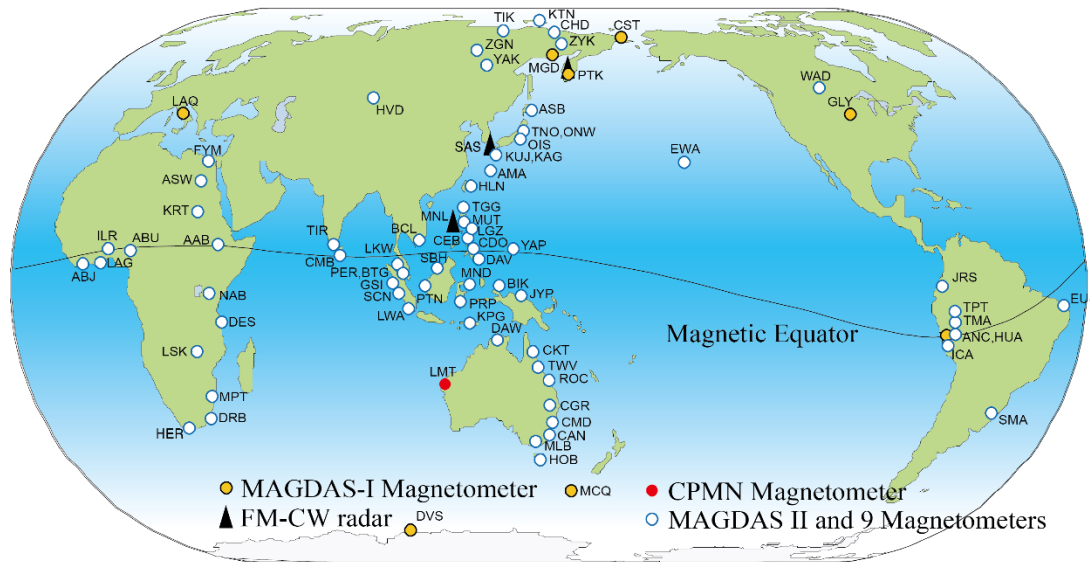




Latest scientific results of MAGDAS project - Seasonal dependence of semidiurnal equatorial magnetic variations -

MAGDAS/CPMN

(MAGnetic Data Acquisition System/Circum-pan Pacific Magnetometer Network)

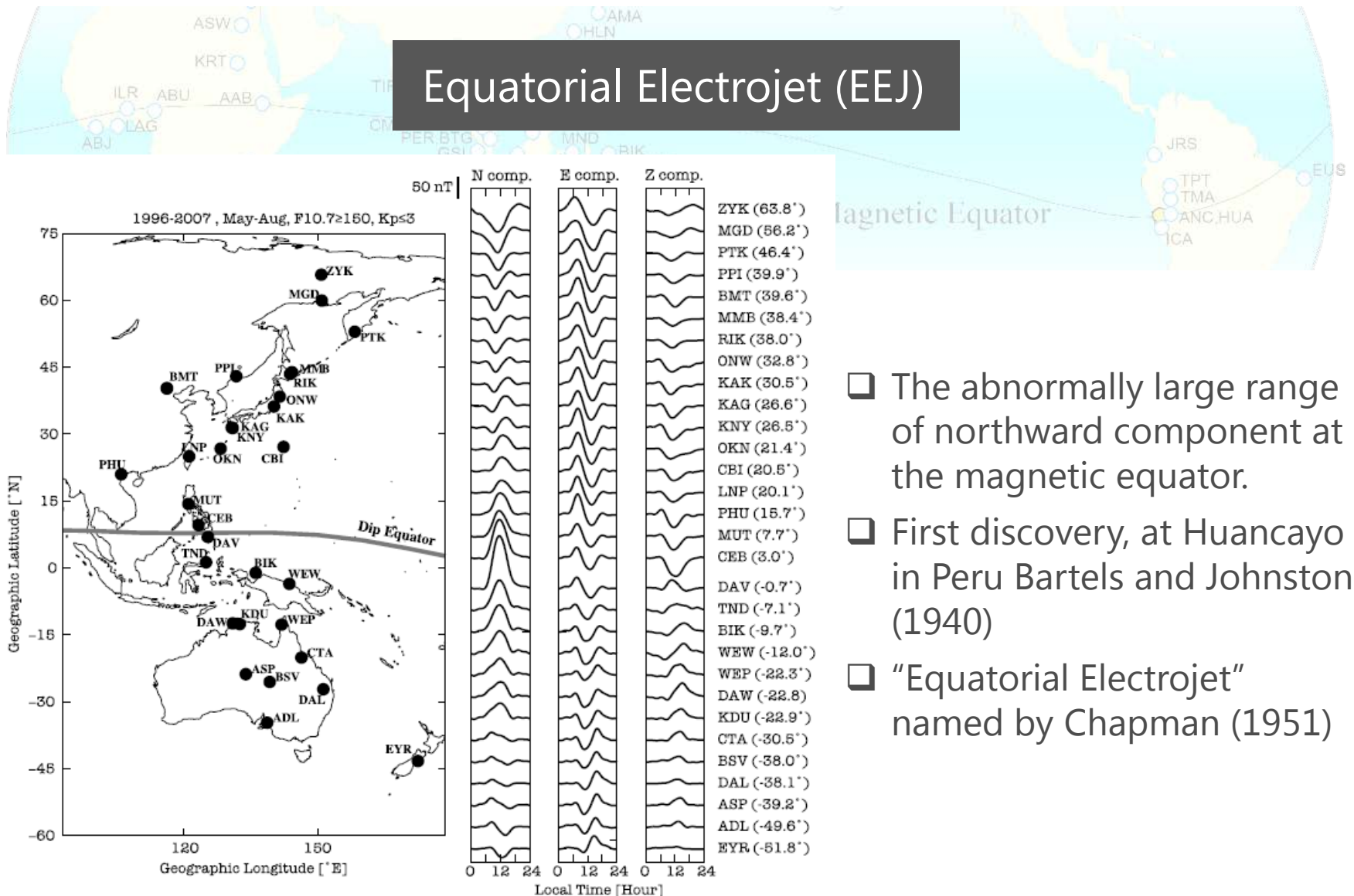


Akiko Fujimoto¹

Akimasa Yoshikawa¹, Teiji Uozumi¹, Shuji Abe¹ and Hiroki Matsushita²

1. International Center for Space Weather Science and Education, Kyushu University
2. Earth and Planetary sciences, Kyushu University

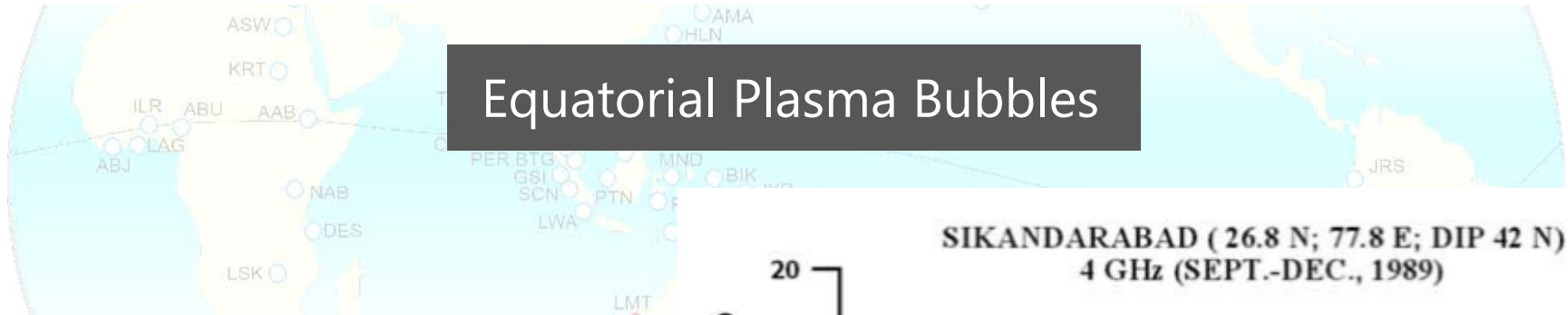
The importance of equatorial magnetic observation



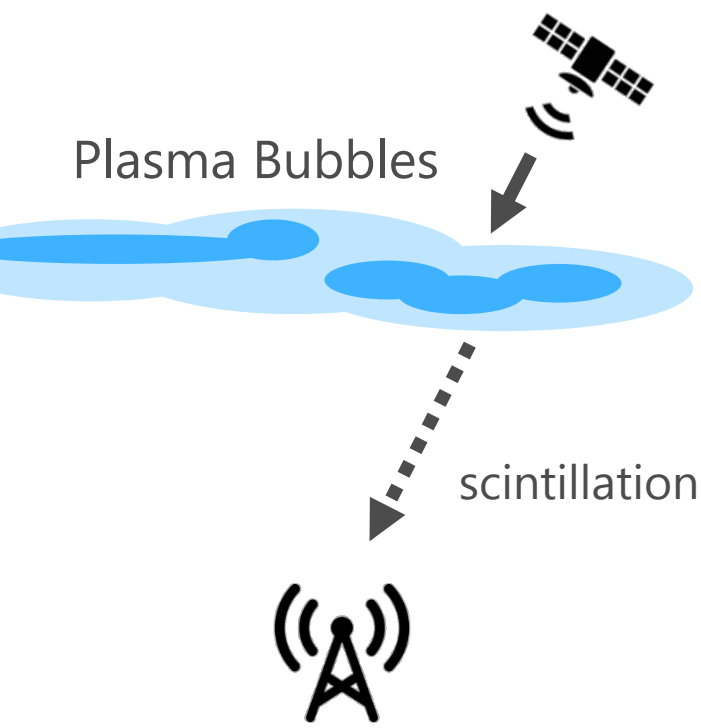
- ❑ The abnormally large range of northward component at the magnetic equator.
- ❑ First discovery, at Huancayo in Peru Bartels and Johnston (1940)
- ❑ “Equatorial Electrojet” named by Chapman (1951)

Fig. 2 Average geomagnetic daily variations in the magnetic-northward (*N*), magnetic-eastward (*E*), and vertically downward (*Z*) components during May–August of 1996–2007. From Yamazaki (2011)

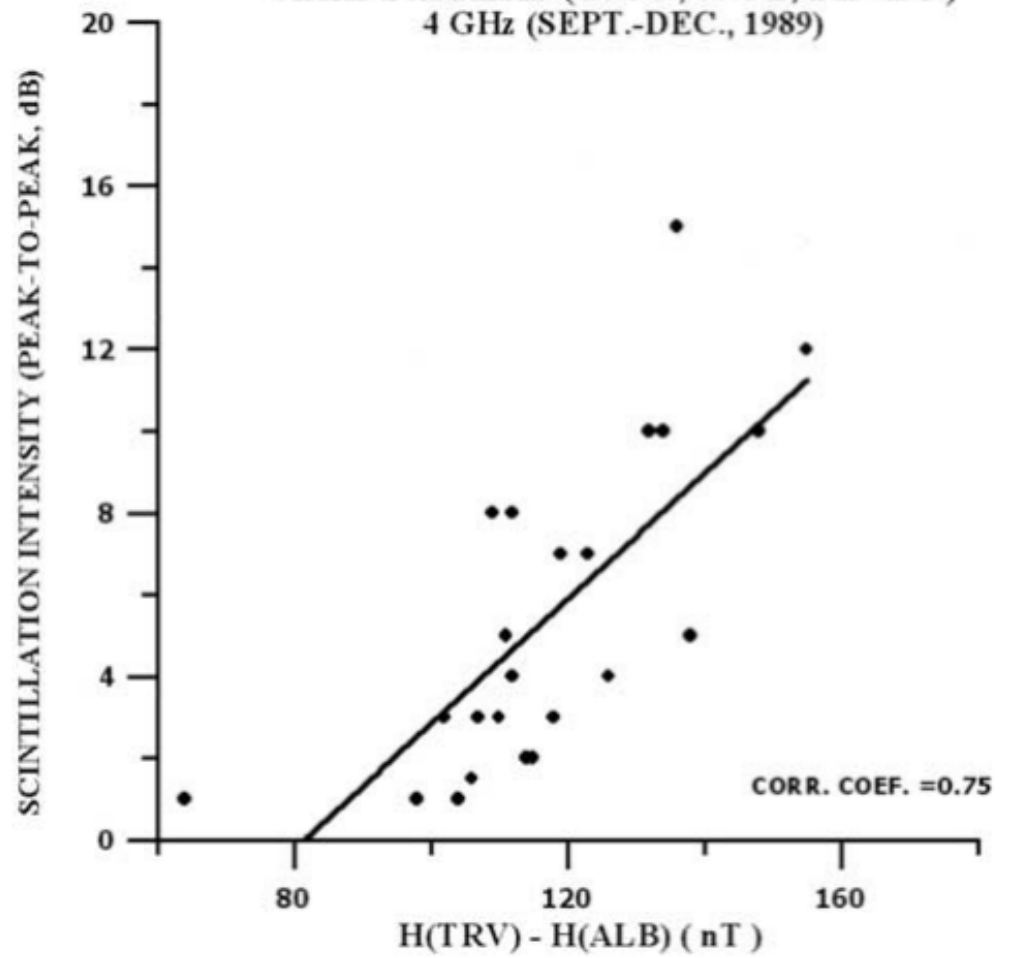
The importance of equatorial magnetic observation



Equatorial Plasma Bubbles

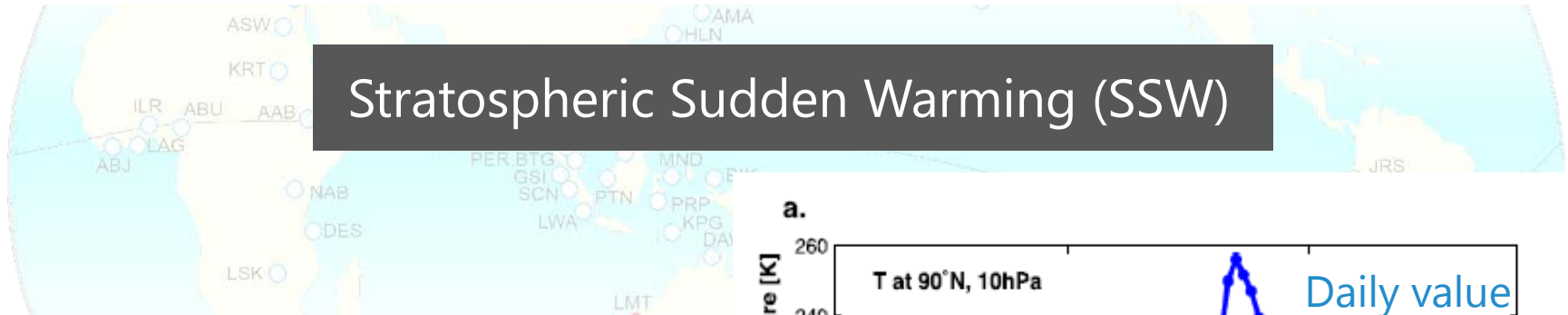


SIKANDARABAD (26.8 N; 77.8 E; DIP 42 N)
4 GHz (SEPT.-DEC., 1989)



[Dabas et al., 2003]

The importance of equatorial magnetic observation



Stratospheric Sudden Warming (SSW)

Solar wind

Magnetosphere

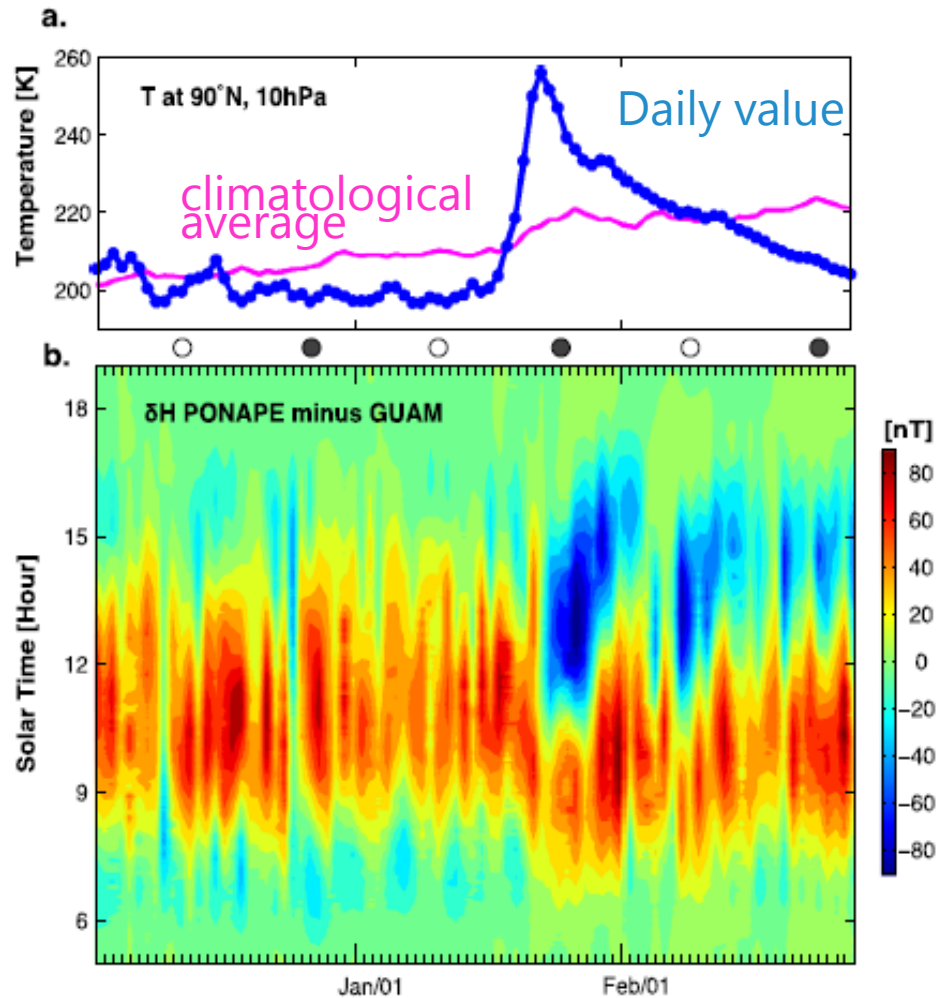
Current

Ionosphere

Atmospheric wave



Ground Magnetometer

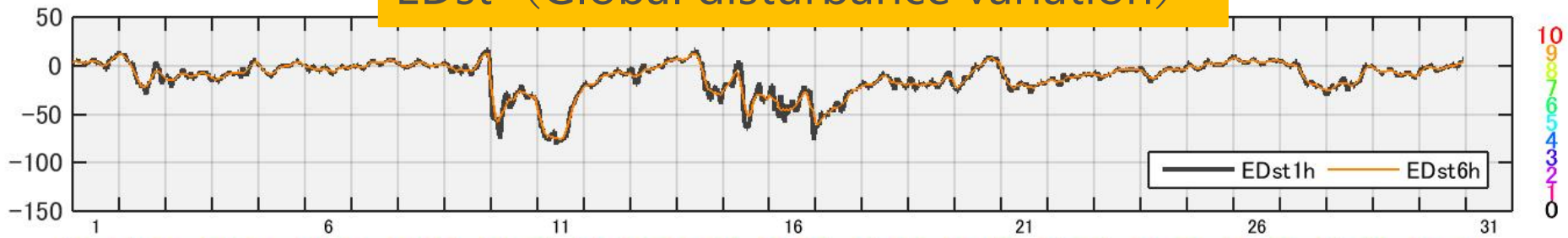


Morphology of EEJ during magnetically quiet periods

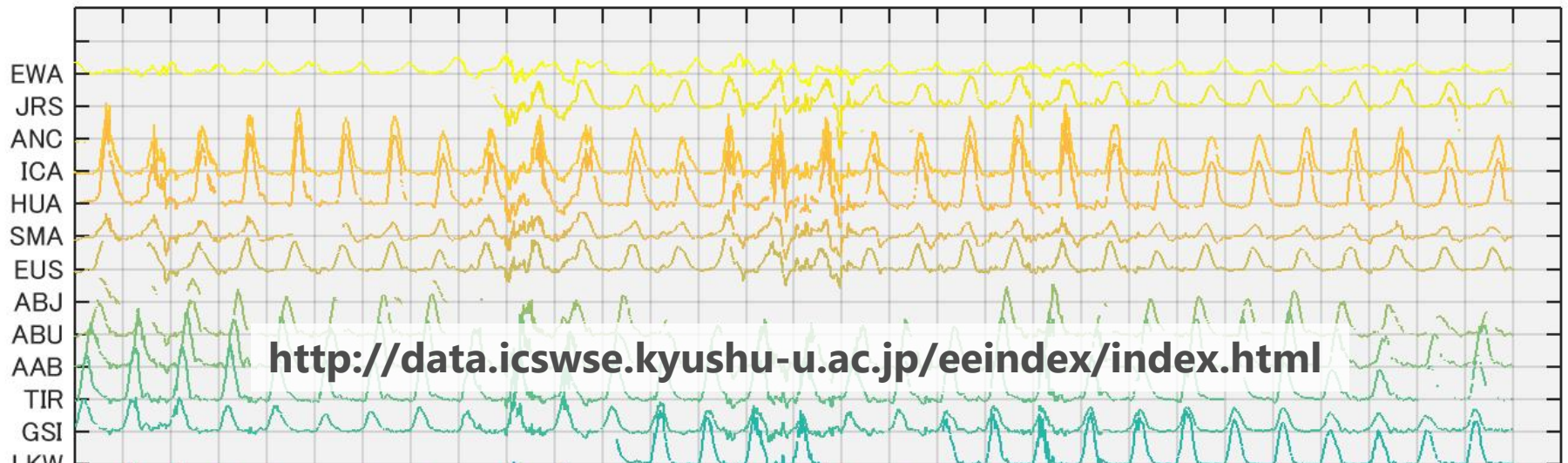
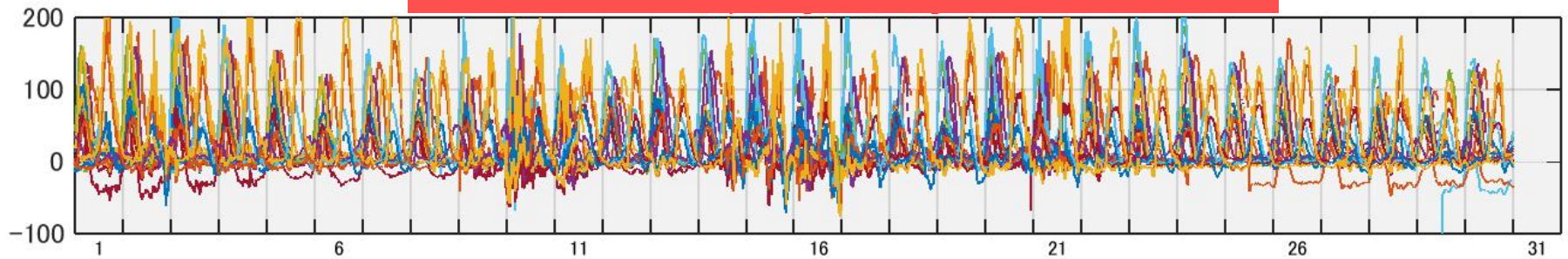
- Dependence of the EEJ intensity on the solar activity
 - 1976: Rastogi and Iyer
 - 1994: Rastogi et al.
- Semi-annual variations of the EEJ intensity
 - 1965: Chapman and Raja Rao
 - 1966: Yacob
 - 1980: Campbell
- Day-to-day variability in the EEJ intensity
 - 1976: Fambitakoya and Mayaud
 - 1980: Kane and Trivedi
 - 1998: Doumouya et al.
 - 2008: Kawano-sasaki and Miyahara
- Relationship of the EEJ to the Sq current system
 - 1992: Onwumechili
 - 1995: Stening

Real-time EE-index

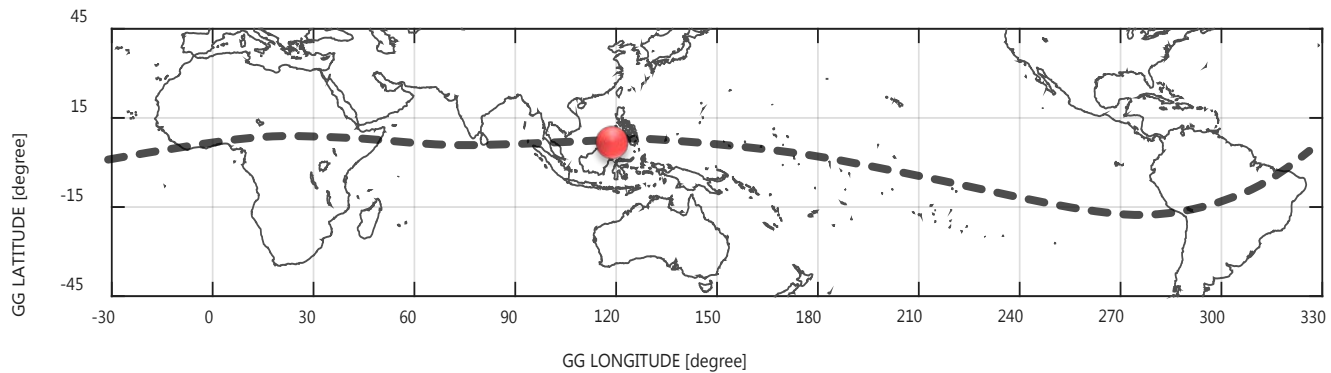
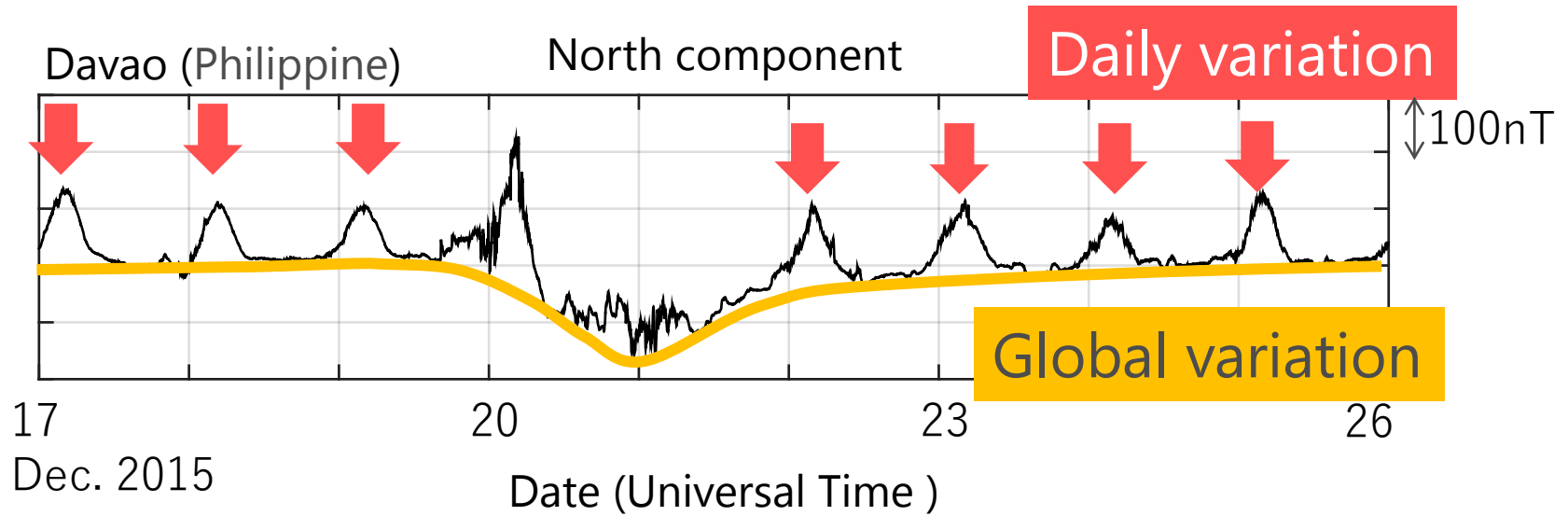
EDst (Global disturbance variation)



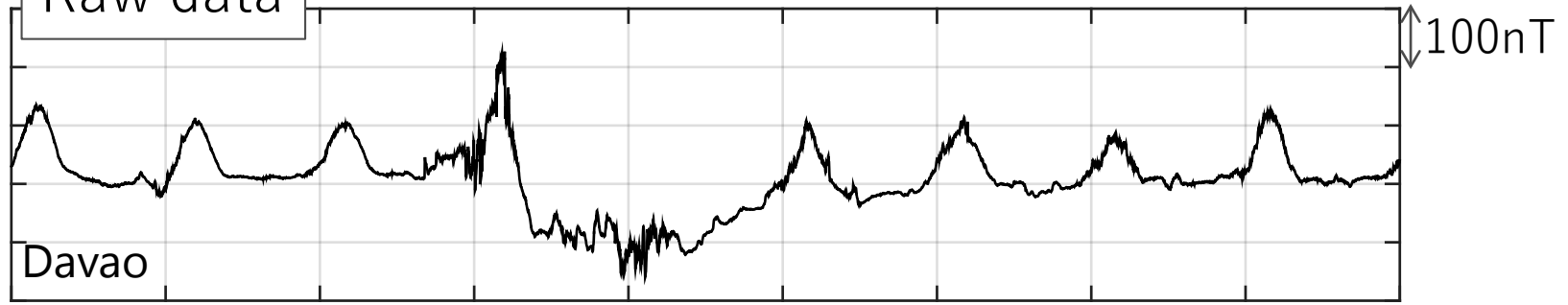
EUEL (Local disturbance variation)



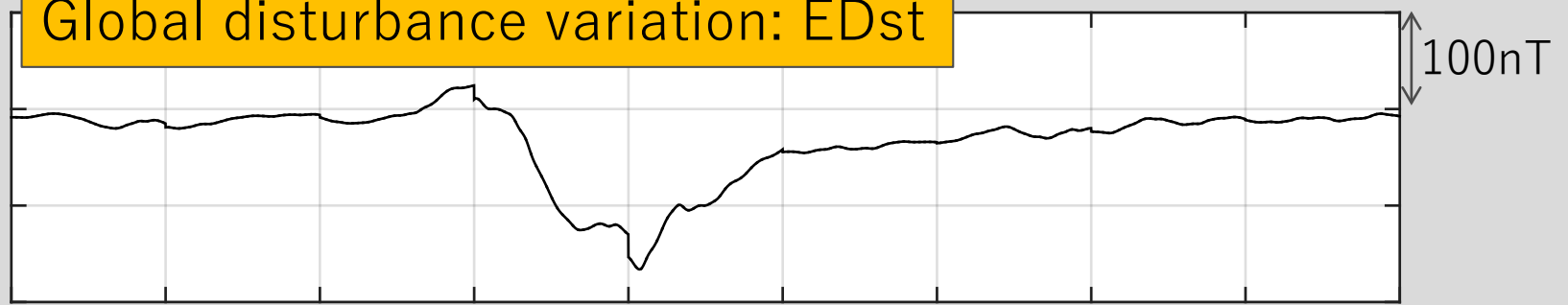
Time series data of magnetic field along the equator



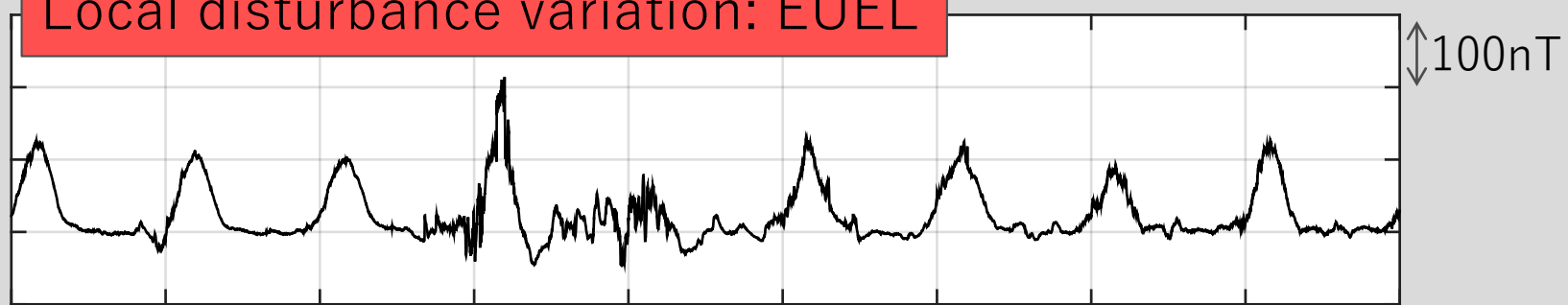
Raw data



Global disturbance variation: EDst



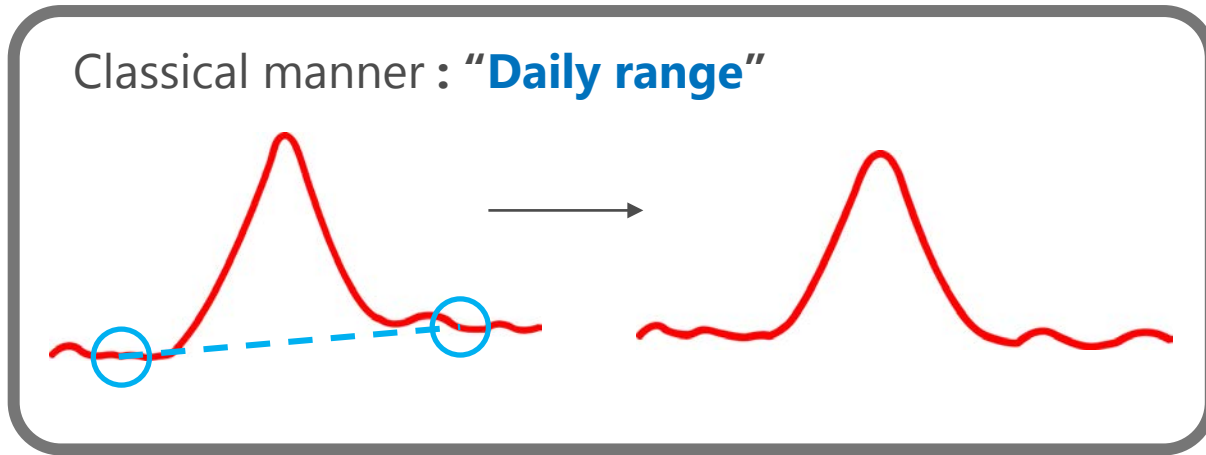
Local disturbance variation: EUEL



EE-index

17 Dec. 2015 20 23 26

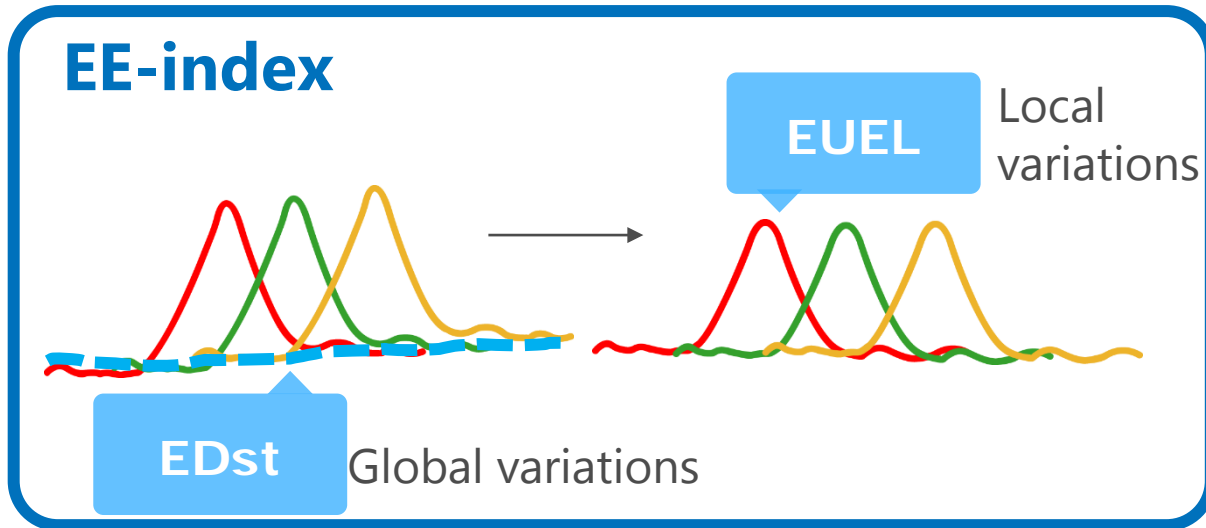
Classical manner vs. EE-index on the EEJ study



Static base line



Quiet time



Drastic base line



Quiet &
Disturbance time

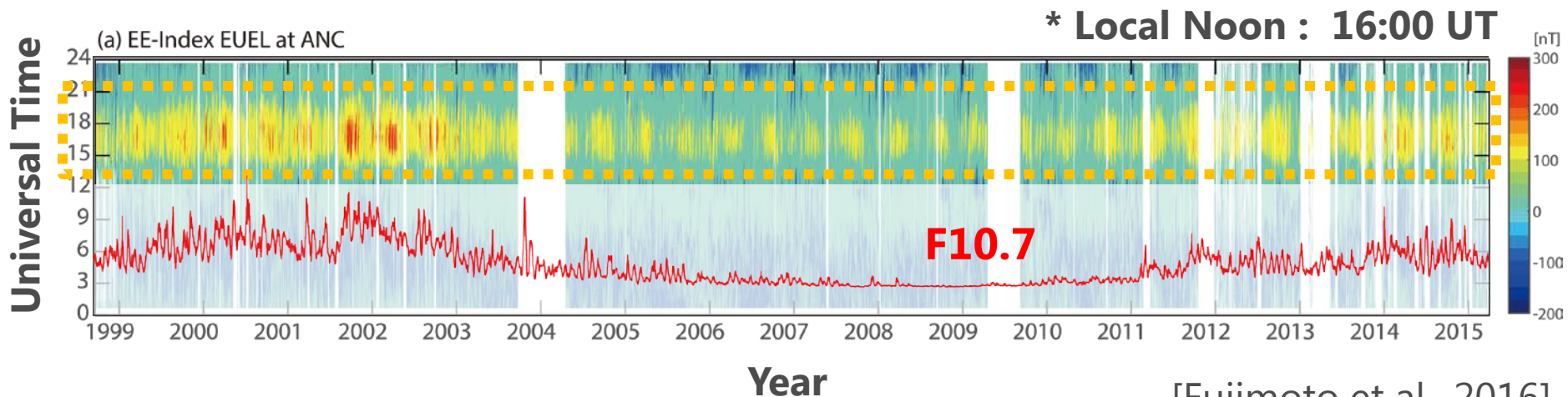


Time series analysis

long-term variation of EUEL from 1998 to 2015

EUEL of EE-index @ Ancon (**ANC**) in Peru

- 1 min. sampling
- period : 1998/09/18 – 2015/03/31
- Hourly averaged EUEL intensity



[Fujimoto et al., 2016]

Data & Analysis

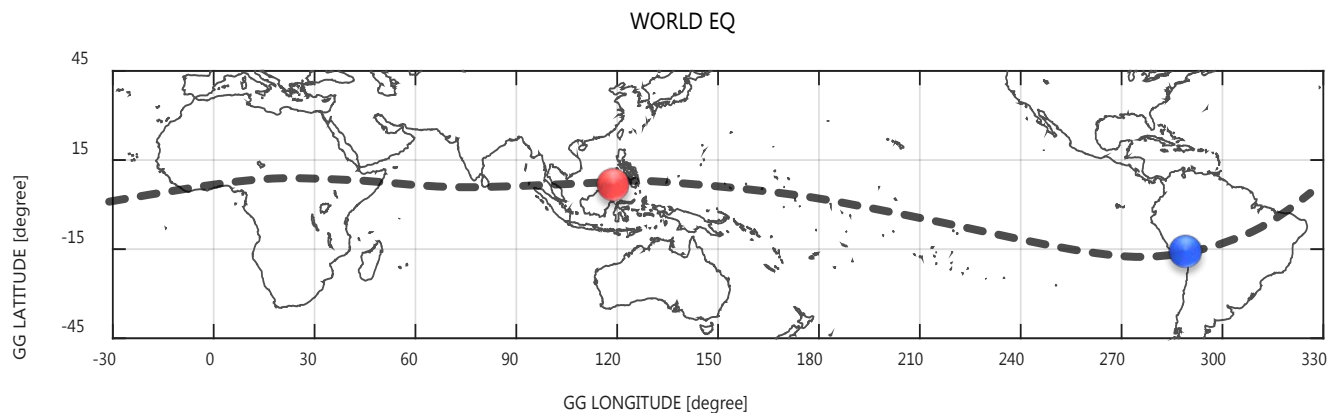
Magnetometer Stations : Ancon (Peru) , Davao (Philippine)

Period : 2004 - 2009 (Ancon), 2005 – 2010 (Davao)

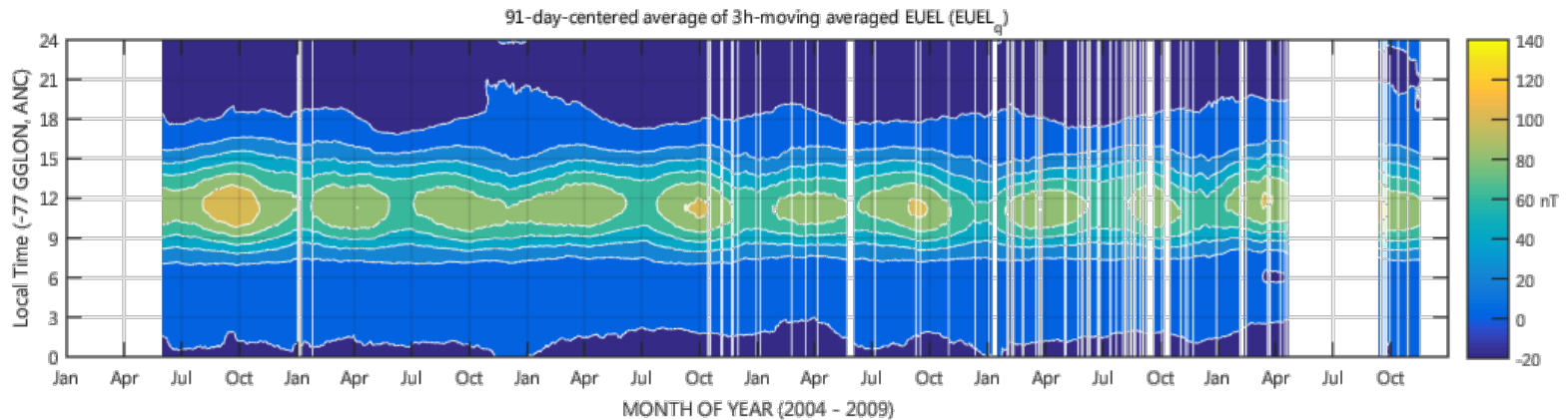
(solar cycle 23 : 1996, Jun. – 2008, Jan. ☒Min : 2008, Mar.)

Data requirements : no lack data in one day

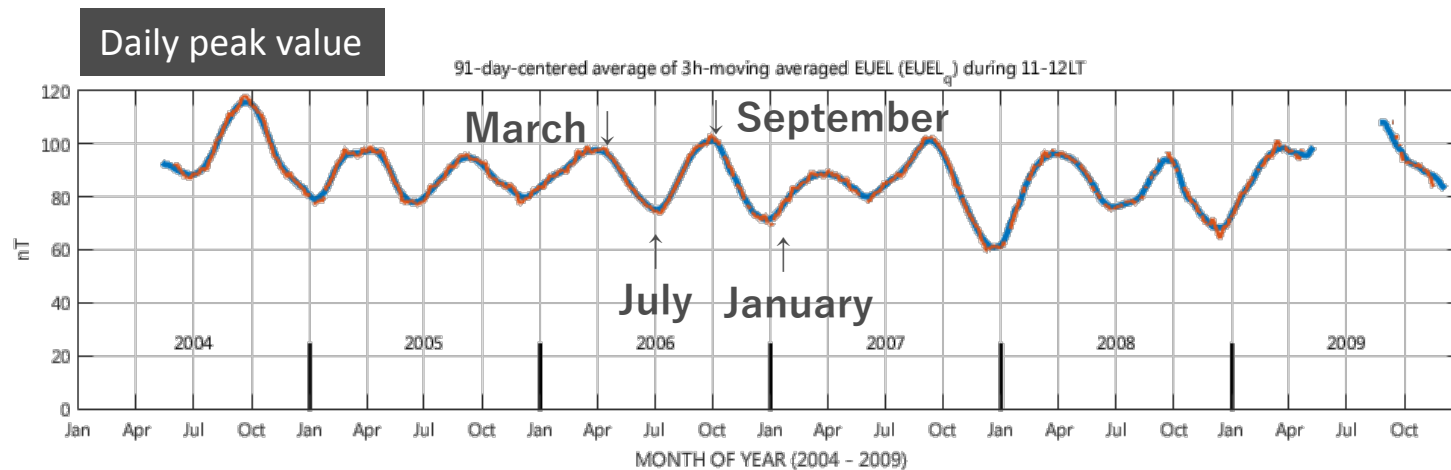
Analysis: Semiannual and Semidiurnal of EUEL



Semiannual variation (Ancon)

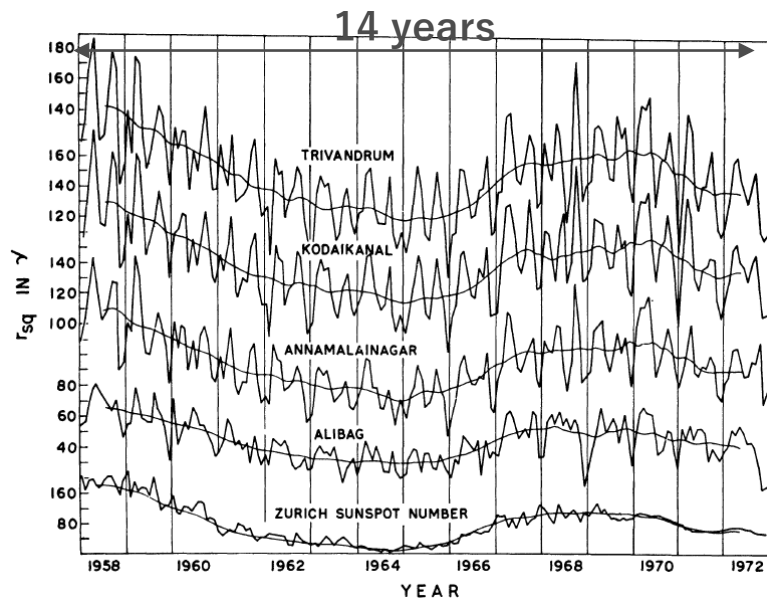
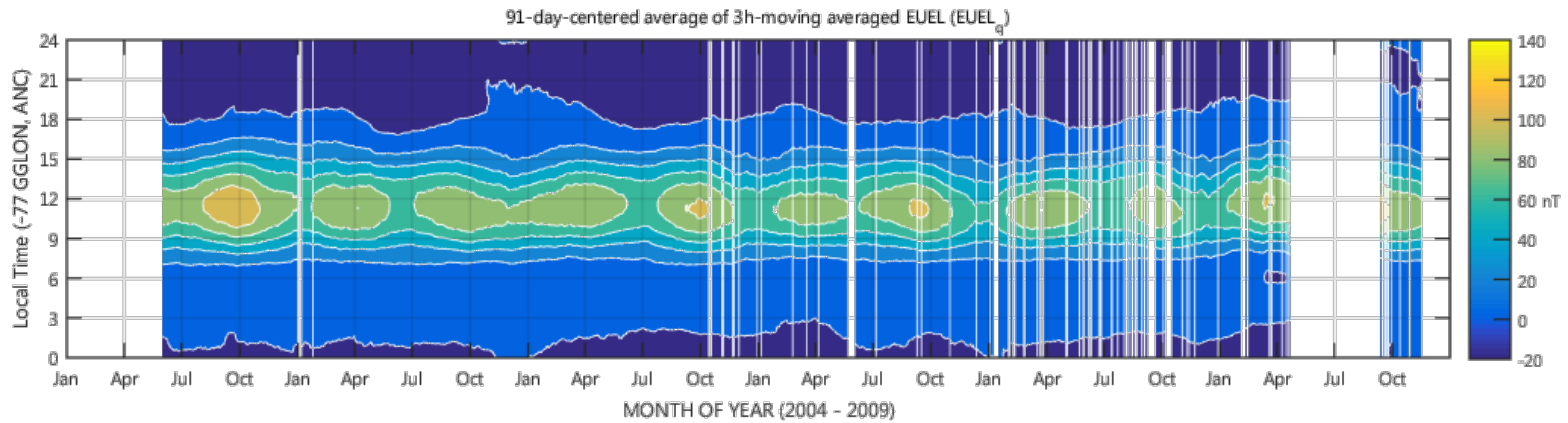


The contour shows the 91-day-centered average of 3-hour moving averaged EUEL. White color path indicates the lack data.



The daily peak value is the maximum between 11:00 and 12:00 local time of the data on the upper panel.

Semiannual variation (Ancon)



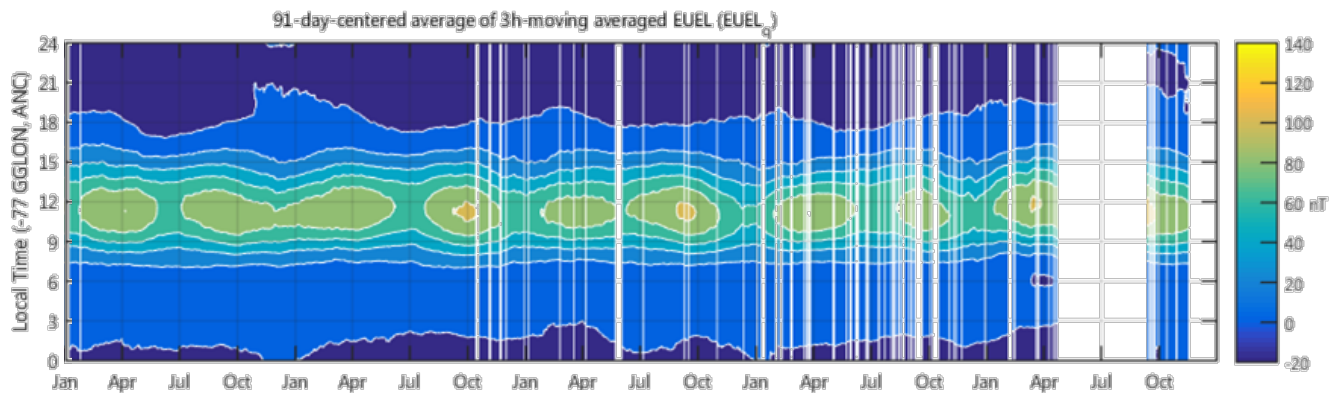
[Rastogi and Iyer, 1976 (JGG)]

The semiannual variation based on EE-index (including the quite/disturbed days), is consistent with the result of Rastogi and Iyer (1976).

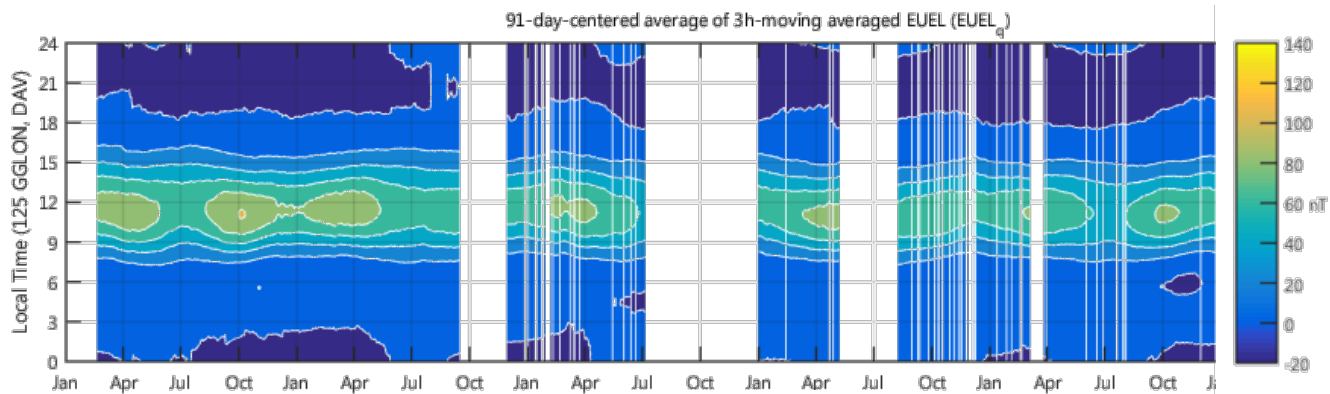
- There is a significant **semiannual variation** with maxima around March and September.
- The semiannual peak value follows the solar activity.

Semiannual variation (2005-2010)

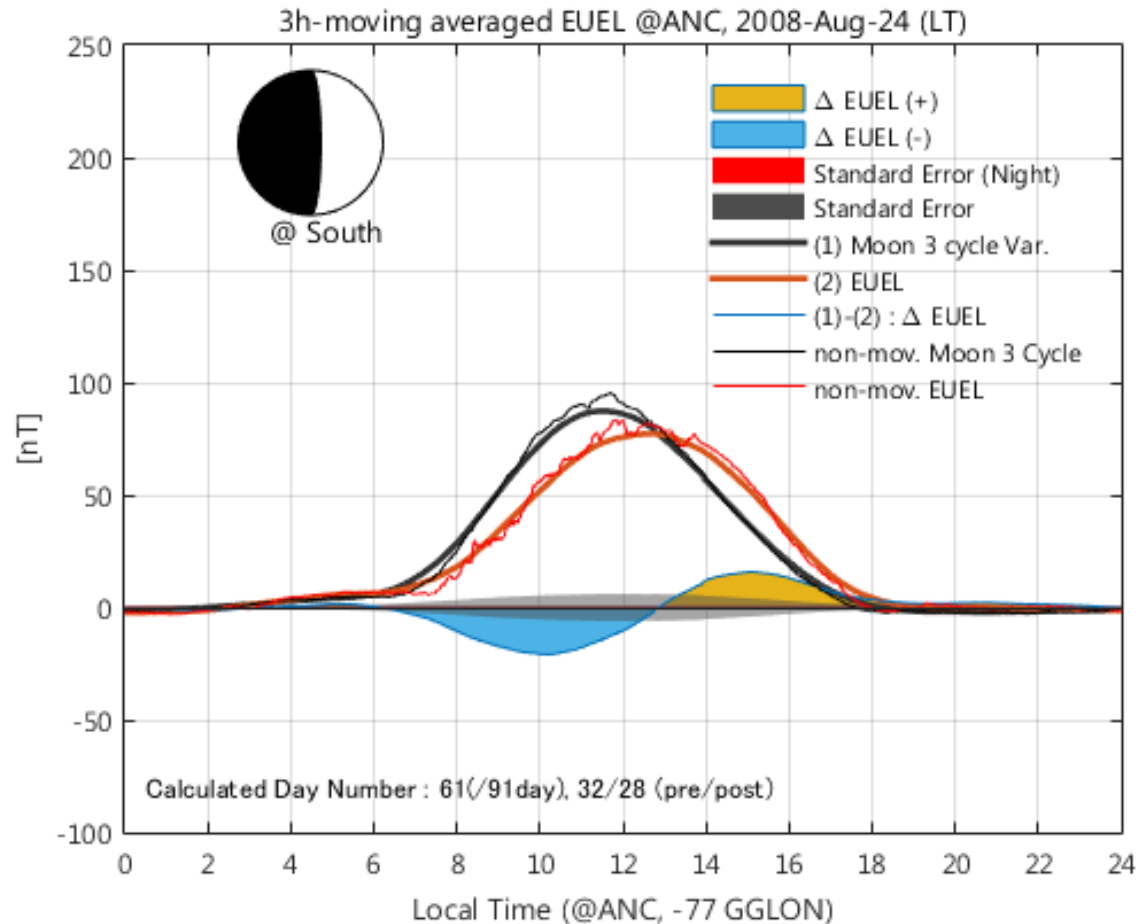
Ancon



Davao



Semidiurnal variation ($\Delta EUEL$)



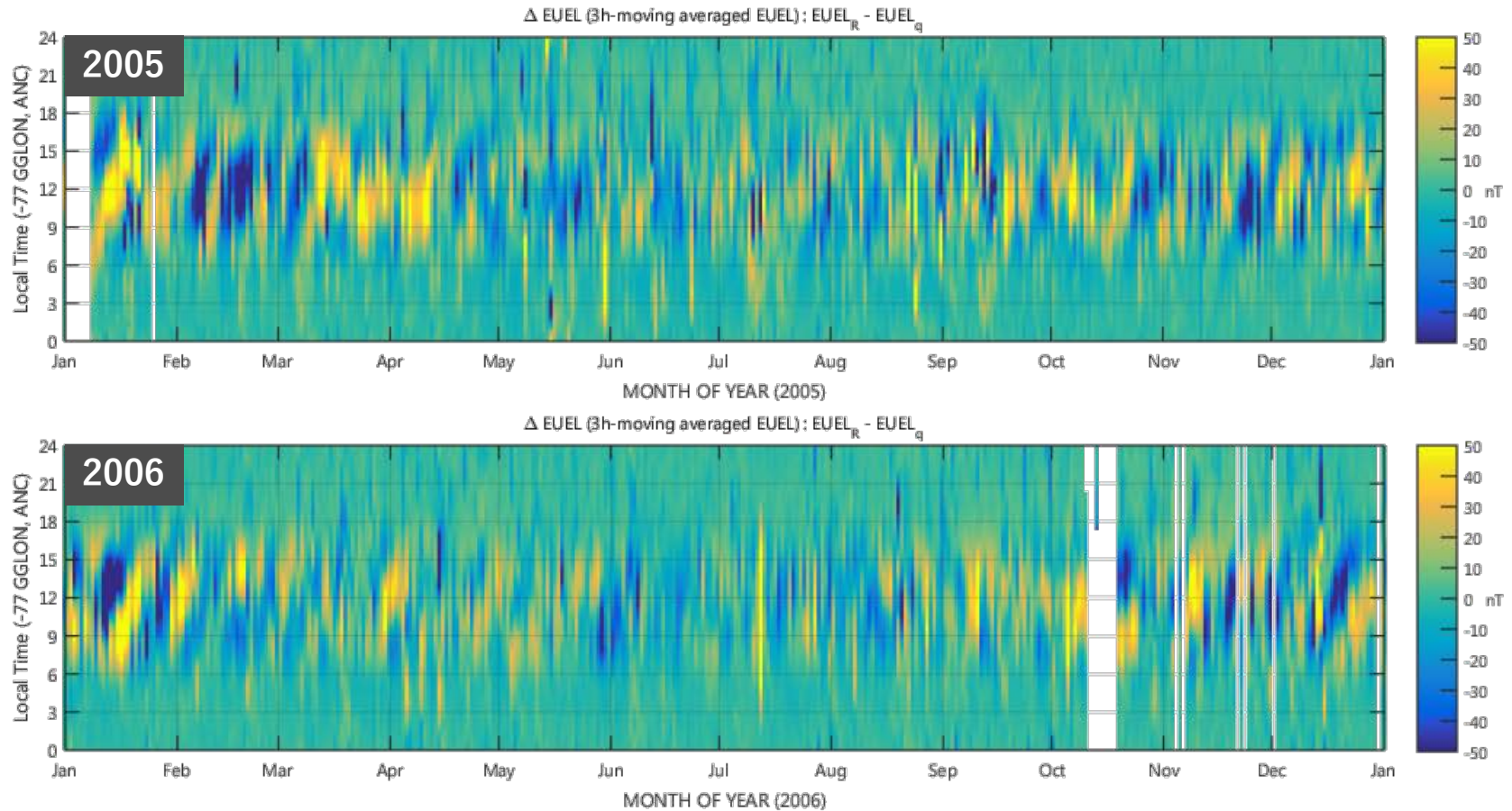
$EUEL_{3h}$: 3 hour moving averaged $EUEL$

\overline{EUEL} : 91 days entered average of $EUEL_{3h}$

$$\Delta EUEL = EUEL_{3h} - \overline{EUEL}$$

Semidiurnal variation (Ancon)

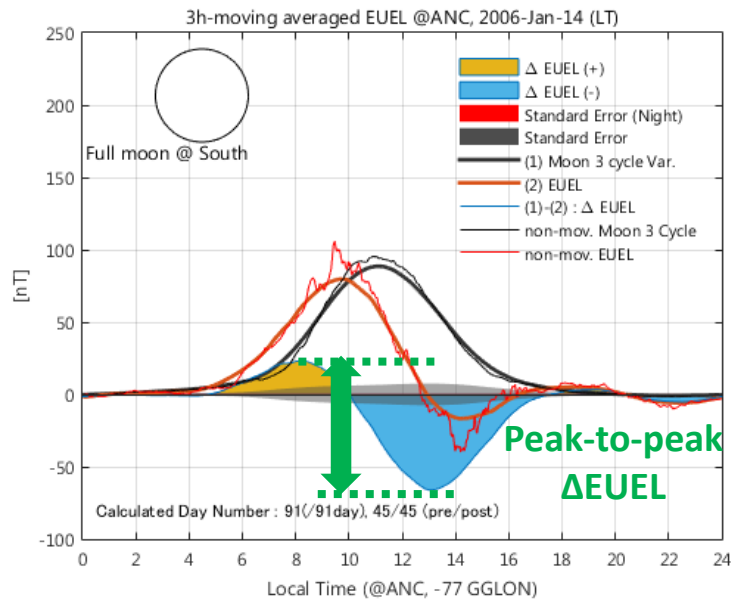
The amplitude of semidiurnal EUEL variations increased in January and decreased around July.



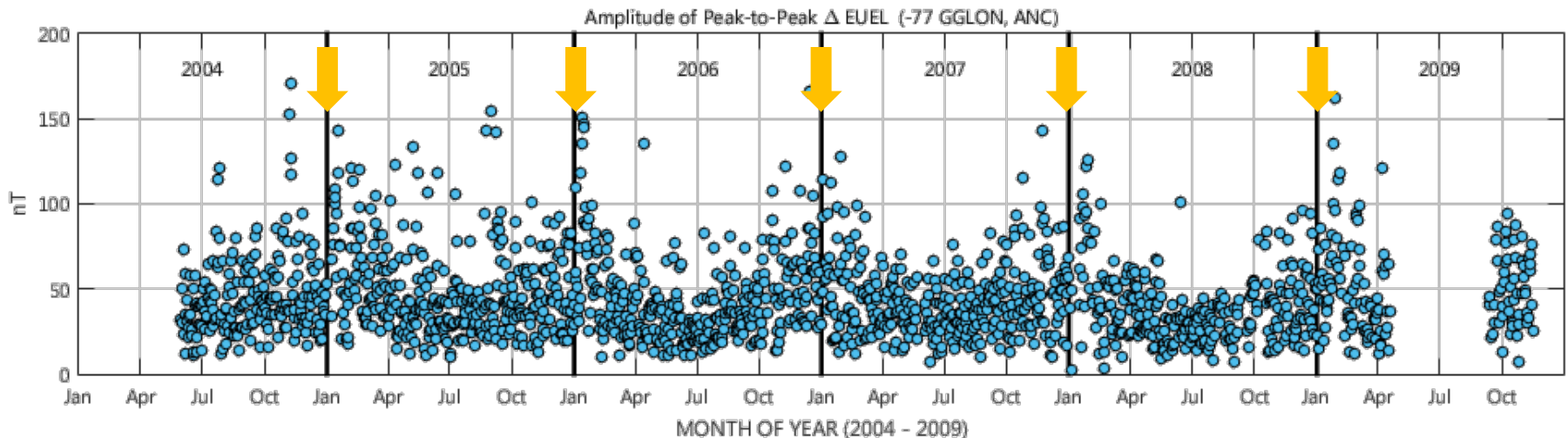
$\Delta EUEL$ is obtained by subtracting \overline{EUEL} from $EUEL_{3h}$.

The yellow color and blue color indicate the positive and negative, respectively.

Semidiurnal variation (Ancon)

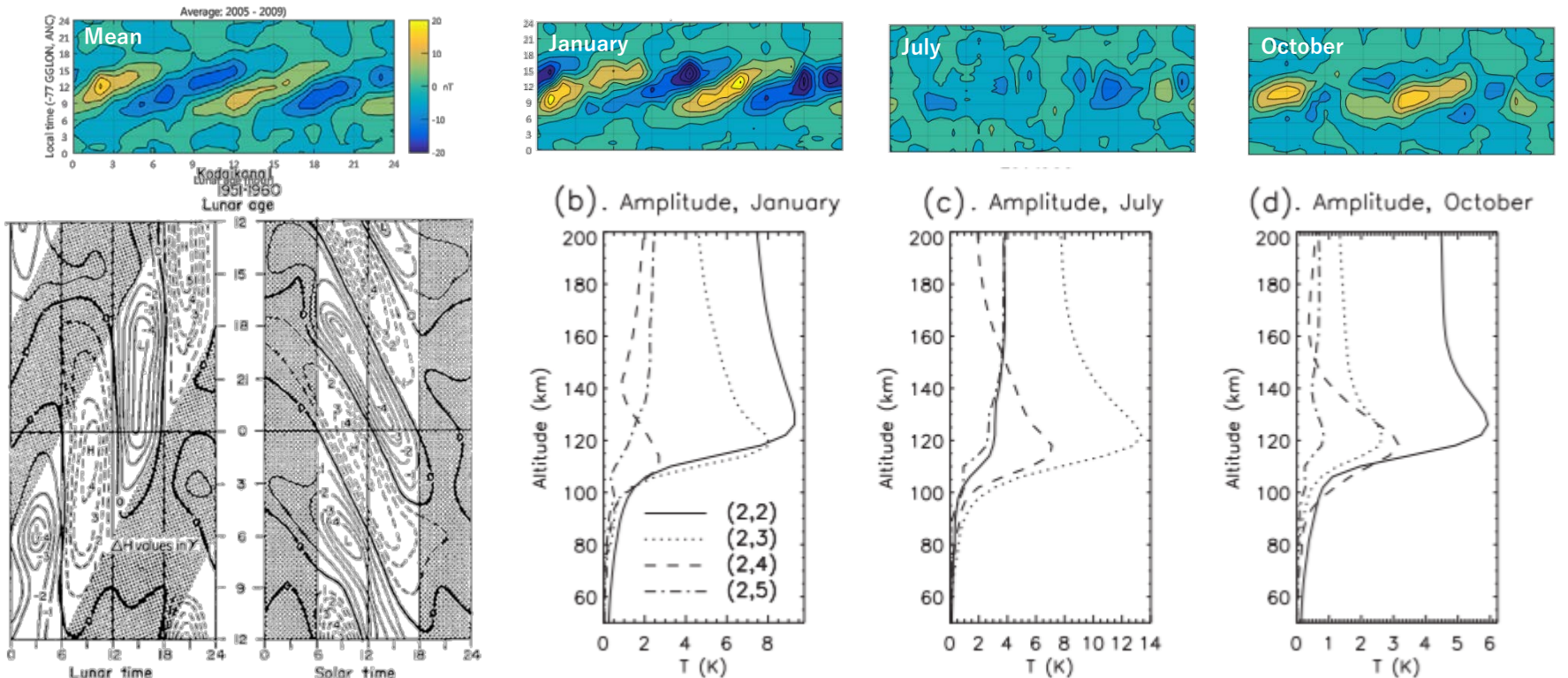


The peak-to-peak (indicated by the green arrows) is the difference between the maximum and minimum of Δ EUEL.



Seasonal dependence of Semidiurnal Var.

- The mean behavior of $\Delta E U E L$ is consistent with the result of Rastogi (1973). We demonstrated the monthly average behavior of $\Delta E U E L$, for the first time based on the time-series magnetometer data.
- The seasonal dependence of semidiurnal variation agrees with the seasonal profile of atmospheric neutral wind (2.2) mode corresponding to the lunar tide.



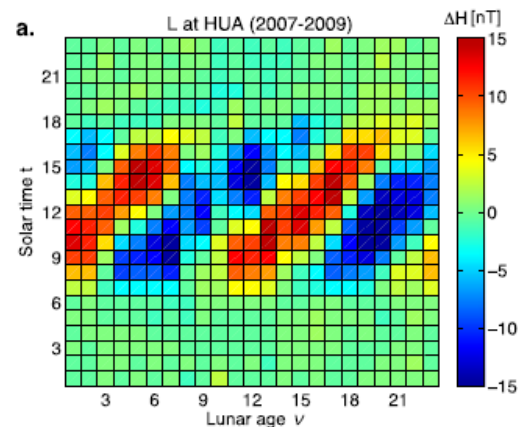
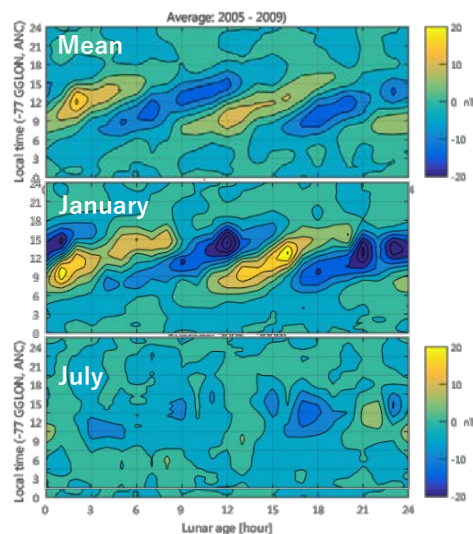
[Rastogi, 1973, PSS]

[Forbes et al., 2013]

Summary

- The latest results on the long-term study of the EE-index:
 - solar cycle variation
 - semiannual variation
 - semidiurnal variation
- The remarkable seasonal dependence of semidiurnal variation :
stronger in January and weaker around July

The seasonal dependence of semidiurnal variation agrees with the seasonal profile of atmospheric neutral wind (2.2) mode corresponding to the lunar tide.



● ○ ● ● ●
[Yamazaki and Maute, 2016]

EE-index : Monitoring index for equatorial electrojet (EEJ)

- 1 Proposed by **ICSWSE** in 2008 [Uozumi et al., 2008]
First Version : 4 stations along the magnetic equator
Latest version : multiple equatorial magnetometer data
[Fujimoto et al., 2016]
- 2 Produced by using **MAGDAS/CPMN** magnetometer network



- 3 Be useful for the study on the long-term time series analysis of the magnetic field along the magnetic equator

We can evaluate the equatorial magnetic field variation with the same ruler during the unquiet time as well as quiet time

<http://data.icswse.kyushu-u.ac.jp/eeindex/index.html>