

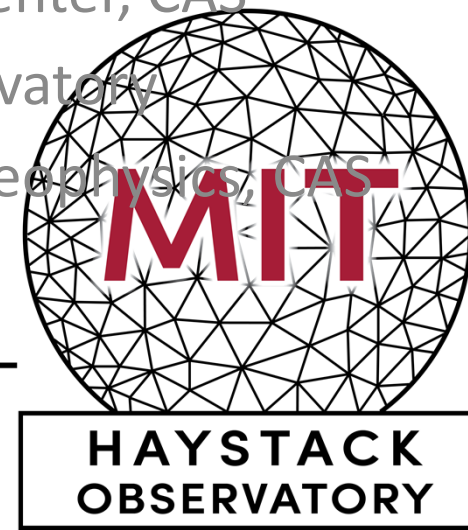
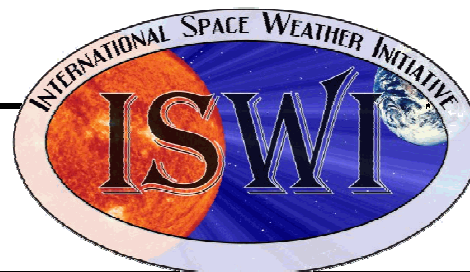
**Monitoring and investigation of geospace
disturbances along the 120E/60W longitudes:
International Meridian Circle Project**

Shunrong Zhang¹, MIT Haystack Observatory

Chi Wang², National Space Science Center, CAS

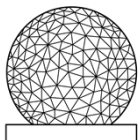
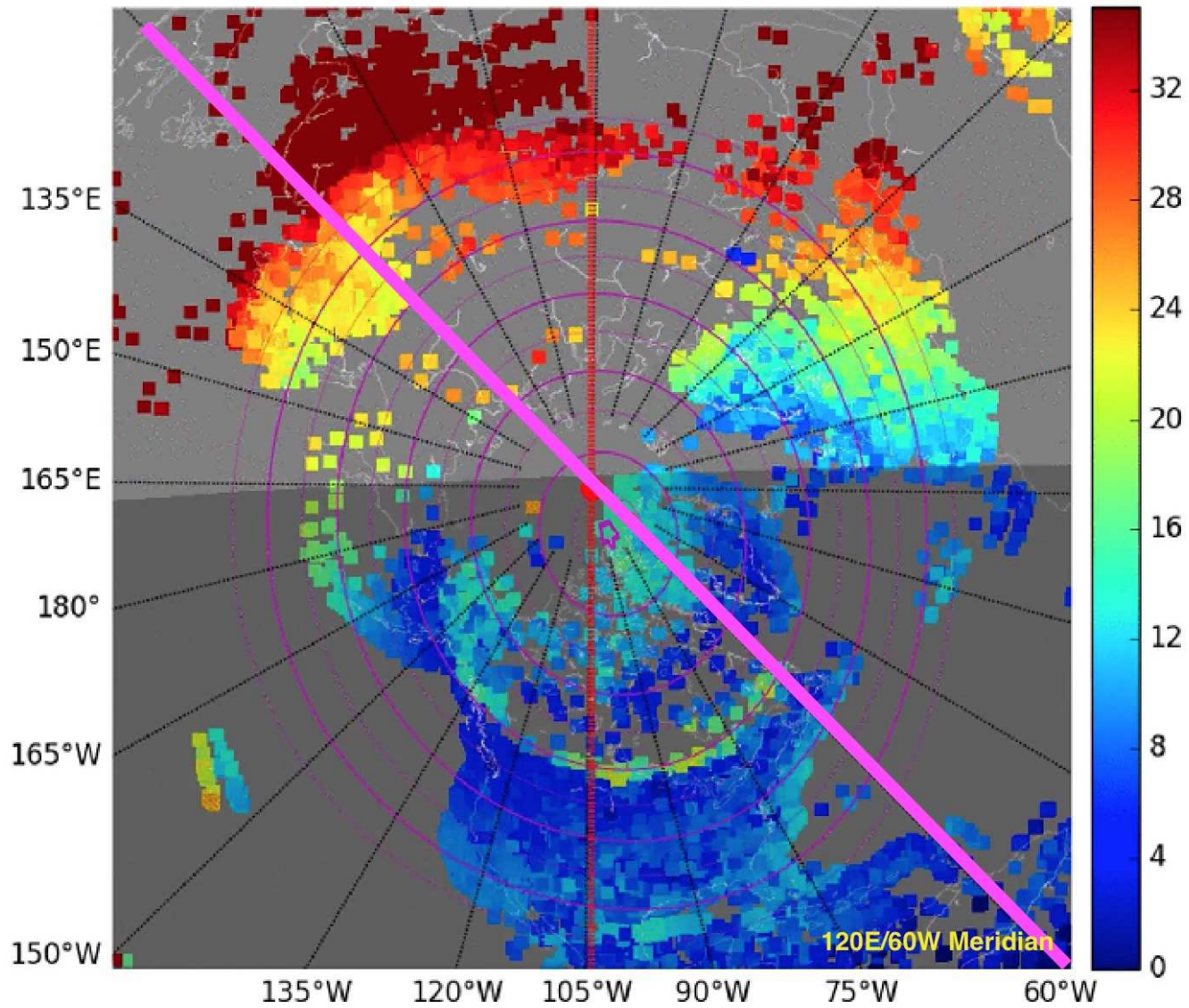
John Foster¹, MIT Haystack Observatory

Weixing Wan³, Institute of Geology and Geophysics, CAS



GNSS TEC MAP

UTC 2015-03-17 06:57:00

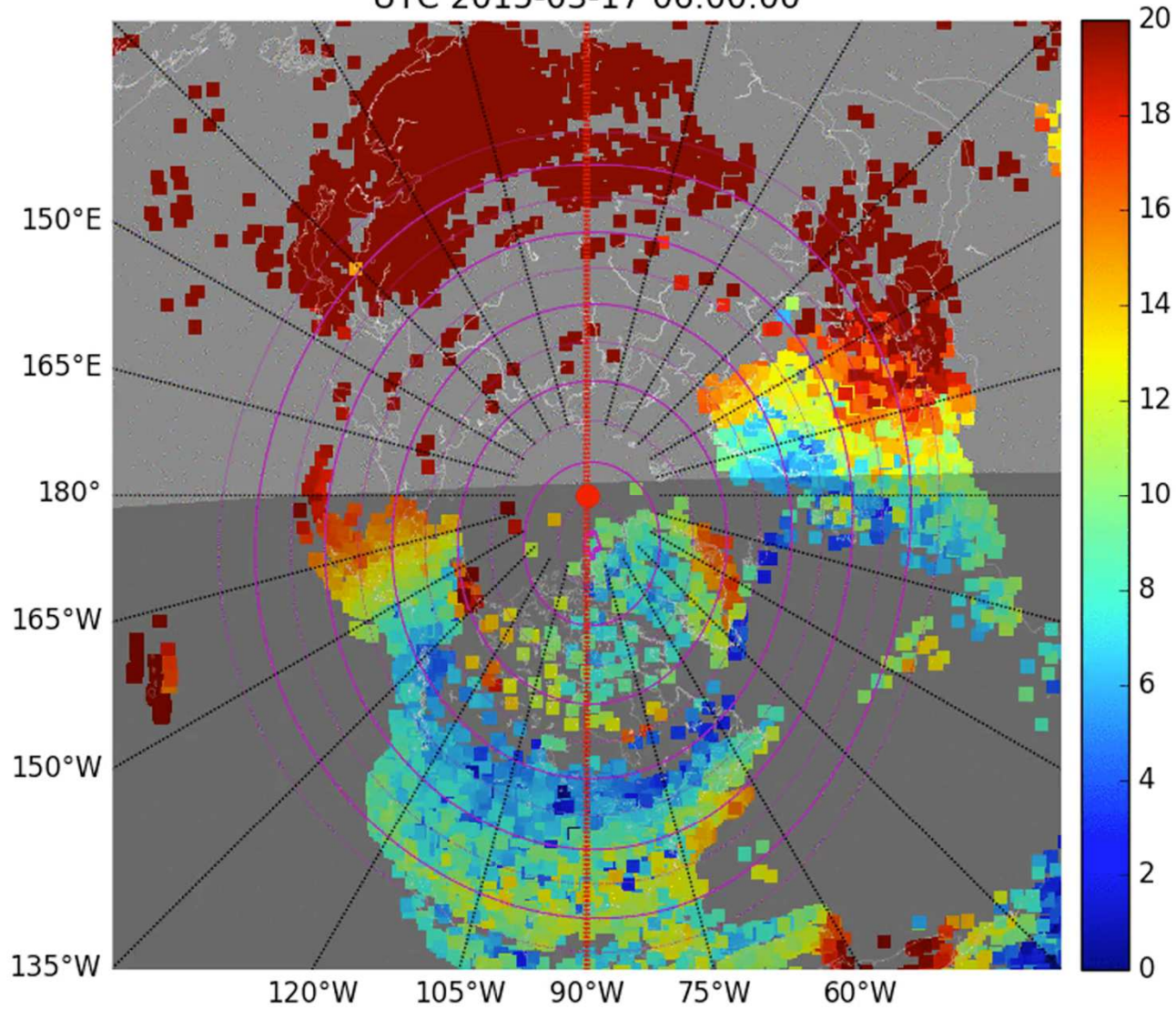


MIT
HAYSTACK
OBSERVATORY

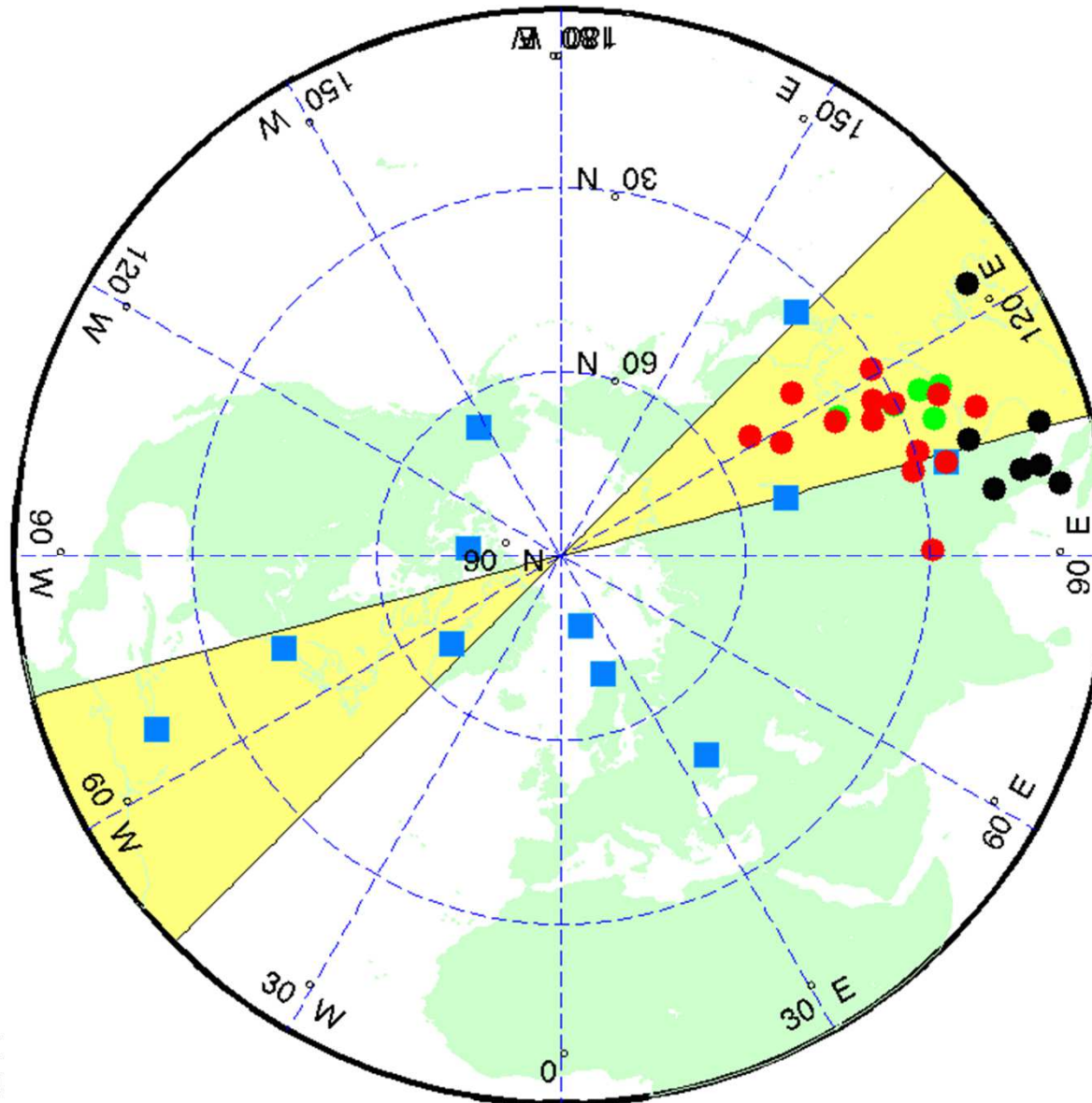
MIT Massachusetts
Institute of
Technology

GNSS TEC MAP (movie)

UTC 2015-03-17 06:00:00

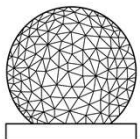
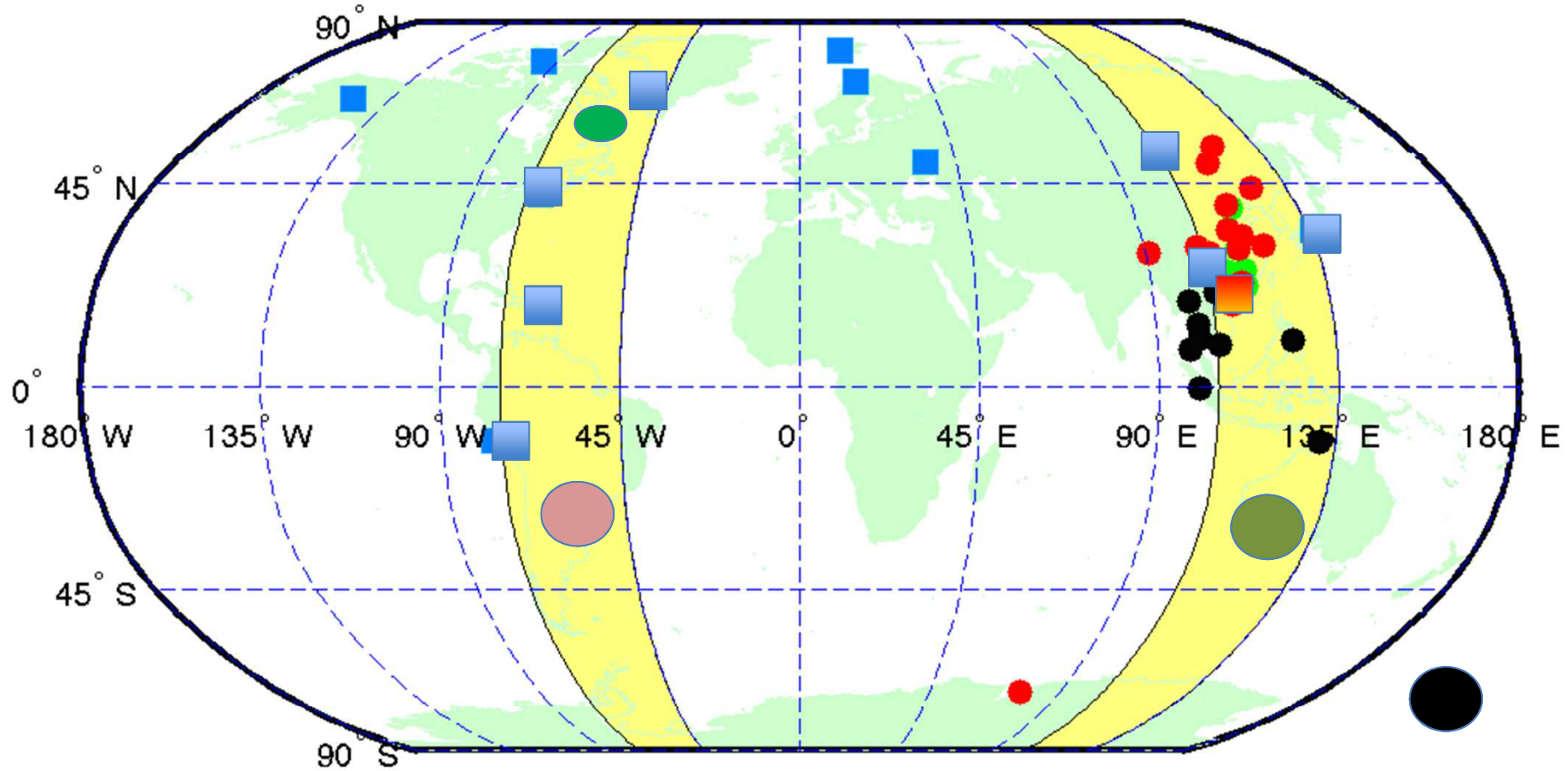


Meridian Circle International Observation: A **Partial** Network (polar view)



Meridian Circle International Observation: A Partial Network

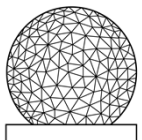
World ISRs and MP sites



MIT
HAYSTACK
OBSERVATORY

MIT
Massachusetts
Institute of
Technology

Soundrestrom ISR



MIT
HAYSTACK
OBSERVATORY

MIT
Massachusetts
Institute of
Technology

Millstone Hill ISR



N_e

T_e

T_i

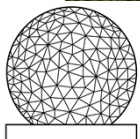
V_0

Ion composition

$T_n \leftrightarrow T_{ex}$

EXB

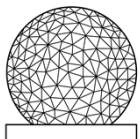
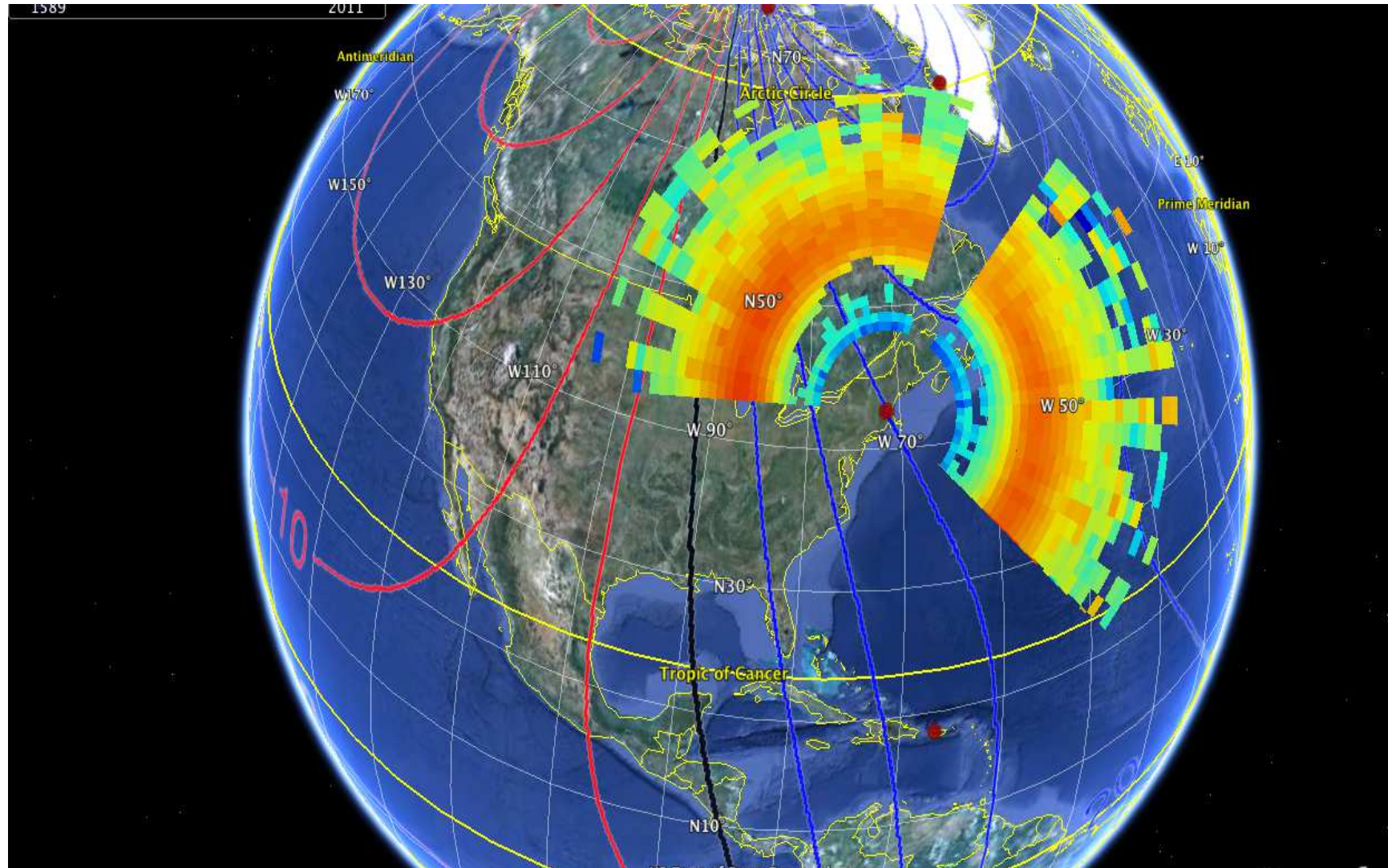
U



MIT
HAYSTACK
OBSERVATORY

MIT Massachusetts
Institute of
Technology

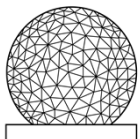
MISA: Super-wide Coverage



MIT
HAYSTACK
OBSERVATORY

MIT
Massachusetts
Institute of
Technology

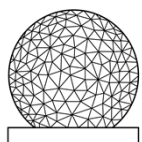
Arecibo ISR



MIT
HAYSTACK
OBSERVATORY

MIT Massachusetts
Institute of
Technology

Jicamarca ISR

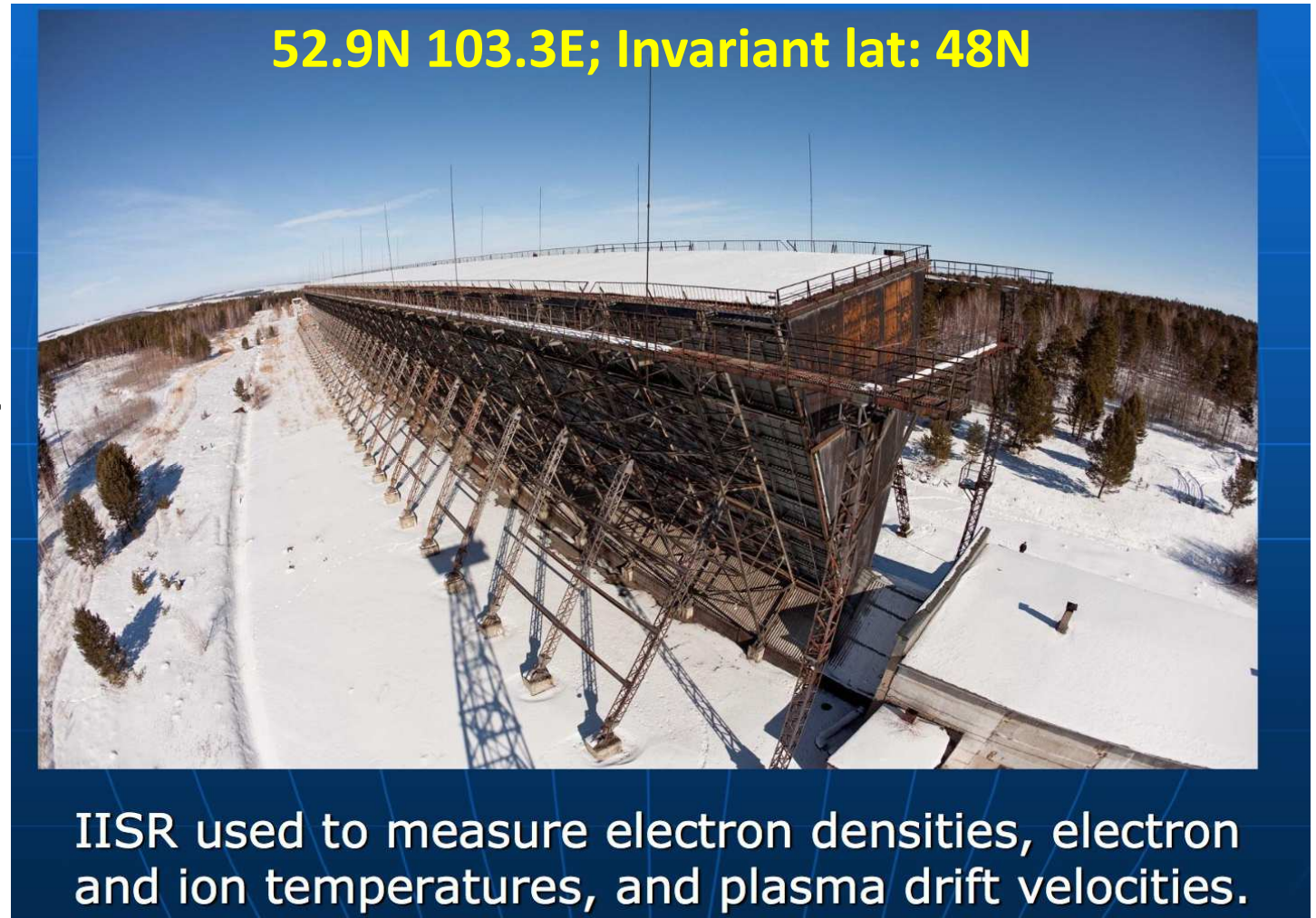


MIT
HAYSTACK
OBSERVATORY

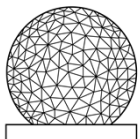
MIT
Massachusetts
Institute of
Technology

Irkutsk Incoherent Scatter Radar (IISR)

- Horn Antenna
246x12 m
- 2 sub-horns
- $f=154-162$ MHz
- Frequency beam steering
- Polarization filter
- $P_t < 3$ MW
- Beam 0.5×10 deg



Medvedev et al. (ISTP)

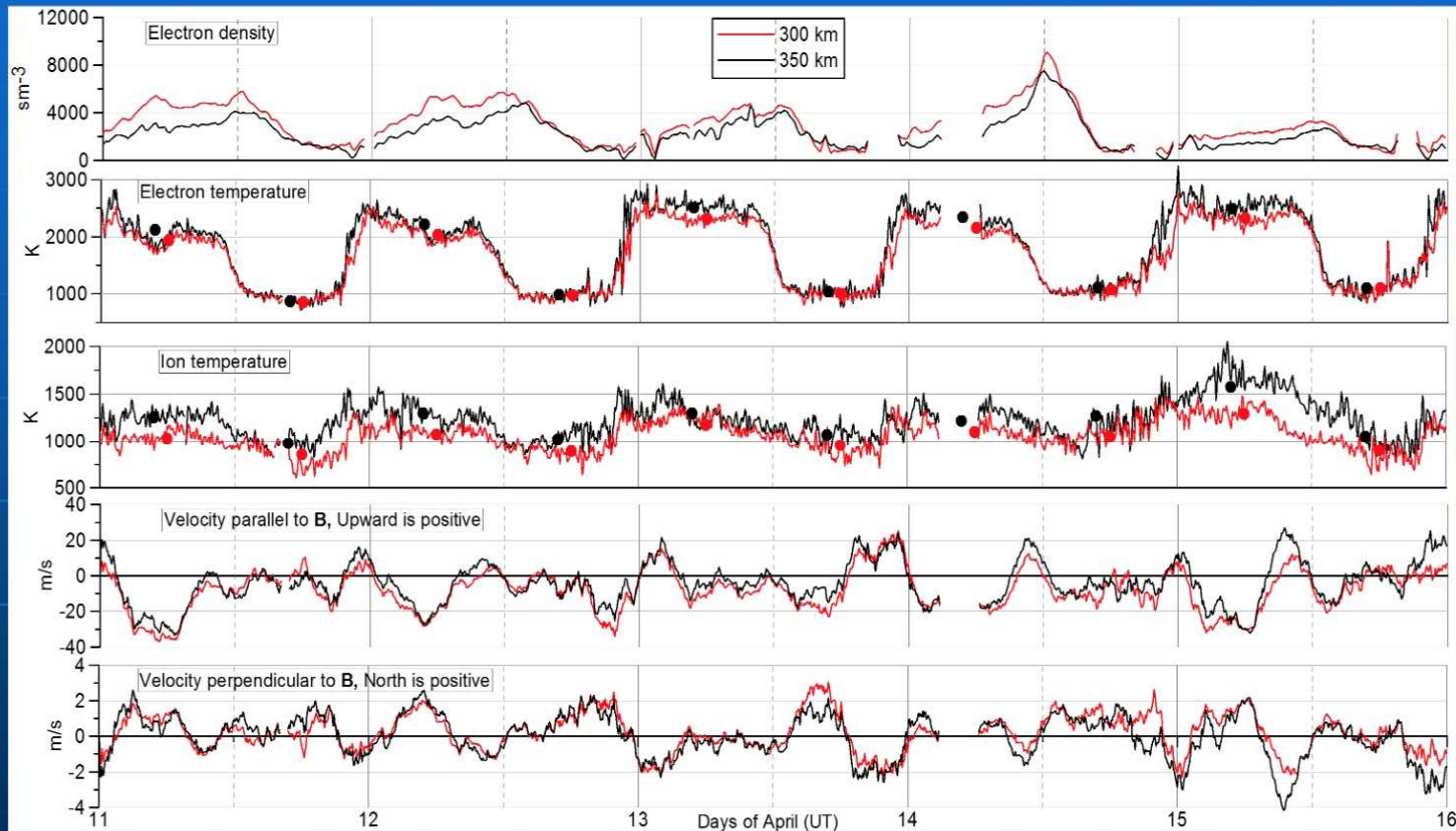


MIT
HAYSTACK
OBSERVATORY

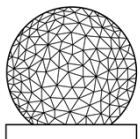
MIT
Massachusetts
Institute of
Technology

IISR Observation

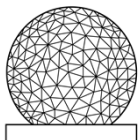
Day-to-day variations of electron density, electron temperature, ion temperature, and drift velocities at 300 and 350 km (Irkutsk ISR data)



Dusk effect is seen in (1) electron density enhancement, (2) decrease (relative to other days at 12 UT) in electron temperature, and (3) increase in upward drift velocity.

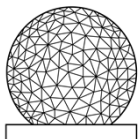
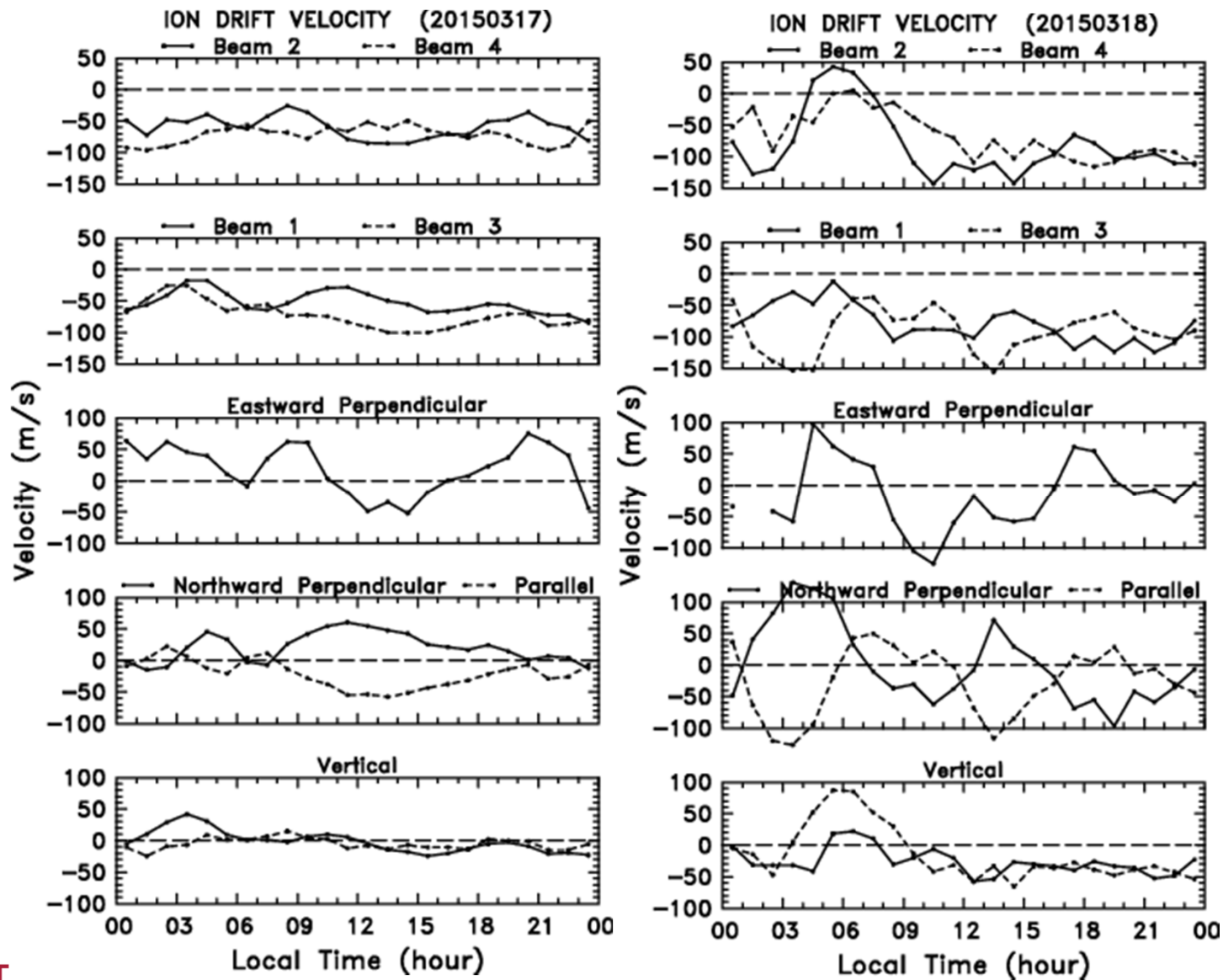


MU radar (Japan)



MIT
HAYSTACK
OBSERVATORY

MU Radar plasma drifts during 2015 St Patrick's Day



Qujing Incoherent Scatter Radar

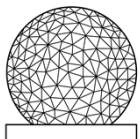
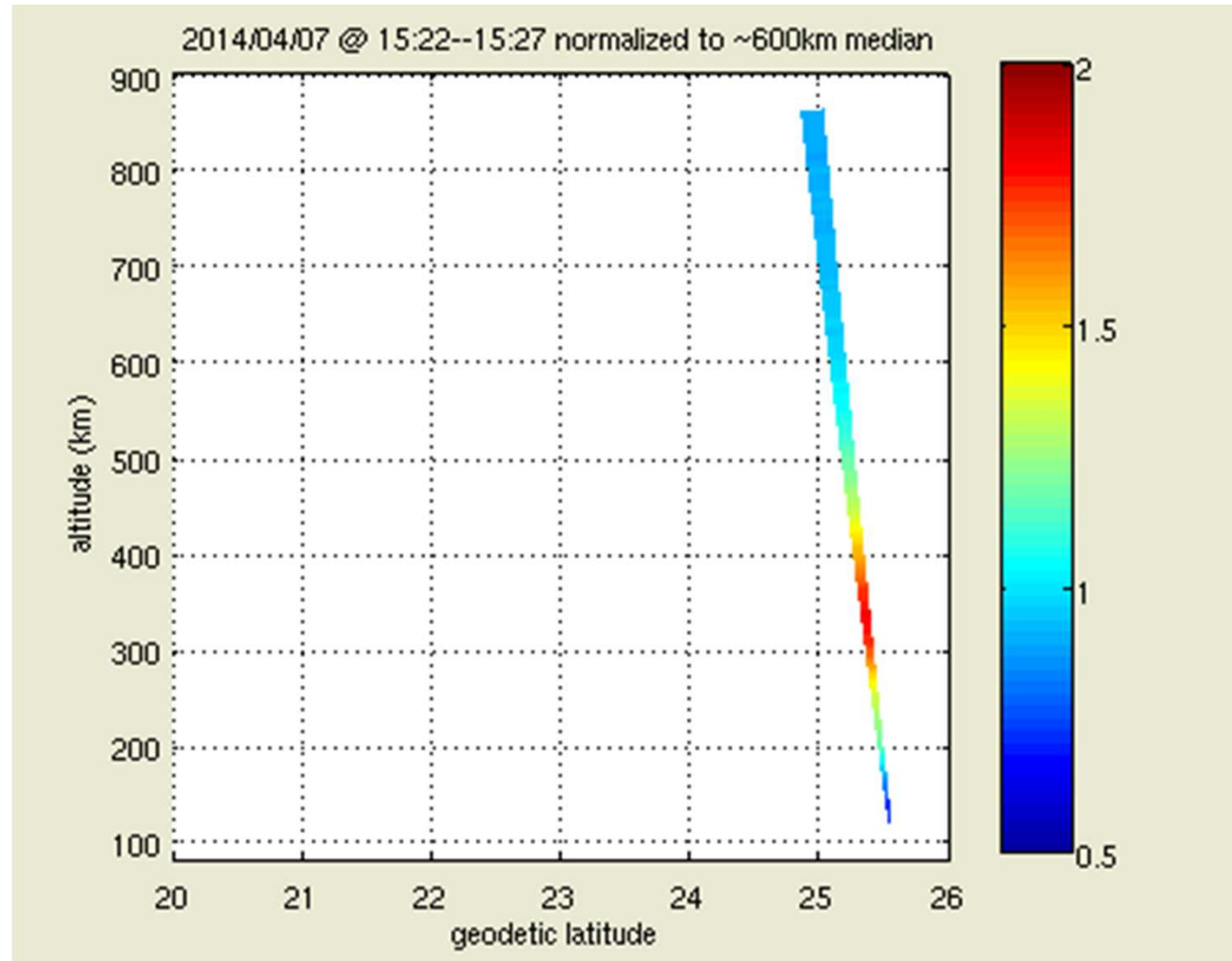


HARVARD-SMITHSONIAN
OBSERVATORY



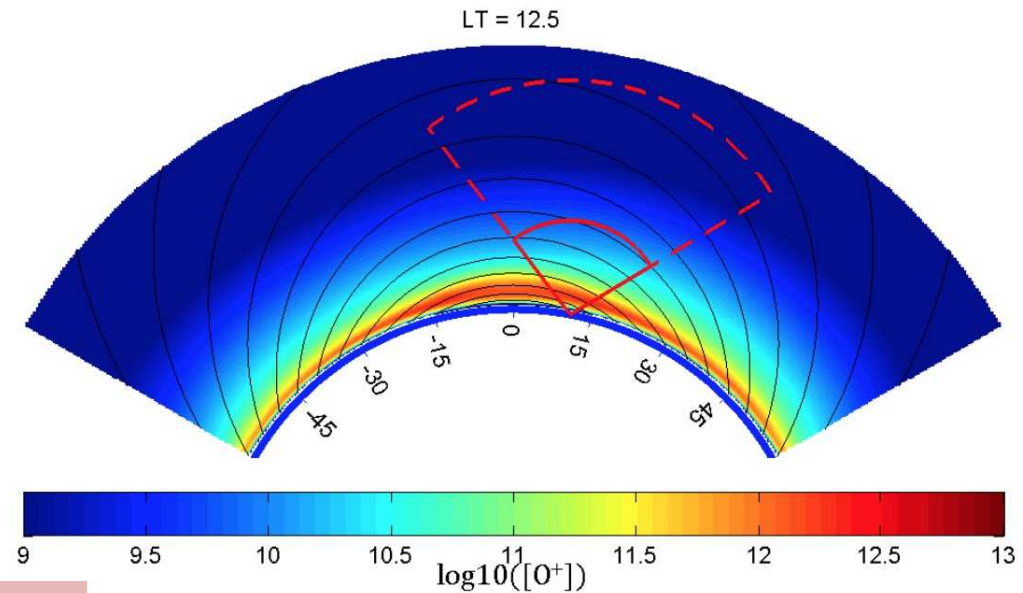
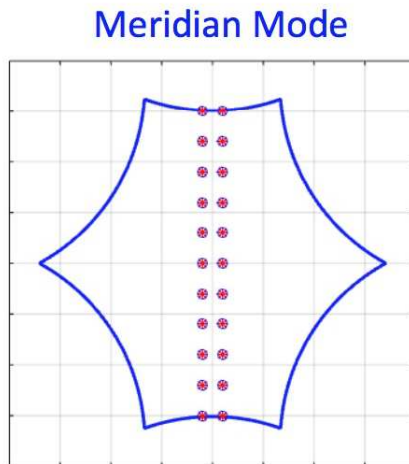
Massachusetts
Institute of
Technology

Qujing ISR elevation scans



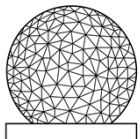
Sanya ISR, under construction

- ◆ Scan zenith: $\pm 48^\circ$
- ◆ Altitude Range: 50-2000 km (possible 5000 km)
- ◆ Observables: $[O^+]$, T_i , V_i



Operational in 2018

Simulated super fountain: O^+ Upflow

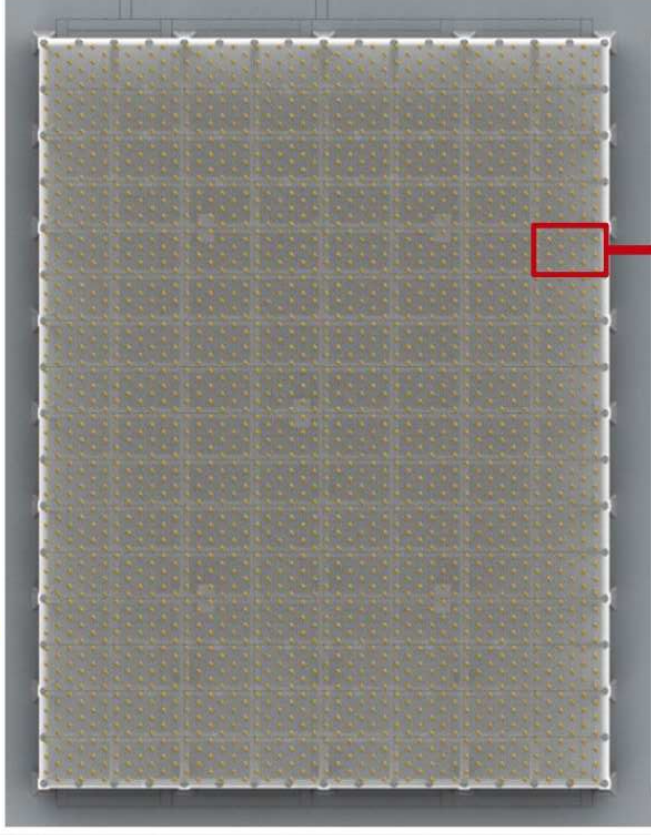


MIT
HAYSTACK
OBSERVATORY

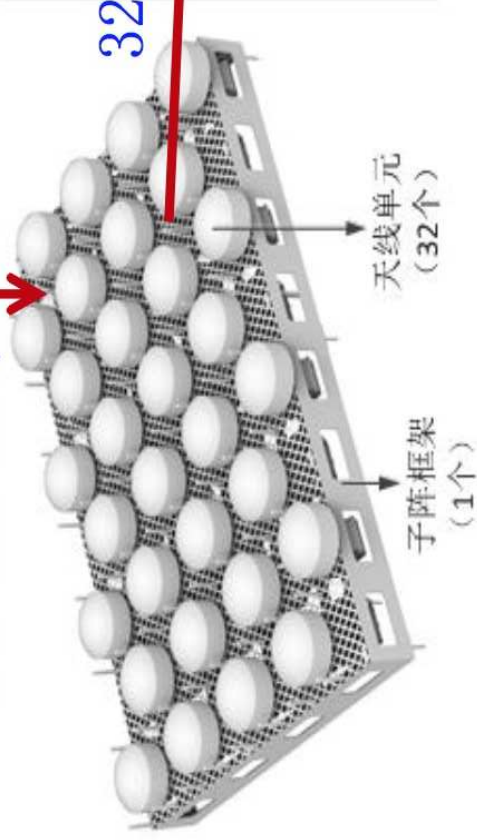
MIT
Massachusetts
Institute of
Technology

SYISR Phase Array

Totally 4096 units, 2 MW

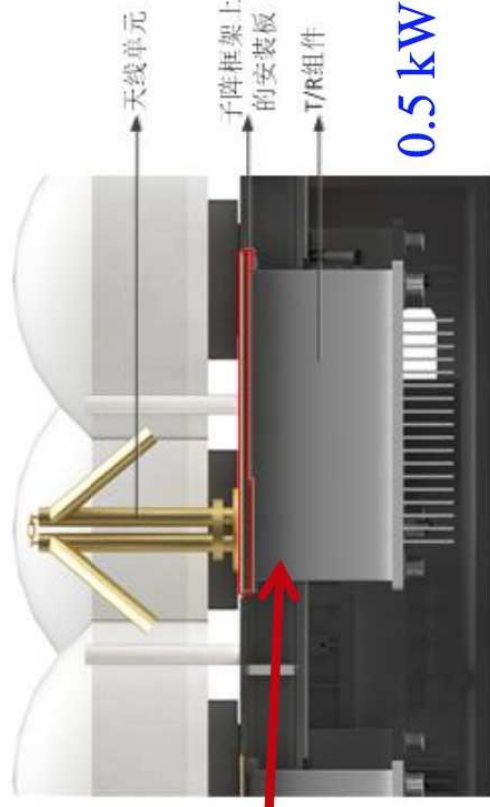


128 sub-arrays

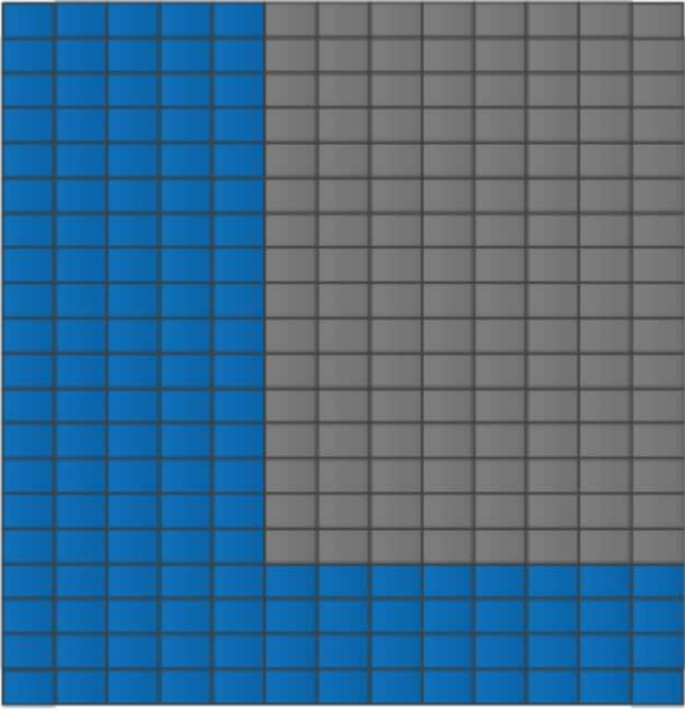


32 units

Frequency	440 MHz
Peak Power	2 MW
Antenna Gain	43 dB
Scan (azimuth)	0 ~ 360°
Scan (zenith)	0 ~ 48°
Max Range	2000 km

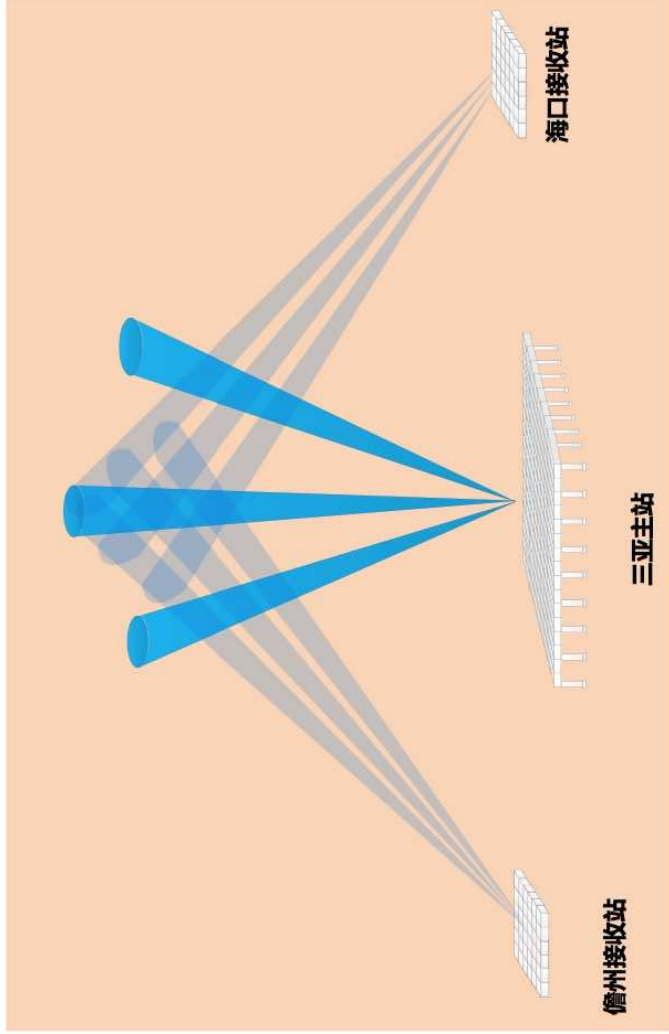


0.5 kW



SYISR Extension in Meridian Project Phase 2 (MP2)

- ◆ Double the phase array
 - 4096 → 8192 units
 - 2 MW → 4MW total power
- ◆ Setup additional 2 receive stations



MP2: Hainan Key Monitoring 子午工程二期之海南低纬重点监测



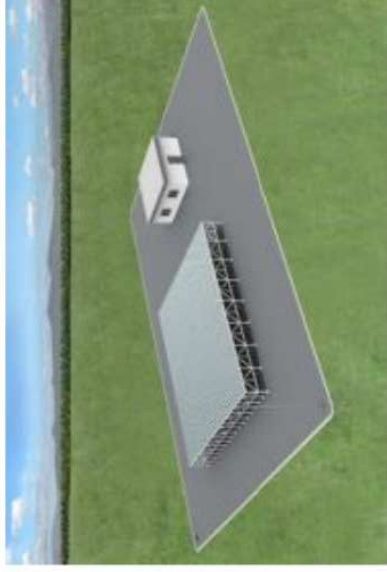
Stations:

- ◆ SanYa (SY)
- ◆ FuKe (FK)
- ◆ HaiKou (HK)

Instruments:

- ◆ ISR (ISR-T/R)
- ◆ HF radar (HF)
- ◆ VHF radar (VHF)
- ◆ Digisonde (DPS)
- ◆ GNSS-TEC (TEC)
- ◆ He Lidar array (HeL)
- ◆ Rocket

Key Instruments



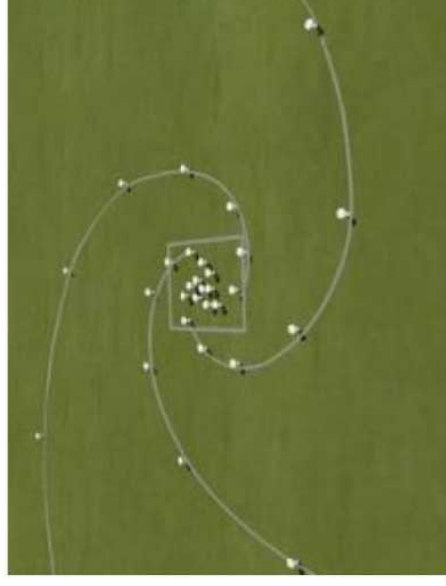
Advanced Modular ISR



Multi-function LIDAR

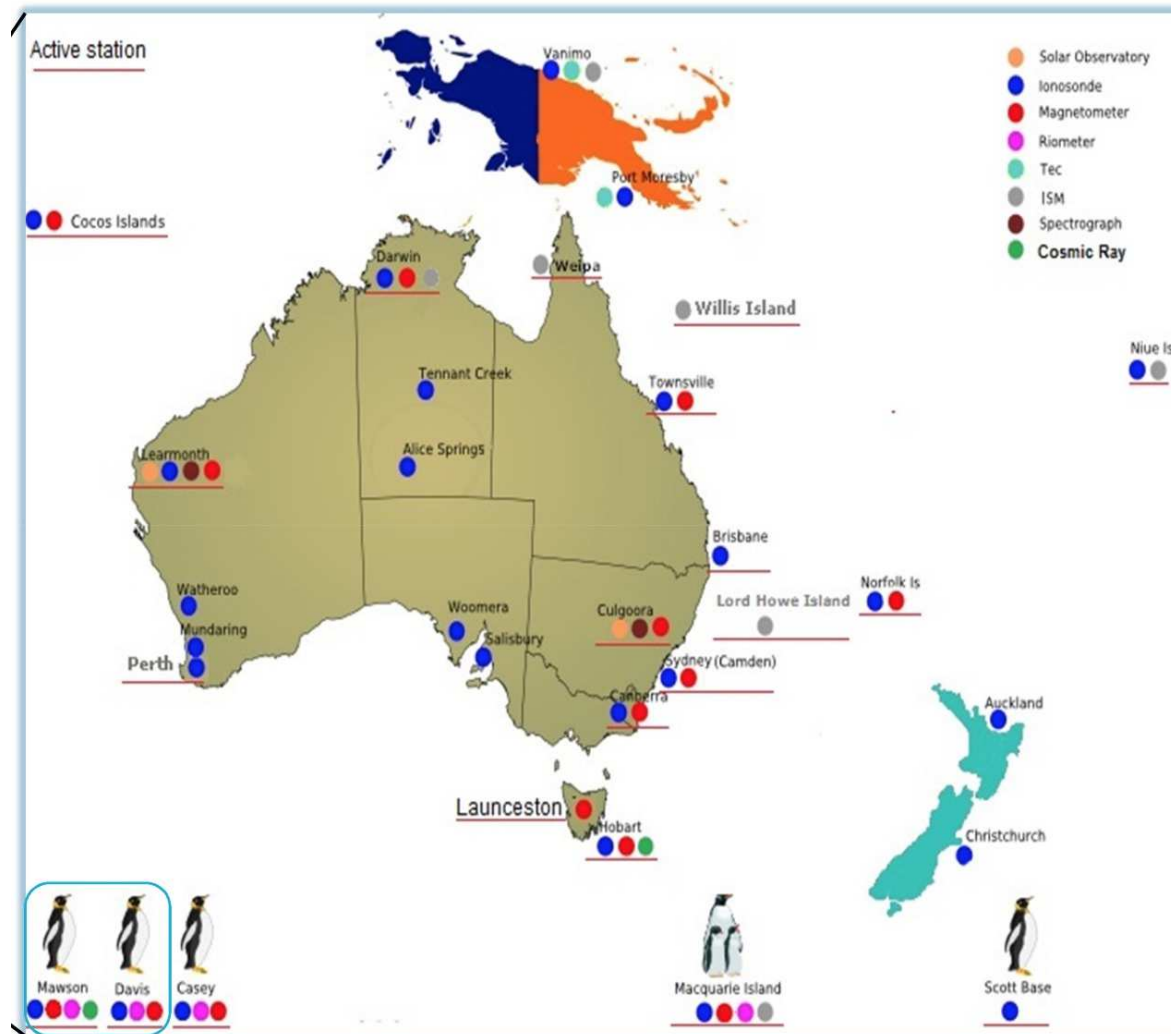


SuperDAWN Radar

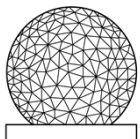


Solar Radio Heliograph

Australia ionospheric stations



Wang (Australia Space Weather Service)



MIT
HAYSTACK
OBSERVATORY



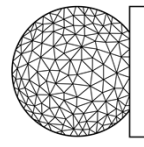


Australia's Space Weather Stations Located in IMCP Belt

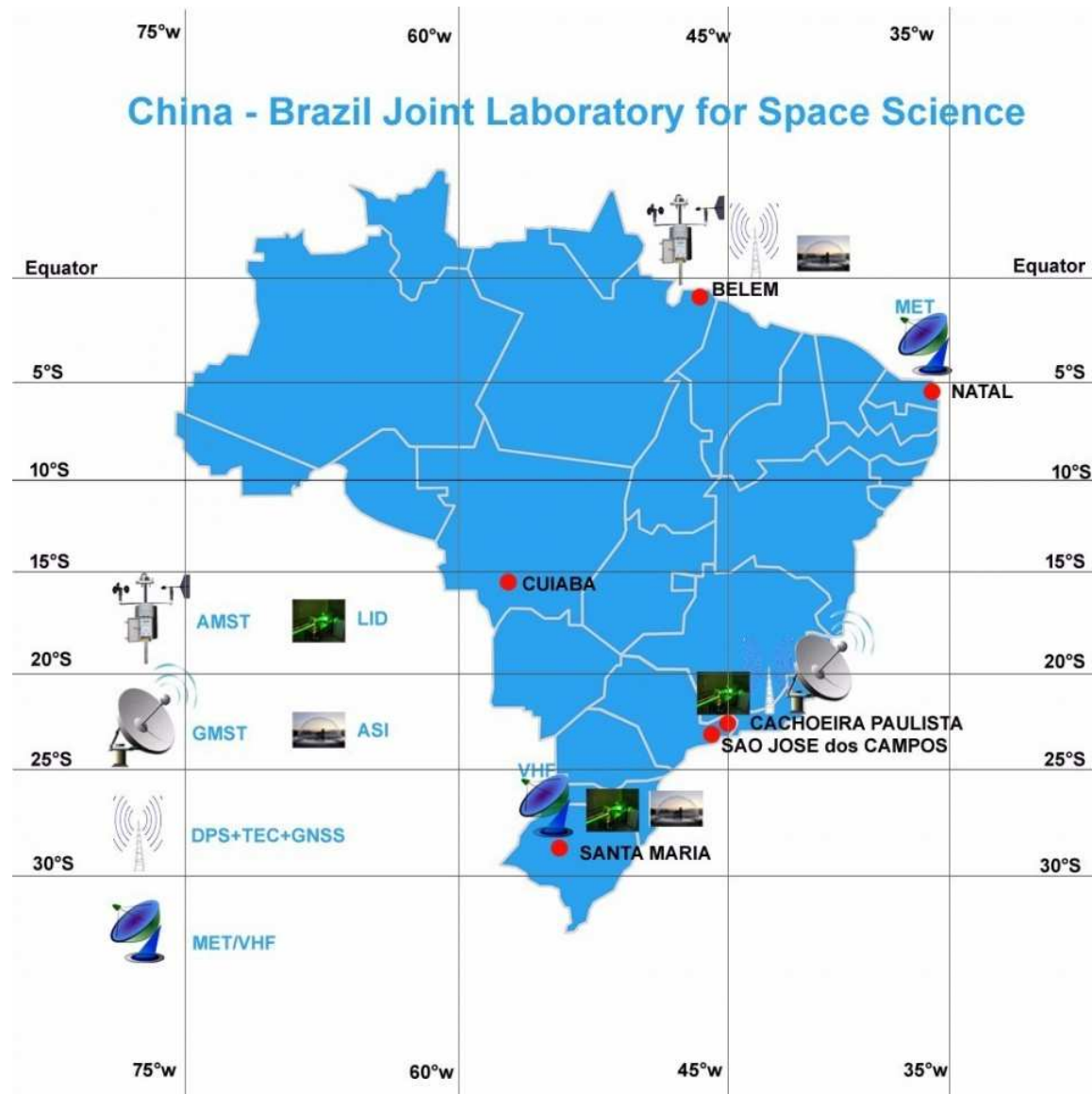
Table 1. Australia's space weather stations list and observations

	Station	Latitude (°)	Longitude (°)	Observation
1	Cocos Islands	12.20 S	96.80 E	Ionosonde, magnetometer
2	Darwin	12.45 S	130.95 E	Ionosonde, magnetometer, ISM*
3	Weipa	12.63 S	141.88 E	ISM*
4	Townsville	19.63 S	146.85 E	Ionosonde, magnetometer
5	Learmonth	22.25 S	114.08 E	Ionosonde, magnetometer, solar, spectrograph
6	Culgoora	30.28 S	149.58 E	Magnetometer, solar, spectrograph
7	Perth	31.94 S	115.95 E	Ionosonde
8	Camden	34.05 S	150.67 E	Ionosonde, magnetometer
9	Canberra	35.32 S	149.00 E	Ionosonde, magnetometer
10	Launceston	41.44 S	147.15 E	Magnetometer
11	Hobart	42.92 S	147.32 E	Ionosonde, magnetometer, cosmic ray
12	Casey	66.30 S	110.50 E	Ionosonde, magnetometer, riometer

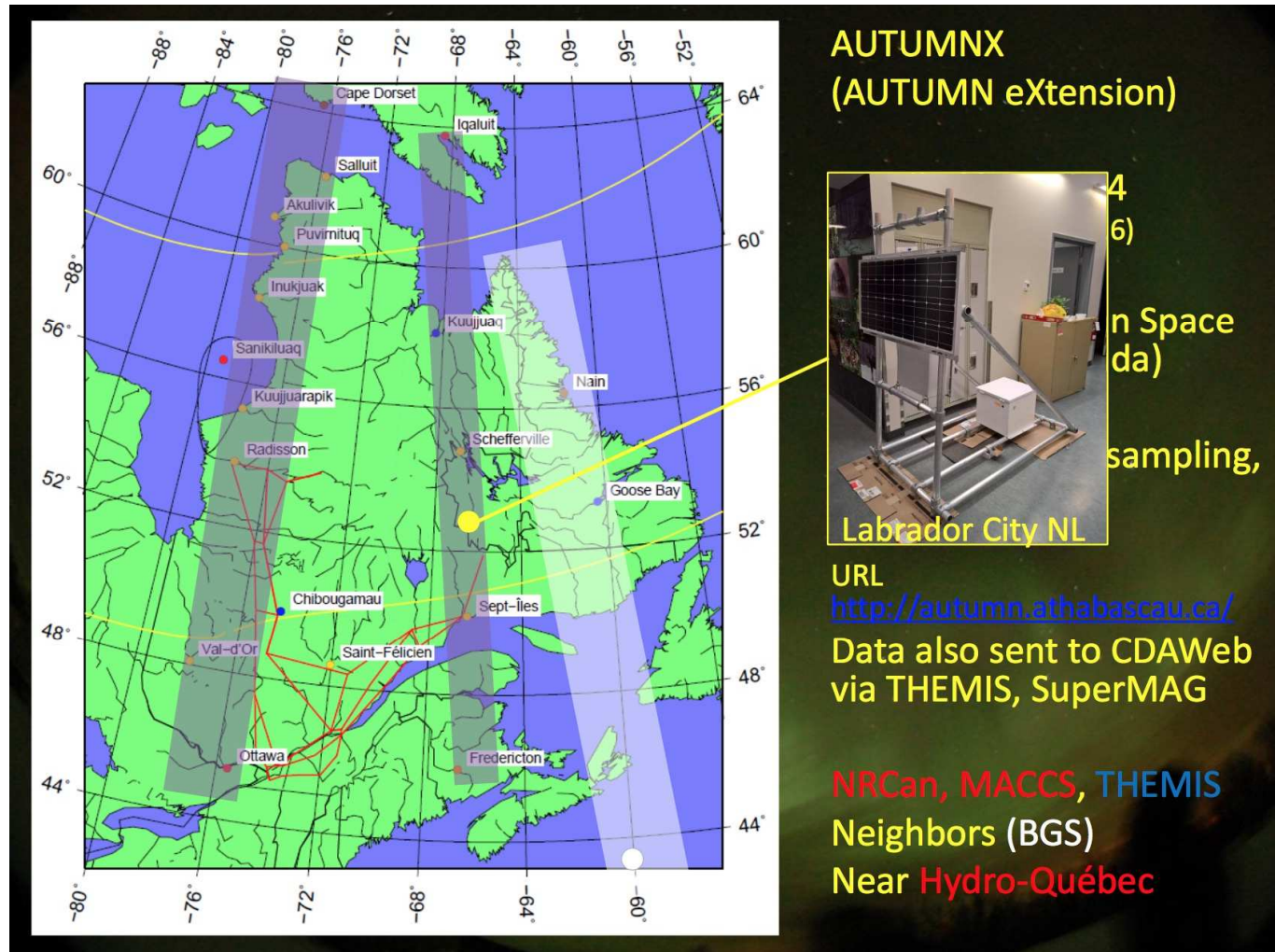
*ISM=ionospheric scintillation monitor



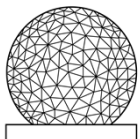
China-Brazil Joint Lab for Space Science



Canadian Magnetometers



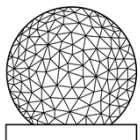
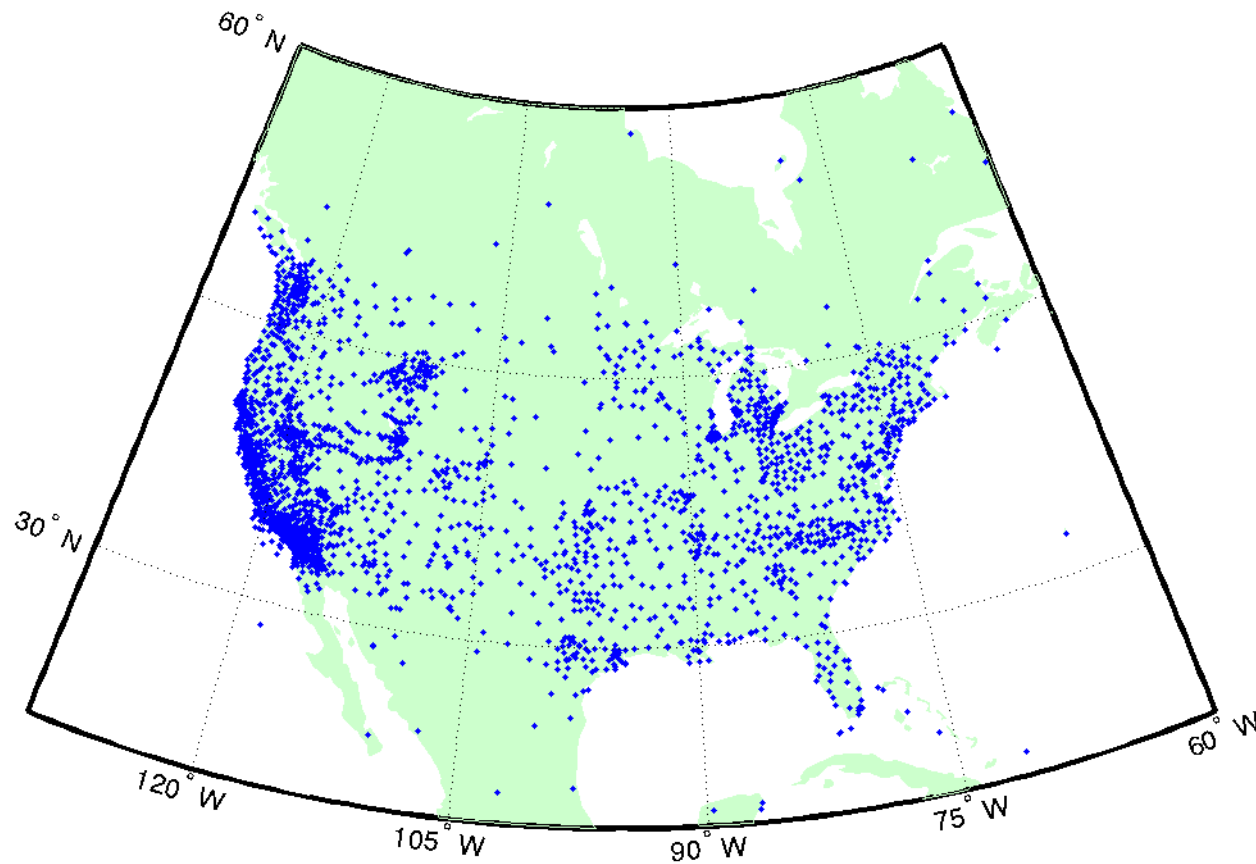
Conners (Athabasca U.)



MIT
HAYSTACK
OBSERVATORY



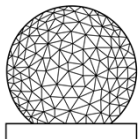
MIT GPS TEC system

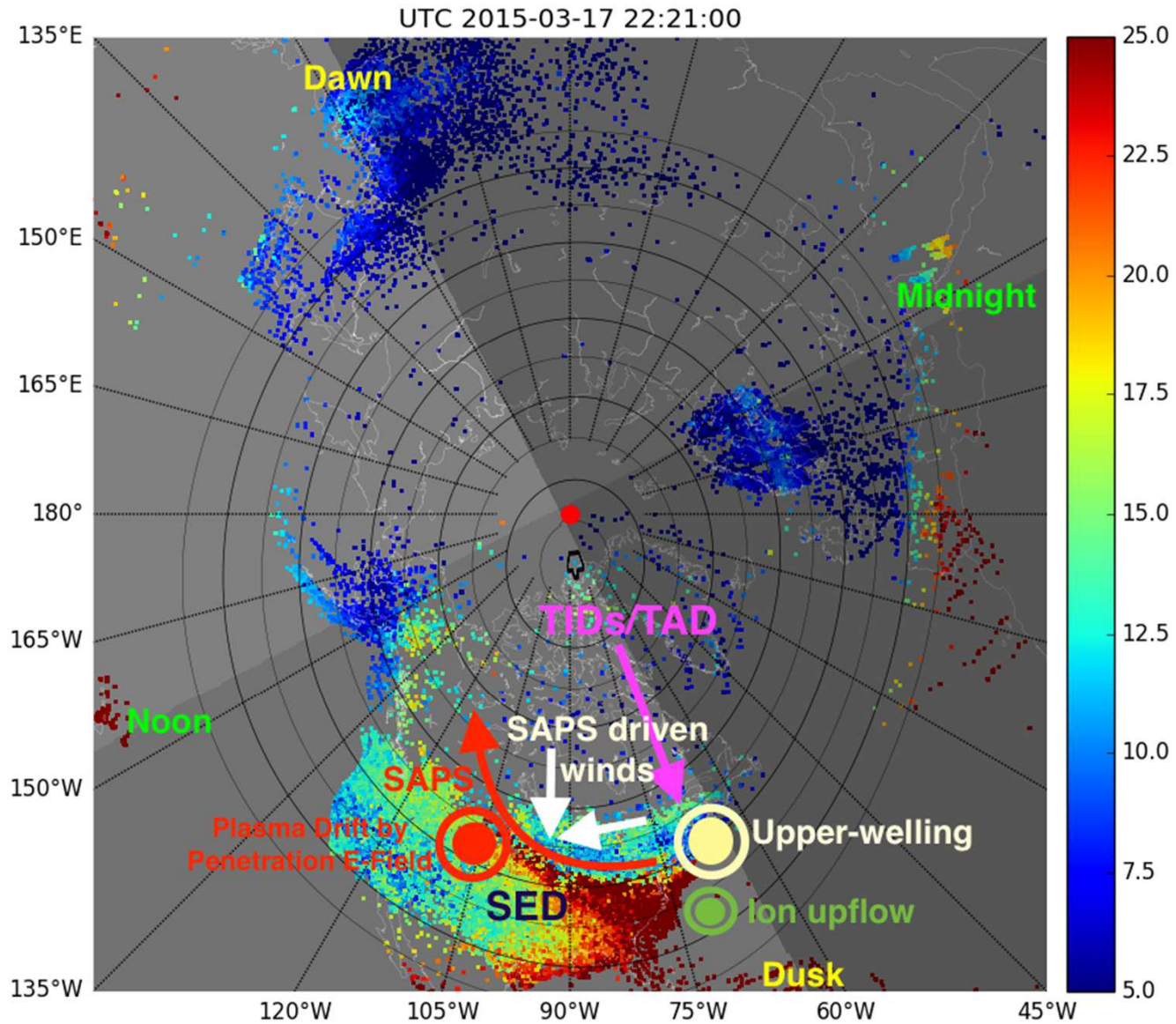


MIT
HAYSTACK
OBSERVATORY

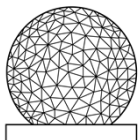
Meridian Circle Campaigns

- 2014
 - March 31 – April 4
 - September 24 – 29
- **2015**
 - **March 17-22, 2015 (St Patrick's Day)**
- 2016
 - March 13 — March 18, 2016
- 2017
 - Sept 13-24, 2017 (window)
- 2018 (June 2018)





Zhang et al. JGR, (2017)



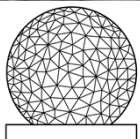
MIT
HAYSTACK
OBSERVATORY



International Meridian Circle Project Workshop, Feb 2011



International Meridian Circle Project Workshop, May 2017



MIT
HAYSTACK
OBSERVATORY



**SA007:
Solar Eclipse Effects on the Upper Atmosphere**

Submit an Abstract to this Session

DEADLINE: today in 10 hours

Session ID#: 26657

Session Description:

For the first time in 26 years, a total solar eclipse will occur in the North American on 21 August 2017. During the eclipse-induced sudden interruption in solar illumination, the upper atmosphere will undergo significant changes beyond what a normal sunset and sunrise process would generate. Although eclipse effects have been studied for many decades, recent major advances in modern observational techniques can provide timely new information on eclipse upper atmospheric system response. Global numerical models have become more capable of capturing important coupling processes on various scales. This session will review existing theories and knowledge of eclipse upper atmospheric effects, examine these against modern eclipse observations, in particular during 21 August 2017, and identify unresolved and challenging problems for future research. We welcome contributions addressing scientific questions of the ionospheric, thermospheric and mesospheric variations during a solar eclipse using ground-based and in situ measurements as well as numerical models.

Please note, the regular AGU abstract submission deadline comes before the 2017 eclipse.

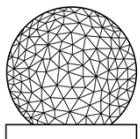
Submissions related to this event must be submitted by the 2 August deadline, however for this session, revision to the submitted abstracts will be possible until September 15, allowing the latest results to be incorporated into the submitted abstracts.

Primary Convener:

Shunrong Zhang, MIT Haystack Observatory, Westford, MA, United States

Conveners:

Larisa P Goncharenko, Massachusetts Institute of Technology, Cambridge, MA, United States
and **Libo Liu**, IGG Institute of Geology and Geophysics, Chinese Academy of Sciences, Beijing, China



**MIT
HAYSTACK
OBSERVATORY**

MIT Massachusetts
Institute of
Technology