Weather





# **Solar Physics activities in Trieste**

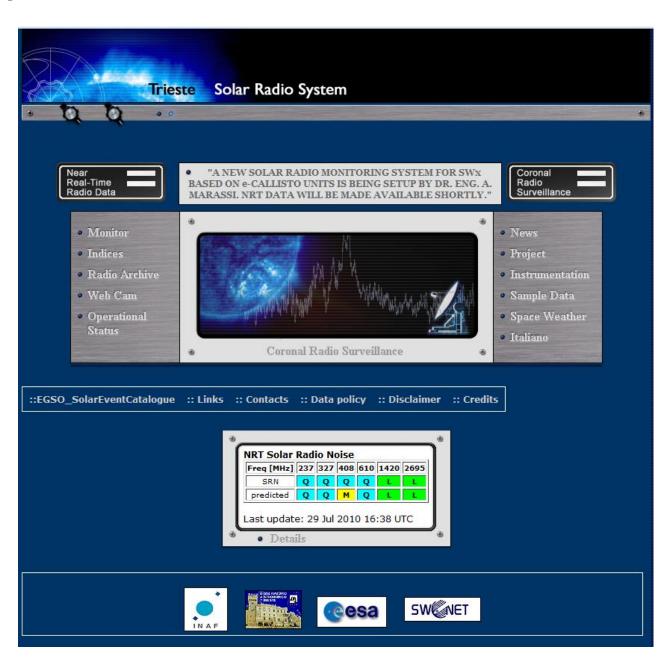
ESA Space Weather Working Team,
Steering Board Member
CESA

European Space Weather Week
Programme Committee, Chair

NATO Science for Peace (SfP) Project 984894 on "Ionospheric Monitoring", Co-Director



M. Bilal, V. Alberti, A. Marassi, E. Cianca, M. Messerotti, Performance assessment of GPS receivers during the September 24, 2011 solar radio burst event, J. of Space Weather and Space Climate, **5**, A32, 16 pp., 2015.



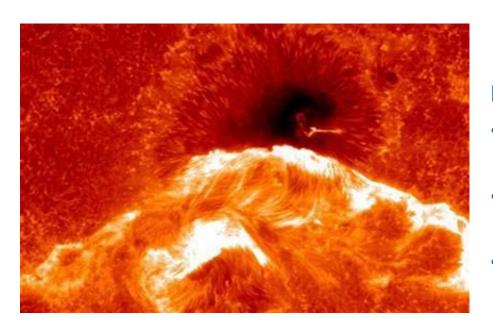
# Solar Physics Group in Catania

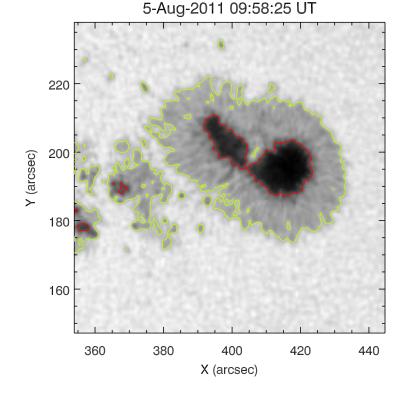
#### Personnel

V. Capparelli (UniCT), A. Compagnino (UniCT), M. Falco (UniCT), S.L. Guglielmino (UniCT), M. Murabito (UniCT), P. Romano (INAF), F. Zuccarello (UniCT).

#### **Main Research Fields**

Participation in the European Solar Telescope Design Phase; Emergence of magnetic flux tubes in the solar atmosphere; Formation and evolution of solar active regions; Flares and Coronal Mass Ejections: drivers and effects on the space environment; Space Weather.





#### **Methods**

- Coordinated observing Campaigns between ground-based and space-based satellites
- Analysis of spectroscopic and spectro-polarimetric data acquired from space and ground.
- Design and development of new instrumentation for future ground based observations.

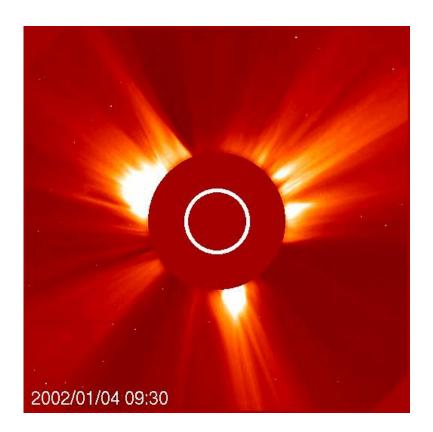
# Solar Physics Group in Catania

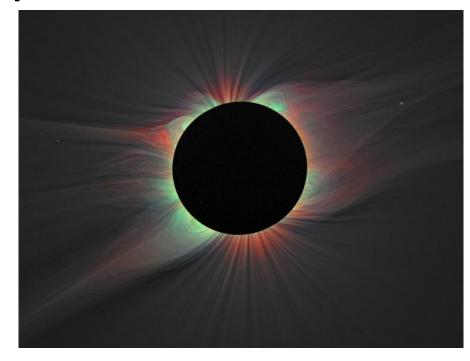
Project Name	Short description	Role	Timeline
SOLARNET	The project brings together and integrates the major European research infrastructures in the field of high-resolution solar physics, in order to promote their coordinated use and development (FP7)	Responsible for WP30: Networking Activities (Leader: F. Zuccarello)	2013 April 1 – 2017 March 31
F-CHROMA F-CHROMA	To acquire, analyse and interpret ground- and space-based observational data of solar flares, test these against model predictions, and create an archive of solar flare observations and models (FP7)	Responsible for WP5: Joint analysis of space-based and ground-based observations (Lead: F. Zuccarello)	2014 January 1 – 2016 December 31
Metis	WL and UV Coronagraph for ESA-Solar Orbiter spacecraft → first close-up (0.3 AU) observations of coronal plasmas	Participant to Science Team	Launch: October 2018, nominal mission 7.5 years

# Solar Physics Group in Turin

#### **Main Research Fields**

 Physics of the solar corona, understanding the origin and evolution of the main drivers of Geomagnetic Storms on Earth: Solar Wind and Coronal Mass Ejections (CMEs).





#### Methods

- Coordination of observational campaigns from space and ground (total solar eclipses)
- Development of diagnostic techniques for the analysis of coronagraphic and spectroscopic data acquired from space and ground.
- Development of new instrumentation for future space missions and ground based observations.

# Solar Physics Group in Turin

## **On-going experimental projects**

Project Name	Short description	Role	Timeline
Metis	WL and UV Coronagraph for ESA- Solar Orbiter spacecraft → first close-up (0.3 AU) observations of coronal plasmas	Leader of the international science consortium (PI: E. Antonucci)	Launch: October 2018, nominal mission 7.5 years
ASPIICS ASPIICS Formation Flying Coronagraph	WL coronagraph for ESA-PROBA3 satellite → first eclipse-like, long-term observations of the inner corona	Italian leader for Formation Flying metrology (Lead Co-I: S. Fineschi)	Launch: 2019, nominal mission 2 years
SCORE	Helium Sounding rocket coronagraph → first determination of coronal Helium abundance	Leader of the italian instrument consortium (PI: S. Fineschi)	First launch: September 2009, Second launch: June 2016
ESCAPE    INLIANDED   PORTON   PORTON	Coronagraph in Antarctica (Concordia base) → first long- term coronal magnetic fields monitoring	Leader of the italian instrument consortium (Co-PI: S. Fineschi)	Deployment: 2017, nominal duration 3 years

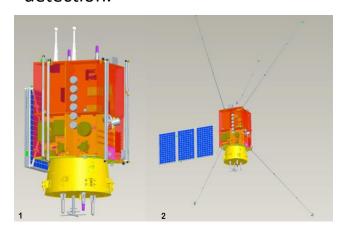


## **University of Rome Tor Vergata**

https://www.fisica.uniroma2.it/solare

F. Berrilli, M. Casolino, D. Del Moro, L. Giovannelli, R. Forte, M. Lovric, M. Martucci, M. Mergè, L. Narici, V. Penza, G. Pucacco, F. Pucci, A. Rizzo, S. Scardigli, R. Sparvoli

Main Projects: FP7-EST, FP7-SOLARNET, H2020-GREST, EU-REACT-SPARC, EU-Ionosphere Prediction Service, PAMELA, ALTEA, CSES Activity: Solar Dynamics and Activity, Sun-Earth interaction, Space Weather, Improve tools for solar synoptic observations and particle detection.



Italian Collaboration to
CSES
China SeismoElectromagnetic
Satellite

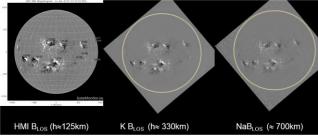
Royal Institute of Technology, Stockholm, Sweden

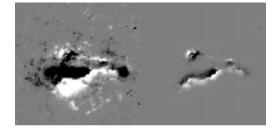


PAMELA satellite in orbit since 2006



Magneto Optical filters at Two Heights (**MOTH**) instrument University of Hawaii, USA Jet Propulsion Laboratory, Japan Eddy Company, USA



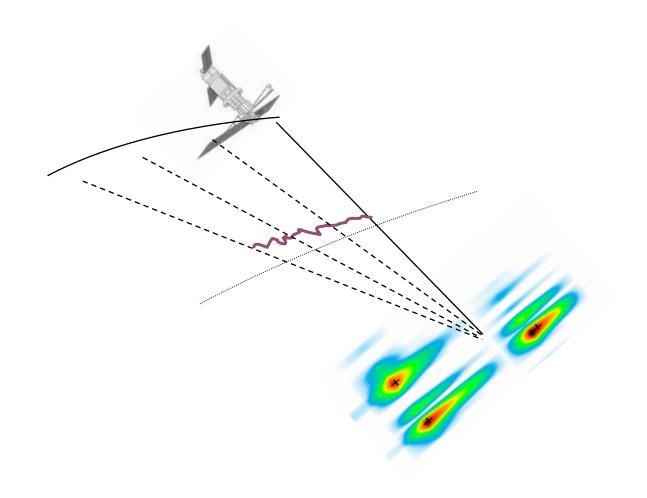


Automated Solar Flare
Forecasting with multiline
MOTH synoptic magnetograms

#### **PAMELA** collaborations

Cosmic Rays Laboratory, Moscow Engineering and Physics Institute, Moscow, Russia Laboratory of Solar and Cosmic Ray Physics, P.N. Lebedev Physical Institute Academy of Sciences, Moscow, Russia Ioffe Physical Technical Institute, St. Petersburg, Russia Physics Department of Siegen University, Germany

# Interplanetary space physics to Space Weather



## Interplanetary Space Physics Group @ INAF/IAPS

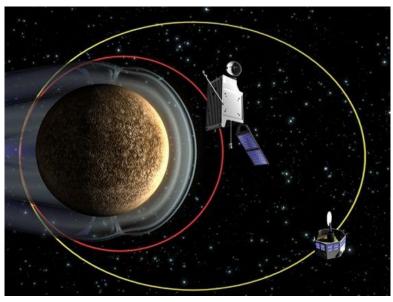




**Solar Orbiter** - A high-resolution mission to the Sun and inner heliosphere. The ISP group participates to the SWA, a plasma feature instrument suite, with the responsibility of the development of the on board DPU.

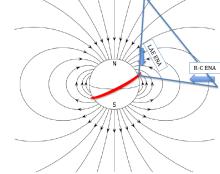
# Super Dual Auroral Radar Network international network of HF ionospheric radars dedicated to the study of the magnetosphere-ionosphere system - The ISP group is responsible for the Dome C East radar located at the research station Dome Concordia in Antarctica.





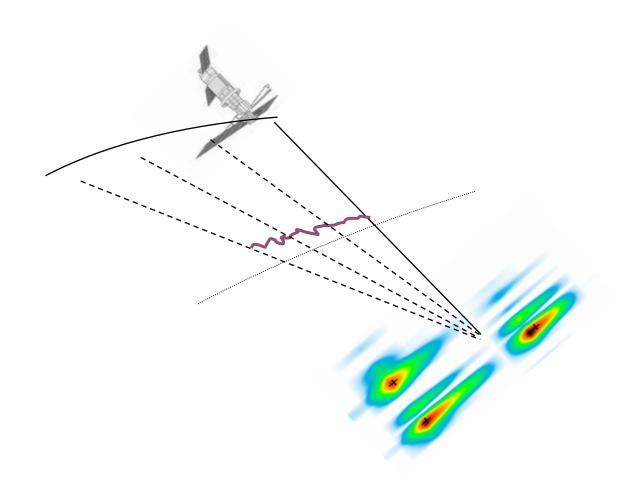
BepiColombo an ESA mission to Mercury – The ISP group is involved in the MEA (Mercury Electron Analyzer) and SIXS (Solar Intensity X-ray and particle Spectrometer) experiments onboard Mercury Magnetospheric Orbiter and Mercury Planetary Orbiter, respectively.





PROPOSAL: Development of an ENA sensor, namely **ENAMISS**, to be uploaded on the International Space Station for continuous magnetosphere observation

# Solar-Terrestrial physics to Space Weather

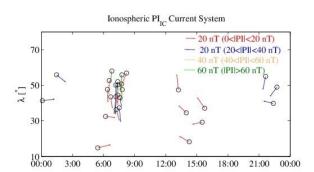


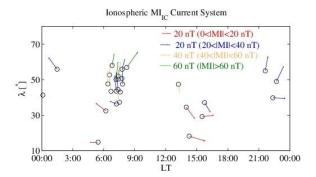




# University of L'Aquila

# Analysis of the ground-based and magnetospheric response to active Solar Wind (SW) conditions



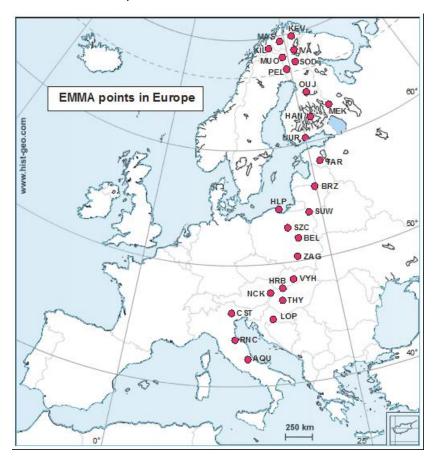


Example of the reconstruction of the ionospheric current flow pattern during November 11, 2000 SI event.

#### **Scientific collaborations**

- Geological and Geophysical Institute of Hungary, Hungary
- Electrical Engineering Department, New Mexico Tech, USA
- Institute of Geophysics-PAS, Poland
- Finnish Meteorological Institute, Finland
- Space Research Institute (IWF), Graz, Austria
- School of Mathematical and Physical Sciences, University of Newcastle, Callaghan, New South Wales, Australia.
- Physics Department, University of Calabria, Rende (CS), Italy.
- National Institute for Geophysics and Volcanology INGV, Rome, Italy.
- National Research Council, Institute for Complex Systems ISC-CNR, Florence, Italy.

EMMA (European Meridional Magnetometer Array) 25 stations, 1.6 < L < 6.1



Villante, U., S. Di Matteo, and M. Piersanti (2015), On the transmission of waves at discrete frequencies from the solar wind to the magnetosphere and ground: A case study, J. Geophys. Res. Space Physics,120, doi:10.1002/2015JA021628.

Lichtenberger, J., M. A. Clilverd, B. Heilig, M. Vellante, J. Manninen, C. J. Rodger, A. B. Collier, A. M. Jørgensen, J. Reda, R. H. Holzworth, R. Friedel, and M. Simon-Wedlund (2013), The plasmasphere during a space weather event: first results from the PLASMON project, *J. Space Weather Space Clim.*, 3, A23, doi: http://dx.doi.org/10.1051/swsc/2013045.

## Geomagnetic Observatories

# Italy



## Antarctica

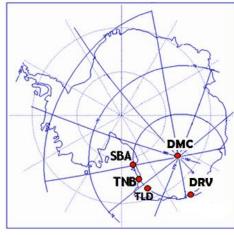


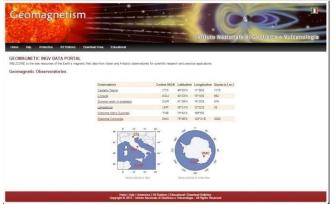


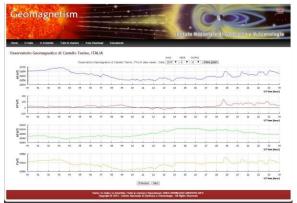






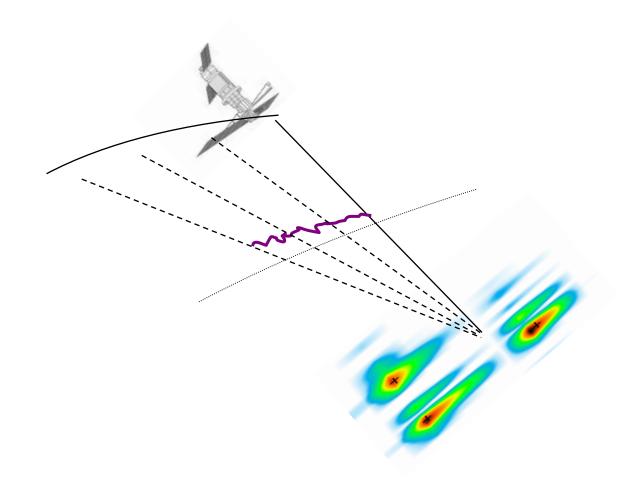






Observatory data are available at the following URL address: http://geomag.rm.ingv.it

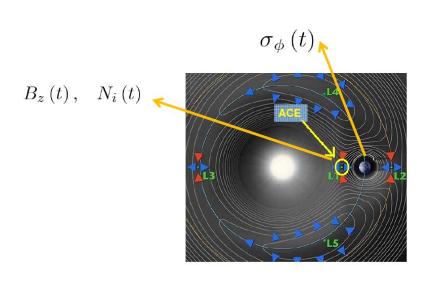
# Upper atmosphere physics to Space Weather



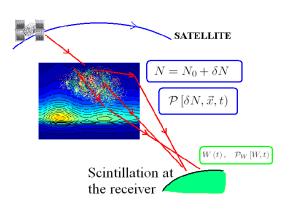
# National Research Council (Institute for Complex Systems) Recent Activities:







Predictive Space
Weather via
information theory
tools for data
analysis



Ionospheric irregularity sensing through multi-scale analysis of radio scintillation on GNSS signals

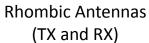
- Space Research Centre of the Polish Academy of Science, Warsaw, Poland (ionospheric irregularities and radio scintillation)
- Centre for Theoretical Physics of the University of Marseille, France (dissipative Magneto-Hydro-Dynamics)
- University of Bath, UK (information theory analysis tools applied to Space Weather)

Wernik, A. W., L. Alfonsi, and M. Materassi (2007), Scintillation modeling using in situ data, Radio Sci., 42, RS1002, doi:10.1029/2006RS003512.

## **Ionospheric Observatories**

### Gibilmanna (Italy)







AIS - INGV

## Rome (Italy)





AIS - INGV

Antennas (TX and RX)

#### S. Miguel de Tucumán (Argentina)

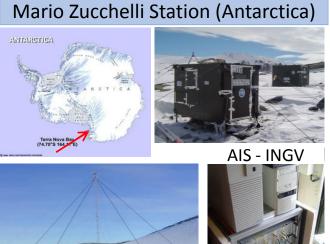


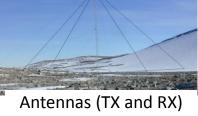
Delta Antennas (TX and RX)



AIS - INGV







## **Space Weather forecast**

vulcani ambiente

Achievement of forecasting and nowcasting three dimensional (3-D) electron density mapping of the ionosphere.

## **INGV GNSS receivers network**

First receiver installed at Ny-Alesund (Svalbard) on 2003

- Polar ionosphere
  - Svalbard islands (3)
  - Antarctica (4)
- Mid latitude ionosphere
  - Chania (Crete)
  - Huelva (Spain) stopped
  - Huelva station moved to Lampedusa
- Equatorial Ionosphere
  - Tucuman (Argentina)

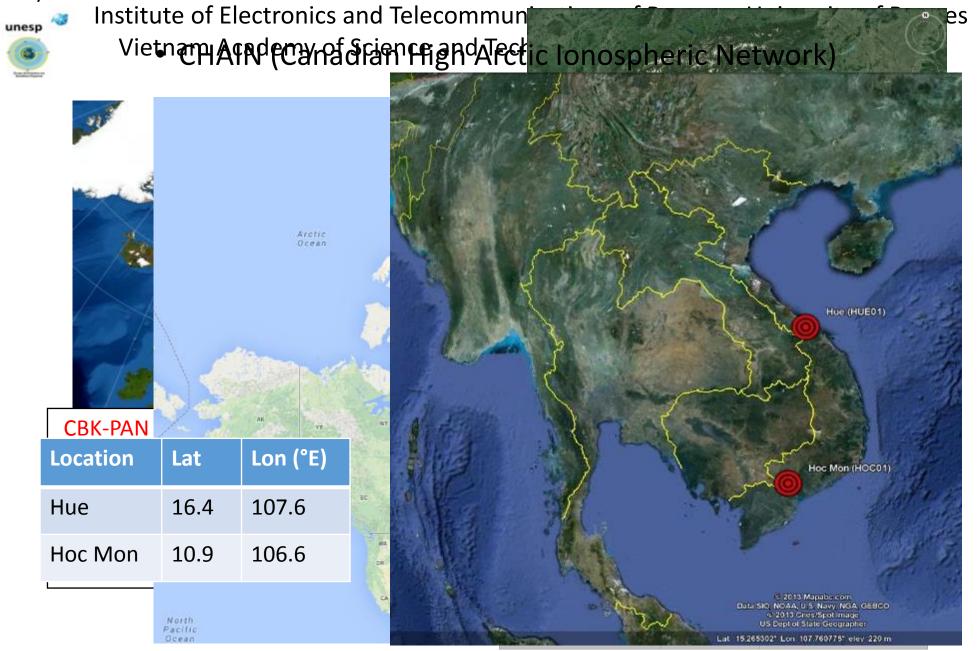


Data are accessible at the *electronic Space Weather upper atmosphere* website <a href="mailto:eSWua"><u>eSWua</u></a>
www.eSWua.ingv.it



## INGV network partners

## CIGALA/CALIBRA Network



## Most relevant past and on-going projects

**CALIBRA**: Countering GNSS high Accuracy applications Limitation due to ionospheric disturbance in BRAzil, FP7–GALILEO–2011–GSA–1

**TRANSMIT**: Training Research and Applications Network to Support the Mitigation of Ionospheric Threats, FP7-ITN Marie Curie

**ESPAS**: Near-Earth space data infrastructure for e-science, FP7-Research Infrastructure

MISW: Mitigation of space weather threats to GNSS services, FP7-Space

**ERICA**: EquatoRial Ionospheric Characterization in Asia, ESA-ALCANTARA

MIMOSA2: Monitoring Ionosphere Over South America, ESA-ALCANTARA

**DemoGRAPE**: Demonstrator of GNSS Research and Application for Polar Environment, PNRA

**GRAPE**: GNSS Research and Application for Polar Environment, SCAR

# GINESTRA – MIMOSA - MEDSTEC COMPETENCE SURVEYS WITHIN THE ESA ALCANTARA INITIATIVES

## MImOSA2

Monitoring Ionosphere Over South America to support high precision applications

## **ERICA**

**E**quato**R**ial **I**onosphere **C**haracterization in **A**sia

















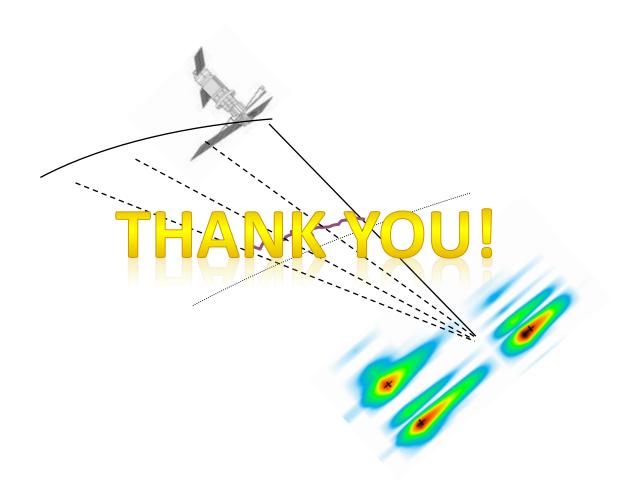












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