



# How does GNSS positioning depend on interoperability?

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# Interoperability problems

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- What is interoperability?
  - PNT services of dual or multiple GNSS systems can be used together to provide **better capabilities** at **the user level** than that would be achieved by relying solely on a **single** system or service



# Interoperability factors

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- Signal-in-space:
  - All generate codes and carrier phase measurements, carrying data messages
  - Each system transmits signals at different frequencies
  - Different frequencies introduce different inter-frequency biases
- Satellite clocks of different systems
  - Are maintained by different groups of ground clocks
- Reference Systems
  - With same or different definitions and parameters
- Reference Frames
  - Different realisation of the same or different reference systems



# What “better capabilities”?


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- Better accuracy
  - Position (absolute or relative)
  - Clock (absolute or relative)
- Higher availability
- Improved integrity (multiple failures etc)
- Better AR reliability (high success rates)
- Sharper timeliness
  - Shorter time to first fix
  - Shorter time to convergence
- Longer baseline/wider coverage
- More ionospheric/tropospheric profiles?
- etc

# Observational Models

GPS L1, L2, L5	Galileo, L1,E6, E5A, E5B
$P_{L1,i}^j = \tilde{\rho} + I_{L1} + T_g^j + \epsilon_{PL1}$ $P_{L2,i}^j = \tilde{\rho} + \alpha I_{L1} + \alpha T_g^j + R_{2i} + \epsilon_{PL2}$ $P_{L5,i}^j = \tilde{\rho} + \beta I_{L1} + \beta T_g^j + R_{5i} + \epsilon_{PL5}$	$P_{L1,i}^j = \tilde{\rho} + I_{L1} + T_g^j + \epsilon_{PL1}$ $P_{E6,i}^j = \tilde{\rho} + \alpha I_{L1} + \alpha T_g^j + R_{E6i} + \epsilon_{PE6}$ $P_{E5a,i}^j = \tilde{\rho} + \beta I_{L1} + \beta T_g^j + R_{5ai} + \epsilon_{PE5a}$ $P_{E5b,i}^j = \tilde{\rho} + \gamma I_{L1} + \gamma T_g^j + R_{5bi} + \epsilon_{PE5b}$
$\Phi_{L1,i}^j = \tilde{\rho} - I_{L1} + T_g^j + \lambda_1 N_1 + \epsilon_{L1}$ $\Phi_{L2,i}^j = \tilde{\rho} - \alpha I_{L1} + \alpha T_g^j + \lambda_2 N_2 + R_{2i} + \epsilon_{L2}$ $\Phi_{L5,i}^j = \tilde{\rho} - \beta I_{L1} + \beta T_g^j + \lambda_5 N_5 + R_{5i} + \epsilon_{L5}$	$\Phi_{L1,i}^j = \tilde{\rho} - I_{L1} + T_g^j + \lambda_1 N_1 + \epsilon_{L1}$ $\Phi_{E6,i}^j = \tilde{\rho} - \alpha I_{L1} + \alpha T_g^j + \lambda_{E6} N_{E6} + R_{E6i} + \epsilon_{e6}$ $\Phi_{E5b,i}^j = \tilde{\rho} - \beta I_{L1} + \beta T_g^j + \lambda_{E5a} N_{e5a} + R_{e5ai} + \epsilon_{e5a}$ $\Phi_{E5b,i}^j = \tilde{\rho} - \gamma I_{L1} + \gamma T_g^j + \lambda_{E5b} N_{E5b} + R_{e5bi} + \epsilon_{e5b}$
$\tilde{\rho} = \rho + c(dt - dT) + d_{trop} + d_{orb}$	$\tilde{\rho} = \rho + c(dt - dT) + d_{trop} + d_{orb}$
$\alpha = (f_1/f_2)^2 = (77/60)^2;$ $\beta = (f_1/f_5)^2 = (154/115)^2$	$\alpha = (f_1/f_{E6})^2 = (154/125)^2$ $\beta = (f_1/f_{e5a})^2 = (154/118)^2$ $\gamma = (f_1/f_{e5b})^2 = (154/115)^2$

# Single Point Positioning Model


$$\begin{bmatrix} P_a \\ P_b \end{bmatrix} = \begin{bmatrix} A_a \\ A_b R_{ba} \end{bmatrix} \delta X + \begin{bmatrix} cJ & 0 \\ 0 & cJ \end{bmatrix} \begin{bmatrix} \delta T_a \\ \delta T_b \end{bmatrix} + \begin{bmatrix} \epsilon_a \\ \epsilon_b \end{bmatrix}$$

$$\delta X \equiv \delta X_a \quad \delta X_b = R_{ba} \delta X_a$$

$$E\left(\begin{bmatrix} \epsilon_a \\ \epsilon_b \end{bmatrix}\right) = 0; \quad \text{Cov}\left(\begin{bmatrix} \epsilon_a \\ \epsilon_b \end{bmatrix}\right) = \begin{bmatrix} \sigma_a^2 Q_a & 0 \\ 0 & \sigma_b^2 Q_b \end{bmatrix} = \sigma_a^2 Q$$

- (1) Determine the same position vector, but different clock biases
- (2) Important to balance between two code variance components for better positioning accuracy
- (3)  $Q_a$  and  $Q_b$  ideally include the effects of SV orbital variance matrix

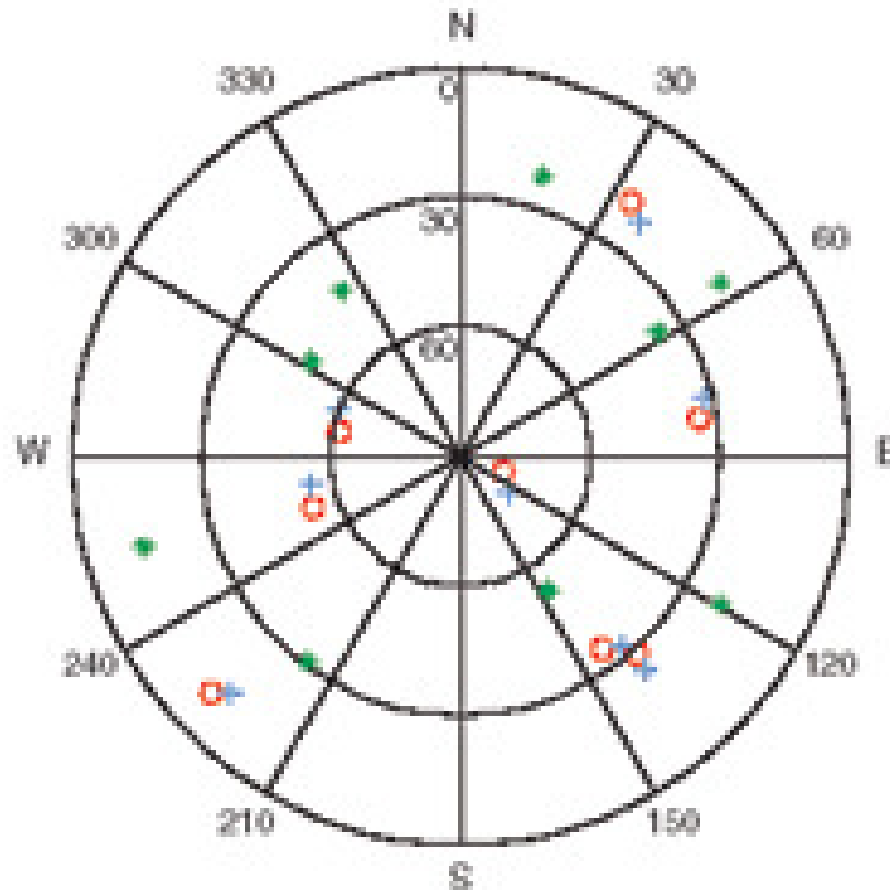
# Double Differenced Models

$$\begin{bmatrix} \Delta P_a \\ \Delta \phi_a \\ \Delta P_b \\ \Delta \phi_b \end{bmatrix} = \begin{bmatrix} A_a & 0 & 0 \\ A_a & -\lambda_a I_a & 0 \\ A_b R_{ba} & 0 & 0 \\ A_b R_{ba} & 0 & -\lambda_b I_b \end{bmatrix} \begin{bmatrix} \delta X \\ \Delta N_a \\ \Delta N_b \end{bmatrix} + \begin{bmatrix} \epsilon_{\Delta P a} \\ \epsilon_{\Delta \phi a} \\ \epsilon_{\Delta P b} \\ \epsilon_{\Delta \phi b} \end{bmatrix}$$

$$E\left(\begin{bmatrix} \epsilon_{\Delta P a} \\ \epsilon_{\Delta \phi a} \\ \epsilon_{\Delta P b} \\ \epsilon_{\Delta \phi b} \end{bmatrix}\right) = 0; \quad \text{Cov}\left(\begin{bmatrix} \epsilon_{\Delta P a} \\ \epsilon_{\Delta \phi a} \\ \epsilon_{\Delta P b} \\ \epsilon_{\Delta \phi b} \end{bmatrix}\right) = \begin{bmatrix} \begin{bmatrix} \sigma_a^2 Q_a & 0 \\ 0 & \sigma_{\phi a}^2 Q_{\phi a} \end{bmatrix} \otimes D_a D_a^T & 0 \\ 0 & \begin{bmatrix} \sigma_b^2 Q_b & 0 \\ 0 & \sigma_{\phi b}^2 Q_{\phi b} \end{bmatrix} \otimes D_b D_b^T \end{bmatrix}$$

1. No cross-constellation differencing avoids requirements for a same frequency
2. Balancing between the variance components of different systems is important for better RTK performance

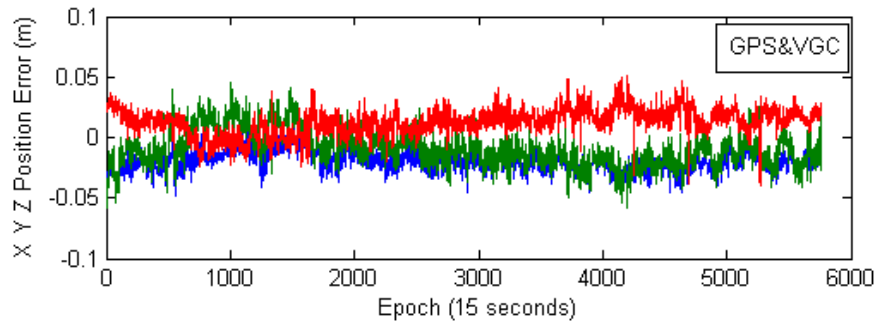
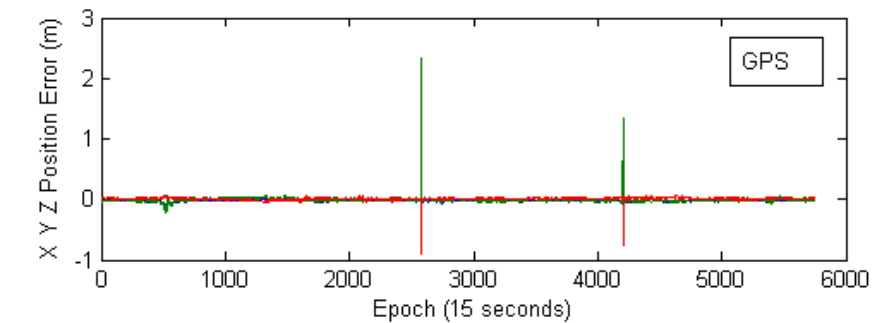
# Virtual Galileo Constellation



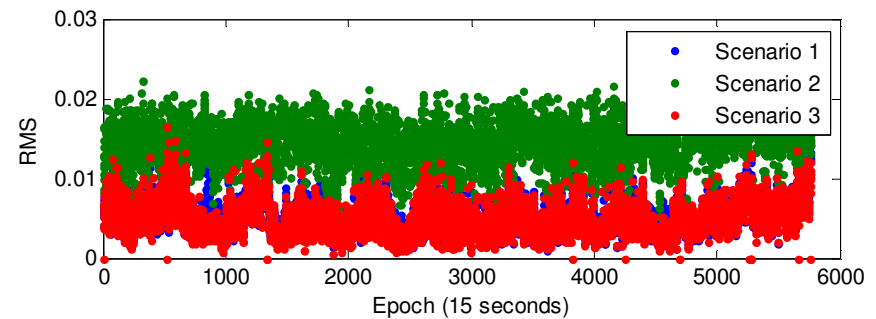
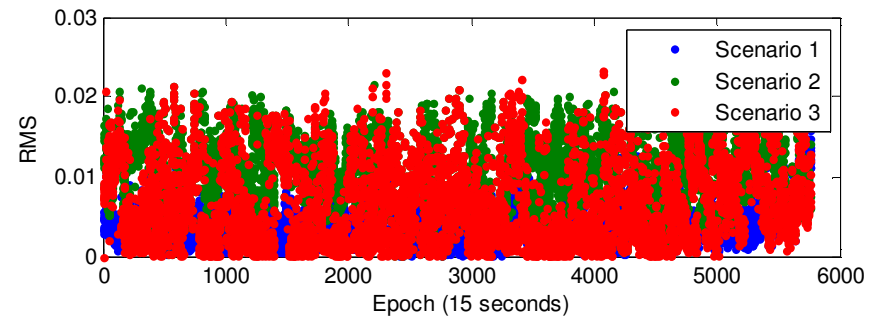
$$V(\epsilon_{Pa}) = AS_x A_a^T + \sigma_{Trop}^2 + \sigma_{iono}^2 + \sigma_{multipath}^2 + \sigma_{ea}^2$$



# AR success rate and RMS values



Comparison of AR success rates



Comparison of RMS values

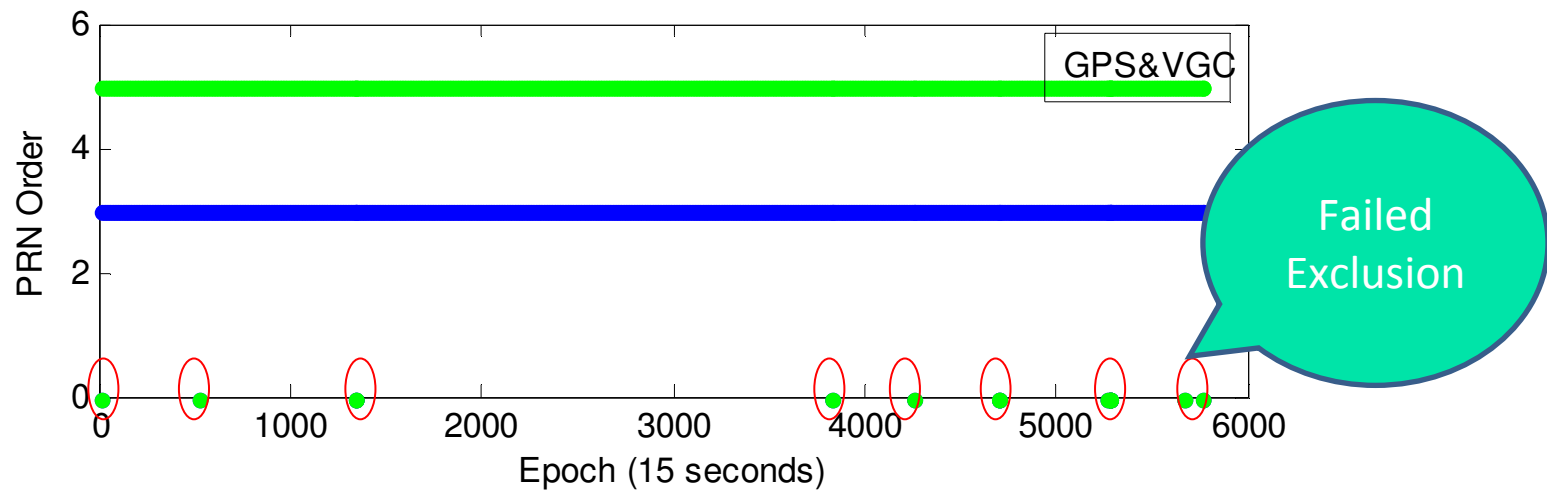
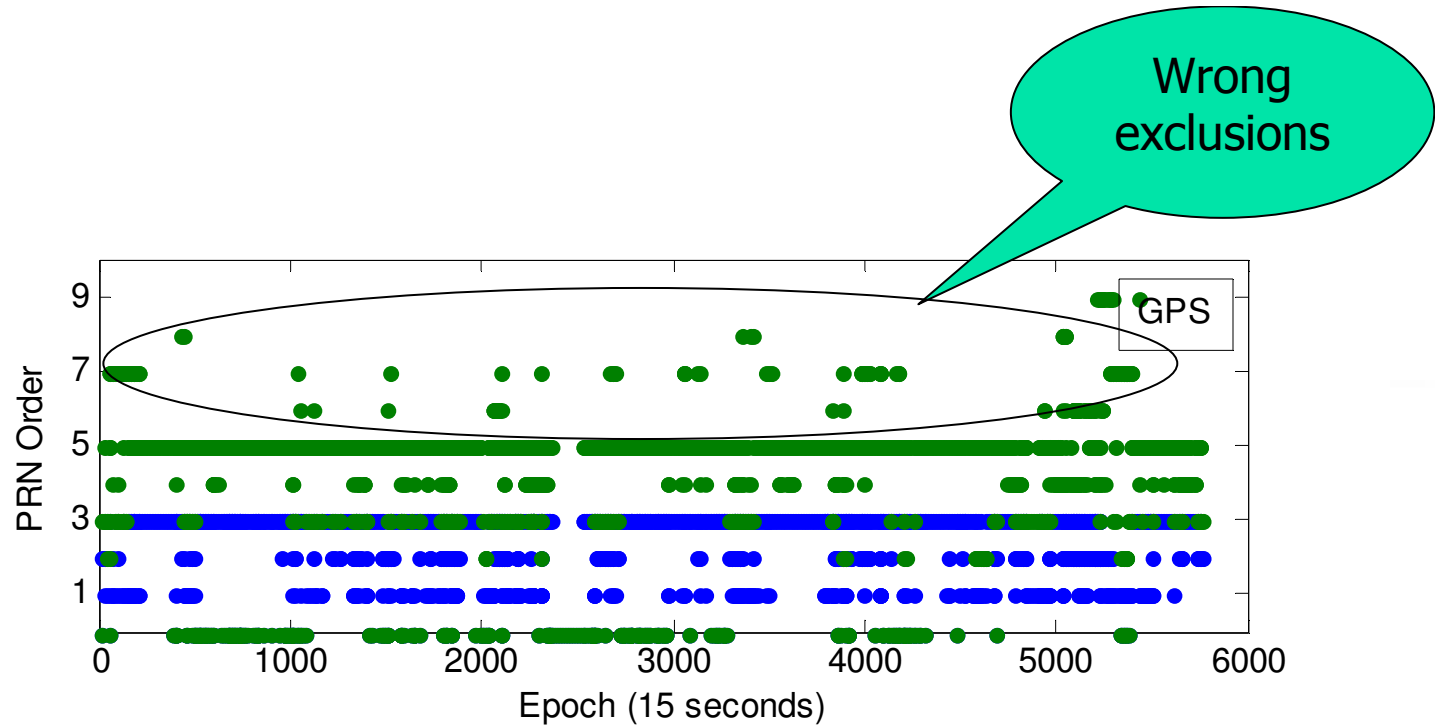
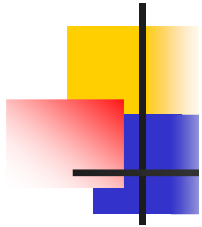
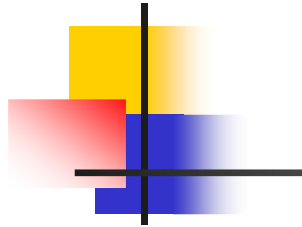



Illustration two detected fault satellites



	Yes/No	Confirmed
Better position accuracy?	Yes	Yes
Better clock accuracy?	Not sure	Not Yet
Higher availability?	Yes	Yes
Improved integrity	Yes	Yes
Better AR reliability?	Yes	Yes
Sharper timeliness	Yes	Not Yet
Longer baseline/wider coverage	Yes	Not Yet
More atmospheric soundings	Yes	Yes
More ionospheric samples	Yes	Yes

# Concluding remarks



Observational models and estimation algorithms can be flexibly designed to achieve most of better capabilities without additional requirements for interoperability

- It would be beneficial to real time users if each system broadcasts SV variance matrices