

United Nations International Meeting on the  
Applications of Global Navigation Satellite Systems, VIENNA, AUSTRIA, 5 - 9 DECEMBER 2022

# IONOSPHERIC TOTAL ELECTRON CONTENT ABOVE ECUADOR

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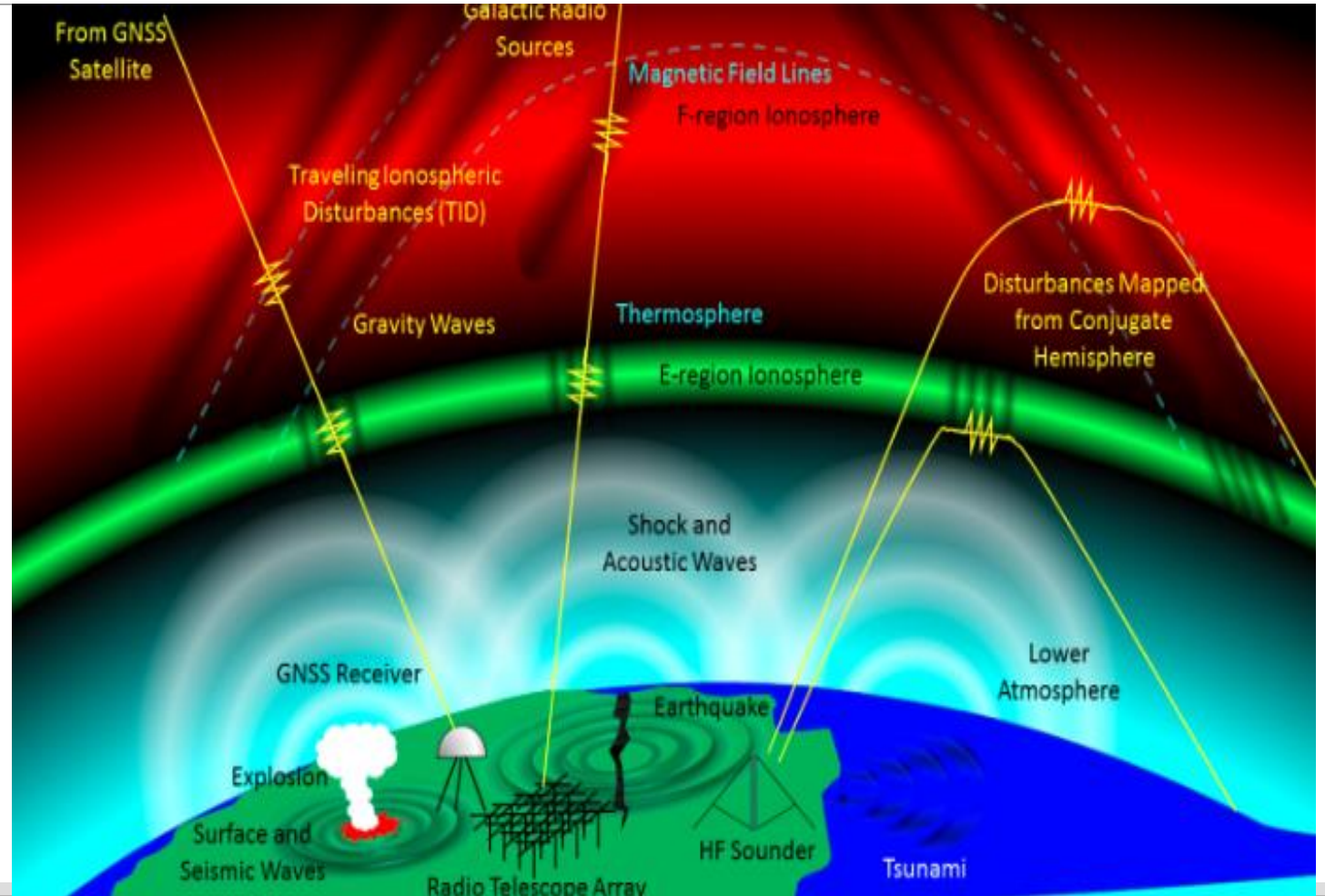
QUITO ASTRONOMICAL OBSERVATORY  
ECUADOR



# INTRODUCTION

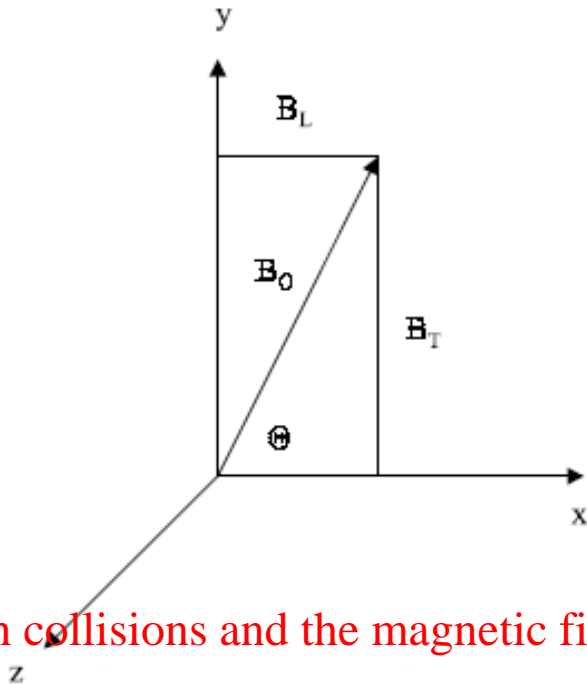
The ionosphere is defined as the region of the upper atmosphere where radio signal propagation is affected by charged particles.

The ionosphere is highly variable in space and time



# IONOSPHERE: DISPERSIVE MEDIUM PROPERTIES

- The ionosphere acts as a refractive medium to radio signals
- The index of refraction depends on the amount of ionisation.
- The Appleton-Hartree equation (see e.g., Komjathy [1997]):



where:

$$n^2 = 1 - \frac{X}{(1 - jZ) - \left[ \frac{Y_T^2}{2(1 - X - jZ)} \right] \pm \left[ \frac{Y_T^4}{4(1 - X - jZ)^2} + Y_L^2 \right]^{1/2}}$$

$$X = \frac{\omega_N^2}{\omega^2} = \frac{f_N^2}{f^2}$$

$$Y = \frac{\omega_H}{\omega} = \frac{f_N}{f}$$

$$Y_L = \frac{\omega_L}{\omega}, \quad Y_T = \frac{\omega_T}{\omega}, \quad \text{and} \quad Z = \frac{\omega_C}{\omega}$$

When both collisions and the magnetic field are negligible

$$n = 1 - \frac{1}{2}X = 1 - \frac{Ne^2}{2\epsilon_0 m \omega^2}$$

The phase refractive index appropriate for the carrier phase observations,

$$n_{\phi} \cong 1 - \frac{40.3 \cdot N}{f^2},$$

The group refractive index appropriate for the pseudorange observations (Langley [1996]):

$$n_g \cong 1 + \frac{40.3 \cdot N}{f^2}.$$

**THE IONOSPHERIC DELAY IS DEFINED AS:**

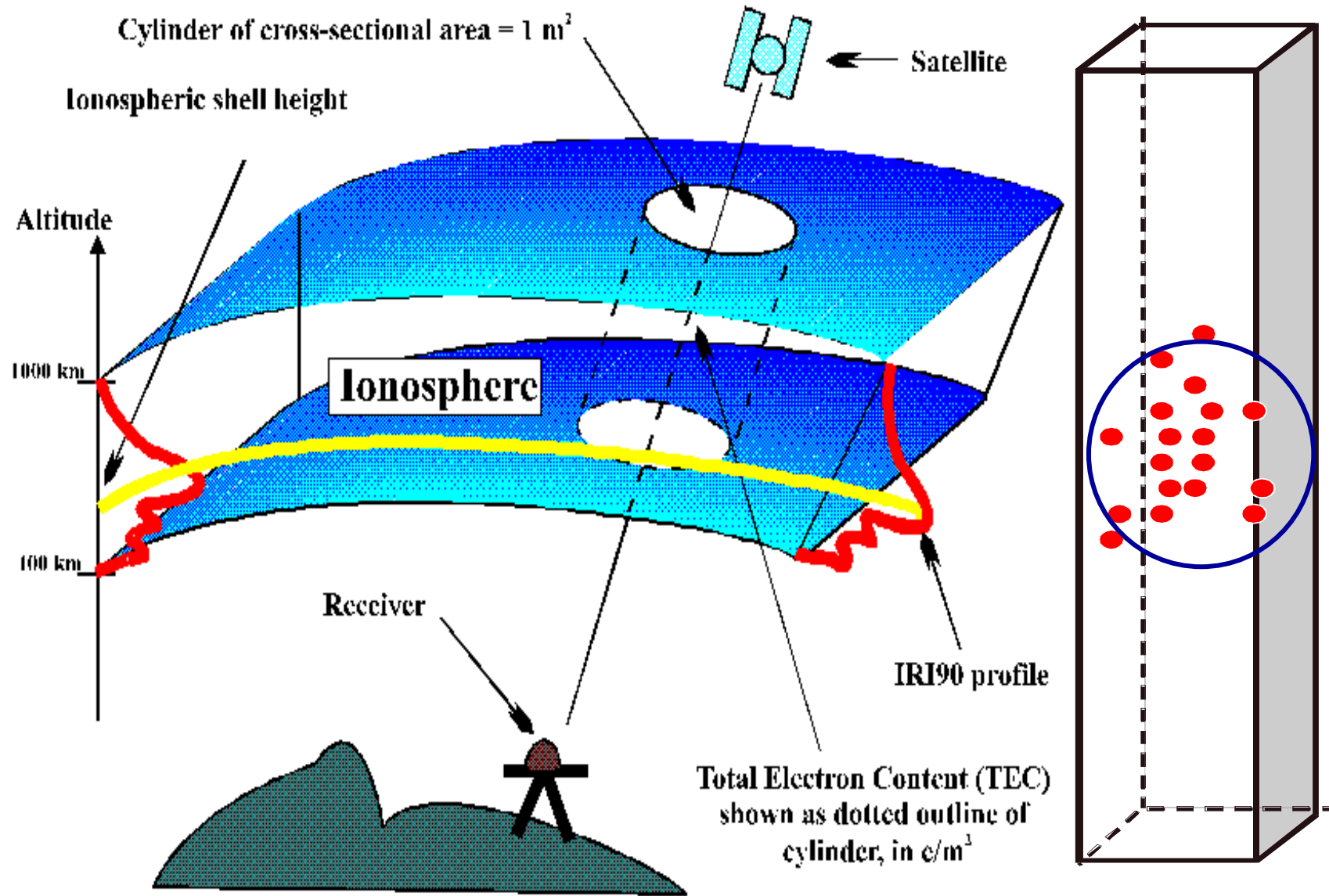
$$d_{ion} = \int_{Sat}^{Rec} (n - 1) ds,$$

The first order approximation of the Appleton-Hartree formula also gives an expression for the ionospheric delay in terms of TEC:

$$d_{ion} = \frac{40.3}{f^2} TEC$$

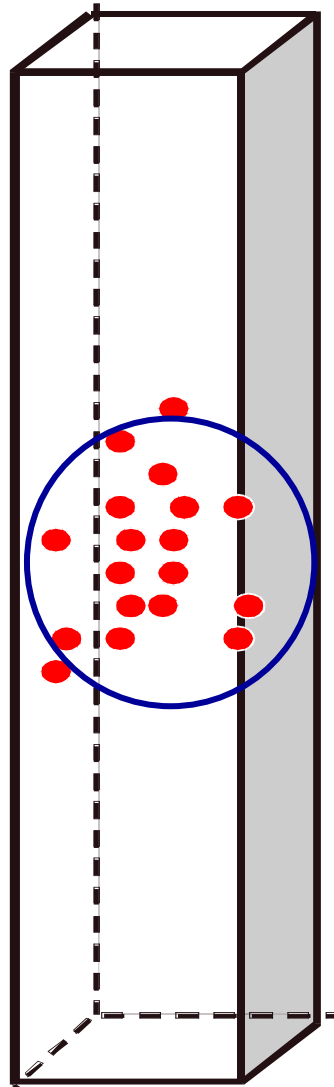
gives cm level accuracy for the ionospheric delay.

where TEC is the total electron content in TEC units (1 TECU =  $10^{16}$  1/m<sup>2</sup>).

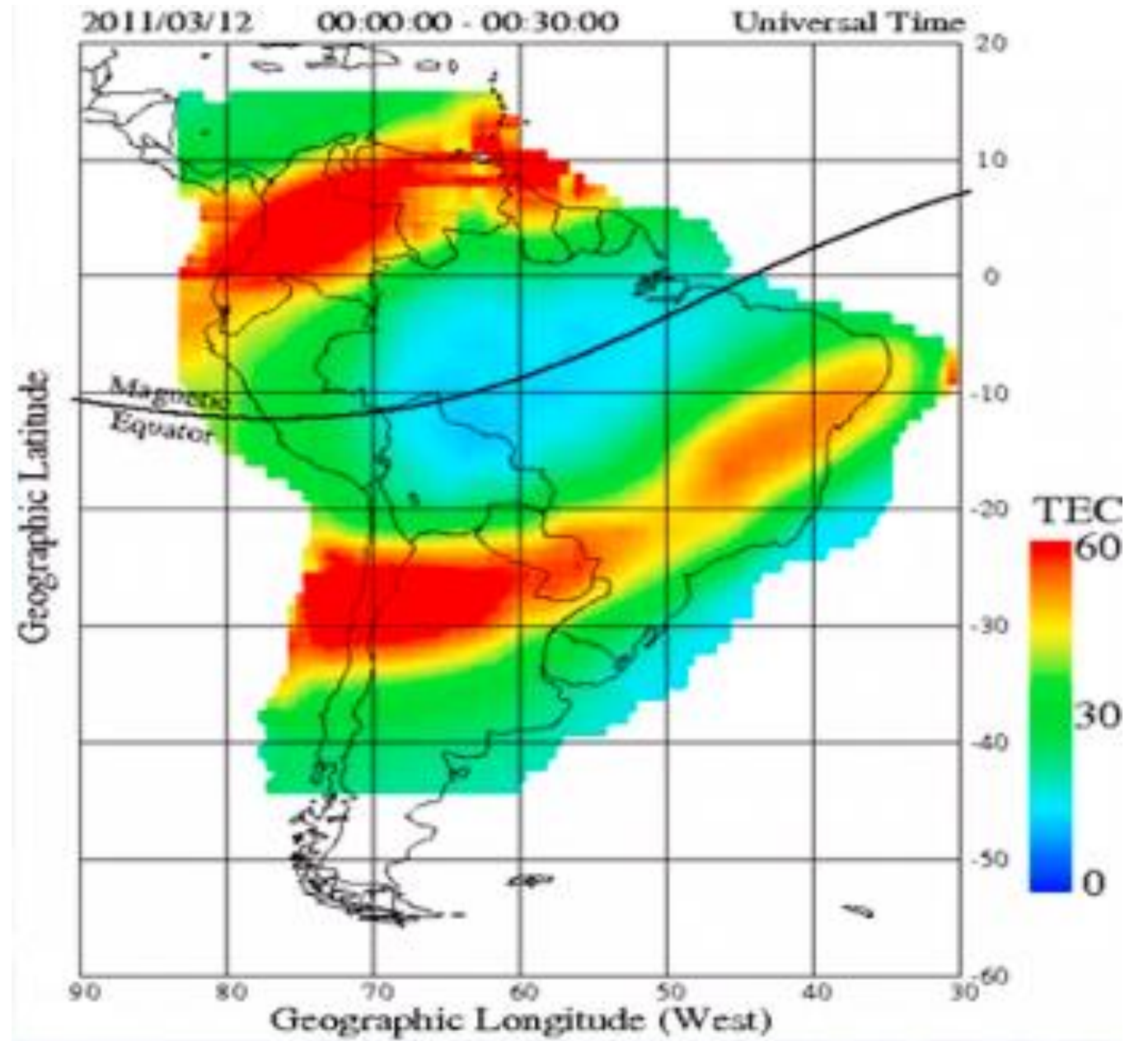
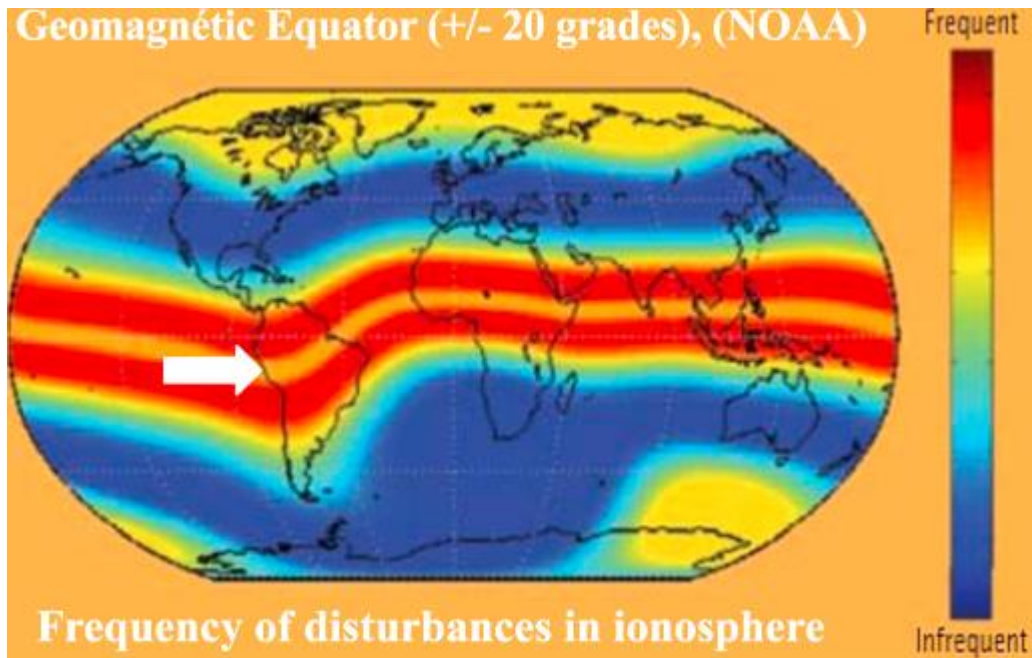


# TOTAL ELECTRON CONTENT (TEC)

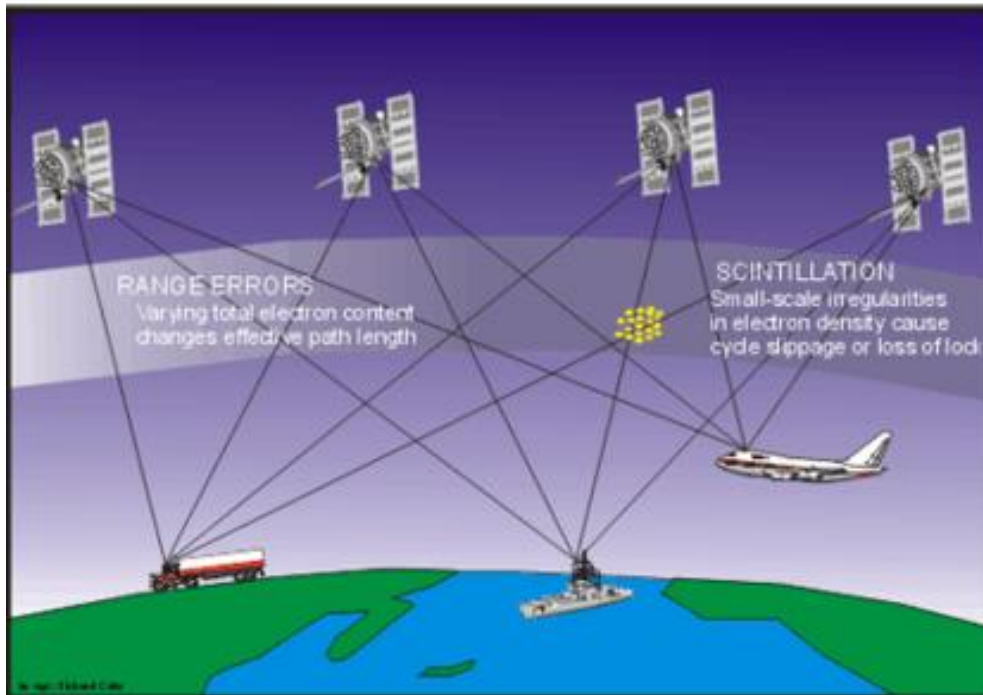
Total number of electrons in a column with cross section of  $1 \text{ m}^2$   
 (1 TECU =  $10^{16}$  electrons/ $\text{m}^2$ )



# TOTAL ELECTRON CONTENT



# SCINTILLATION



0.19 TEC causes problems for the GPS receiver.

Principal impacts of ionospheric scintillation on GPS performance:

- Loss of lock / outages
- Induced ranging errors

Consequences of these effects on GPS positioning accuracy depends on constellation geometry

For example, losing multiple satellites in the same region of the sky can lead to large errors

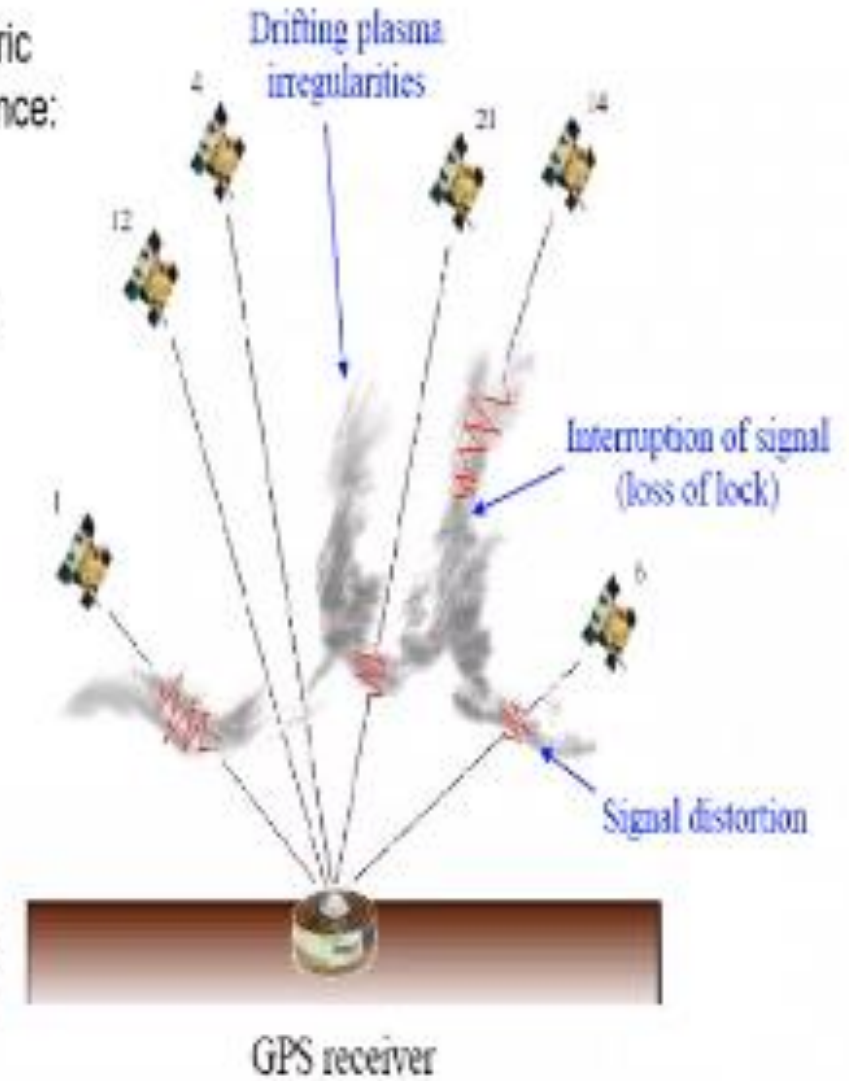
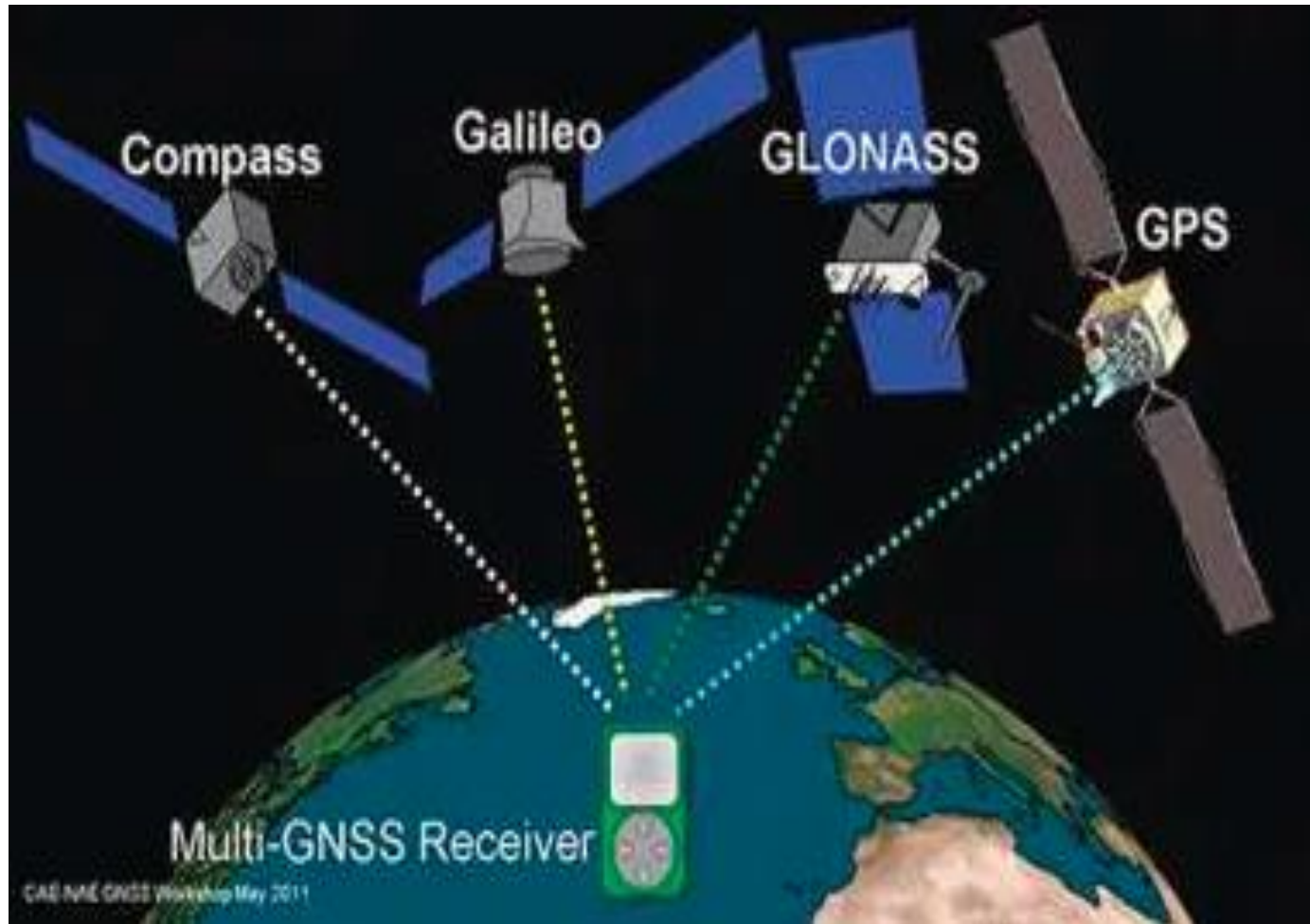


Figure Courtesy of C. Carrano, BC



# ONE TOOL USED TO MEASURE TEC IS THE GLOBAL NAVIGATION SATELLITE SYSTEM (GNSS).



Modelling TEC gives scientists the opportunity to study the temporal and spatial variation of the global ionosphere

ionospheric corrections for navigation, surveillance and communication systems

# IONOSPHERIC MODELING USING GPS DATA

The Total Electron Content (TEC) is the amount of free electrons along the path of the electromagnetic wave between each satellite and the receiver.

TEC is calculated by:

$$TEC = \int_A^B N(s) ds,$$

where  $N$  is the electron density in ( $e/m^3$ ),  
 $ds$  is the path integral differential a  
the integral is calculated along the path that joins points A and B.

It is an important geophysical parameter:

$$TEC = \int_{\text{receiver}}^{\text{satellite}} N \cdot ds$$

\* The TEC also has been measured for decades using the Faraday Rotation effect on a linear polarized propagating plane wave (Klobuchar, 1985 and 1996).

# IONOSPHERIC OBSERVABLE

The GPS pseudorange ( $P$ ) and carrier phase ( $\Phi$ ), observable parameters, are given, for the two GPS frequencies:  $L1 = 1575.42$  MHz and  $L2 = 1227.6$  MHz.

The GPS observation equations are well known from the various references such as Langley [1992; 1996], Wells et al. [1987], Chao et al. [1996], etc

$$P_1 = \rho + c(dT - dt) + d_{ion,L_1} + d_{trop} + \eta_{P_1} + \epsilon_{P_1}$$

$$P_2 = \rho + c(dT - dt) + d_{ion,L_2} + d_{trop} + \eta_{P_2} + \epsilon_{P_2}$$

$$\Phi_1 \lambda_1 = \rho + c(dT - dt) + \lambda_1 N_1 - d_{ion,L_1} + d_{trop} + \eta_{\Phi_1} + \epsilon_{\Phi_1}$$

$$\Phi_2 \lambda_2 = \rho + c(dT - dt) + \lambda_2 N_2 - d_{ion,L_2} + d_{trop} + \eta_{\Phi_2} + \epsilon_{\Phi_2}$$

BUT TODAY TEC MEASUREMENTS ARE MADE MOSTLY USING GNSS DATA

Pseudorange and raw carrier phases that can be measured with high precision.

$$TEC_P = 9.52 \cdot (P_2 - P_1)$$

unambiguous but noisy and therefore imprecise observable

Modern receivers are typically at the 0.9 mm level

$$TEC_\Phi = 9.52 \cdot (\Phi_1 \lambda_1 - \Phi_2 \lambda_2).$$

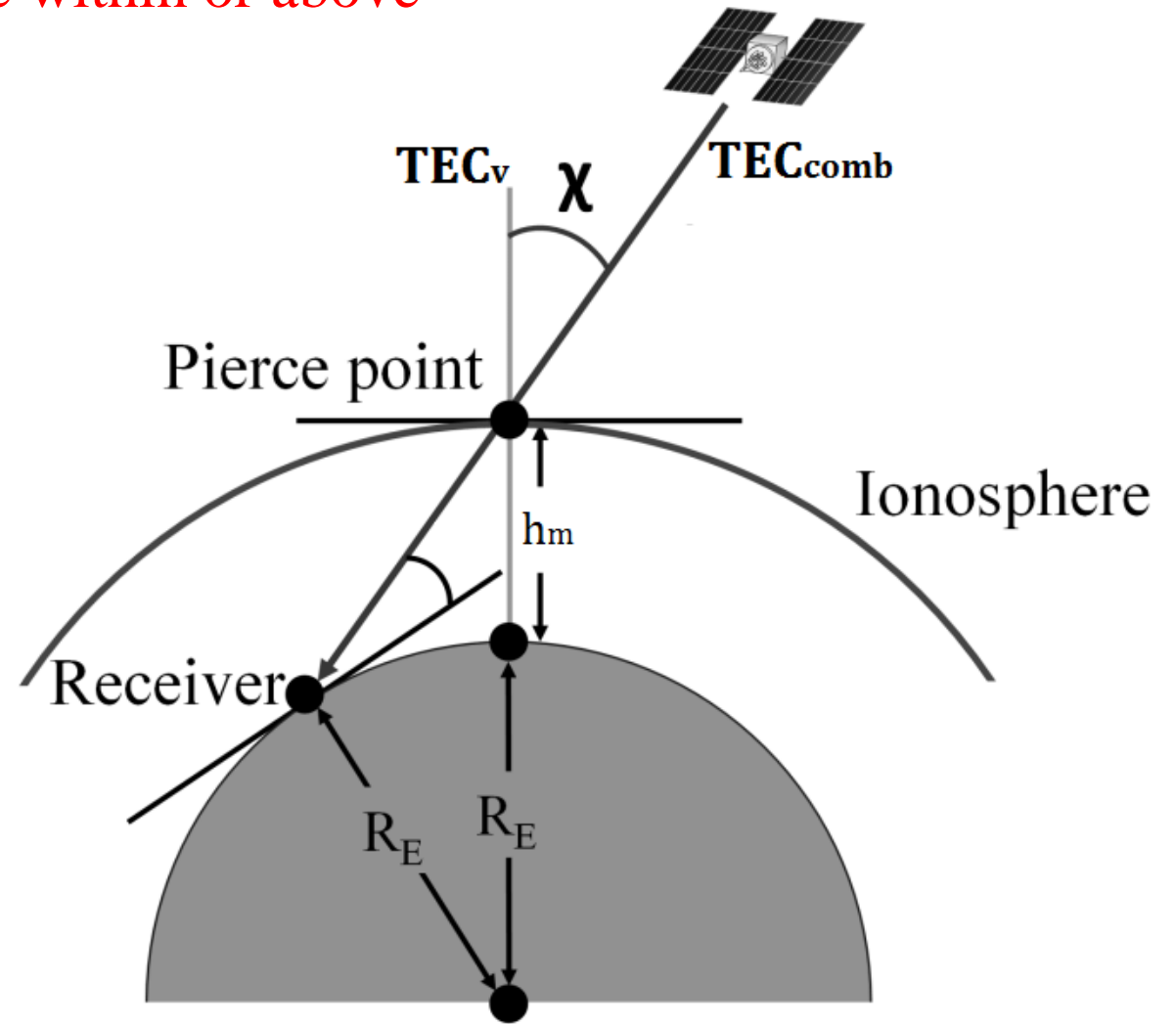
very precise but ambiguous observable

Satellite-based techniques available to measure TEC up to the altitude of the satellite which can be anywhere within or above the ionosphere.

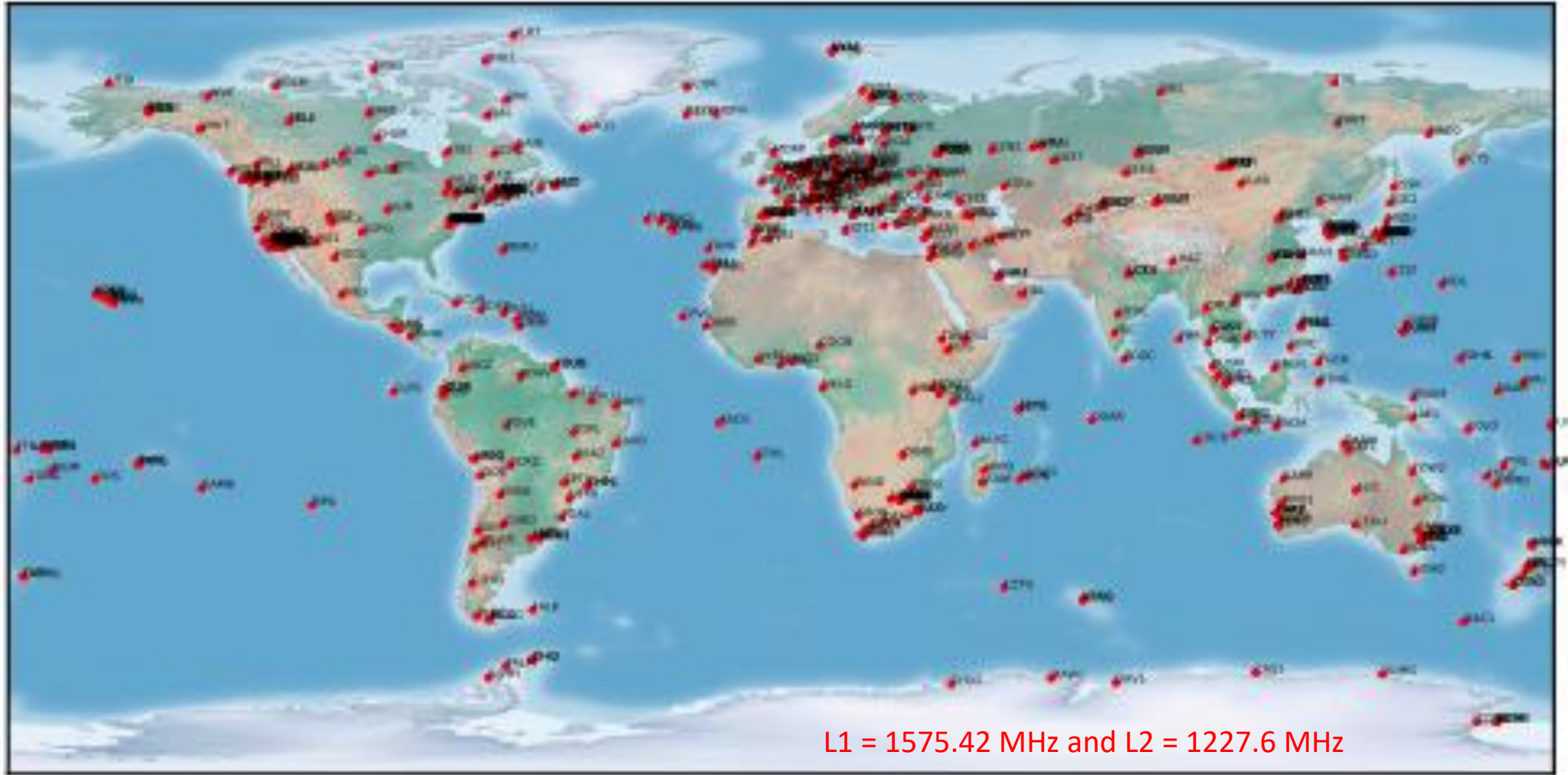
$$TEC_{comb_i} = TEC_{\Phi_i} - \frac{\sum_{j=1}^n p_j [TEC_{\Phi_j} - TEC_{P_j}]}{\sum_{j=1}^n p_j}$$

$$TEC_v = TEC_{comb} (\cos \chi')$$

$$\sin \chi' = \frac{R_e}{R_e + h_m} \sin(z)$$

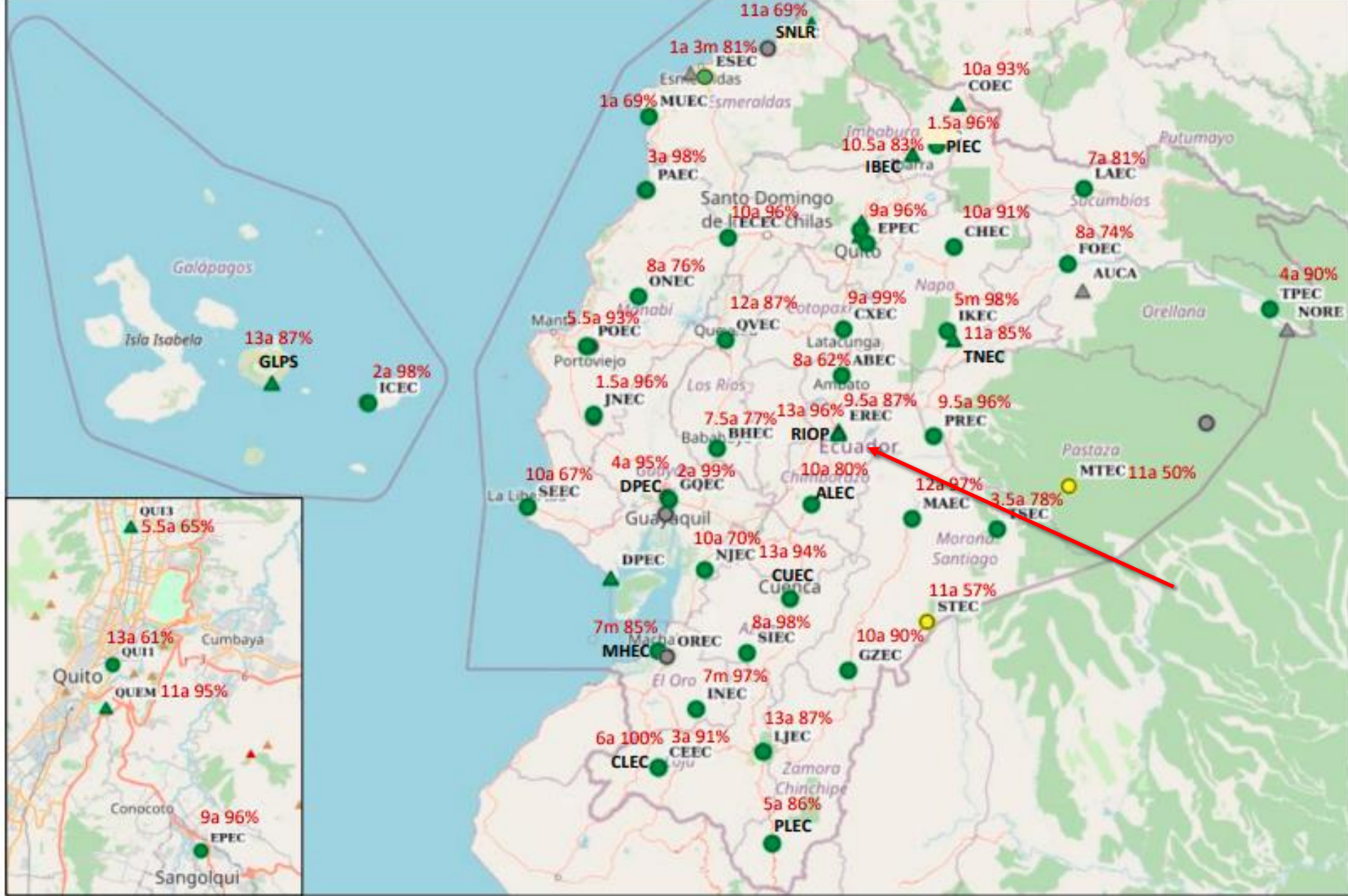


SPACE WEATHER EFFECTS ON GPS



L1 = 1575.42 MHz and L2 = 1227.6 MHz

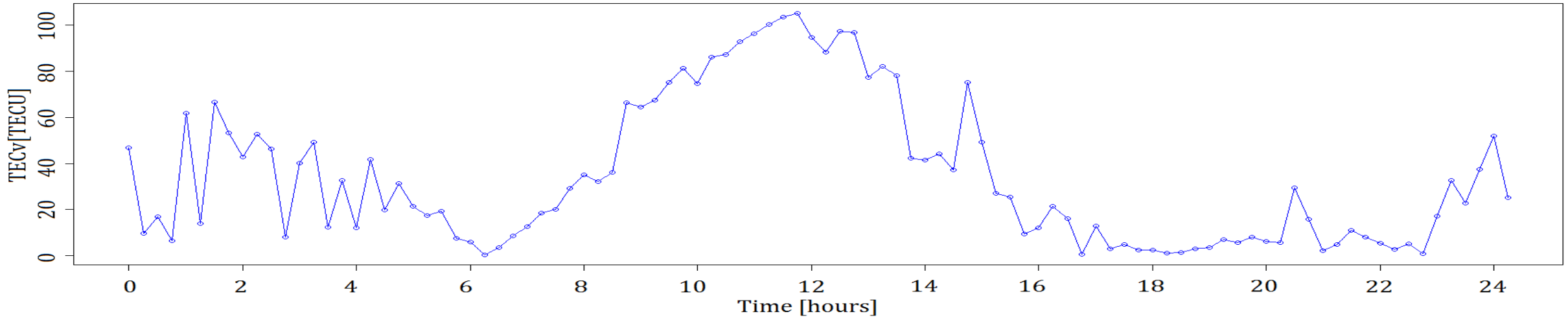
# GPS stations in Ecuador



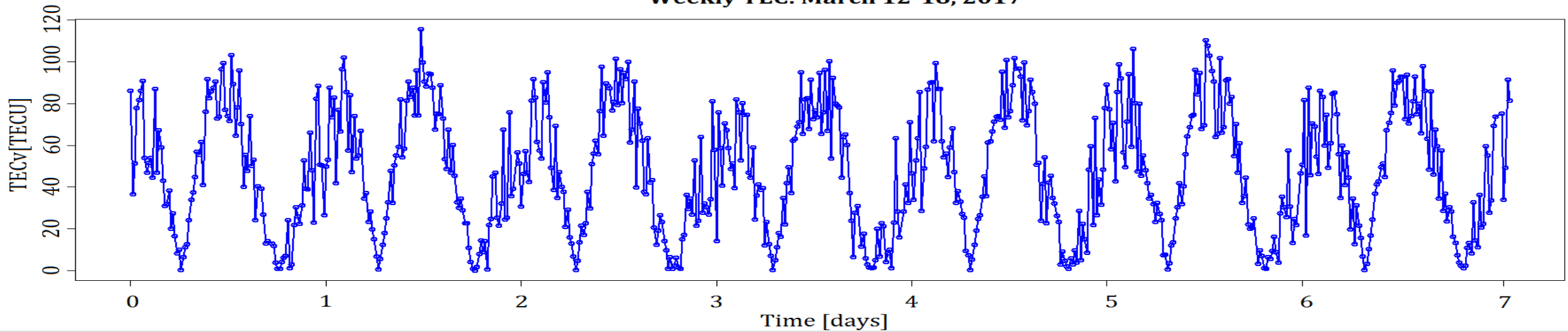
# ESTIMATION OF THE TEC OVER ECUADOR

Riobamba RIOP station:  
(latitude -1.65 longitude -78.65)

**Daily TEC: September 18, 2011**

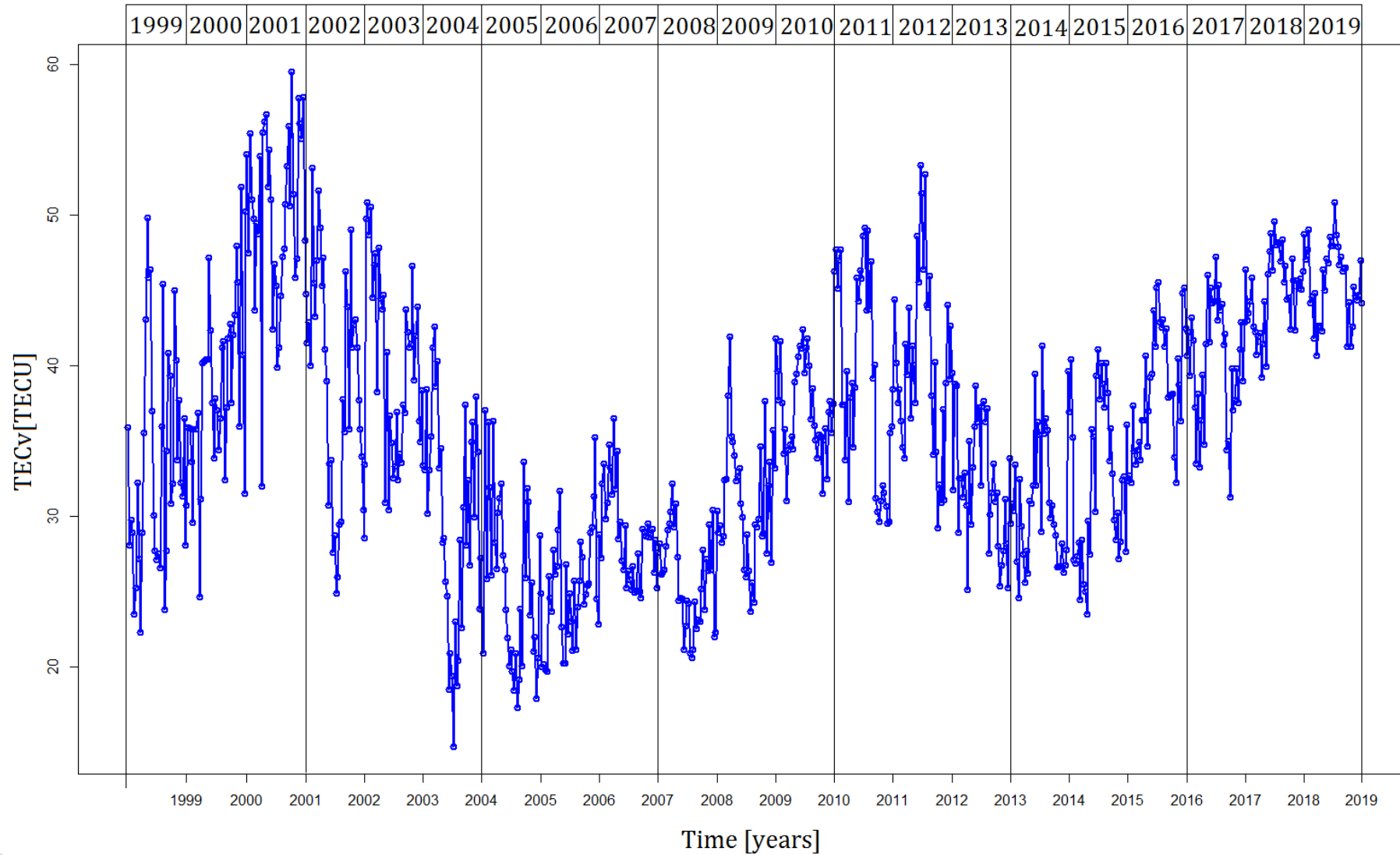


**Weekly TEC: March 12-18, 2017**



# The cumulative average TEC for the whole period from 1999 to 2019

Cumulative average TEC: 1999-2019





# TEC calculated vs. observed

The chi square test:

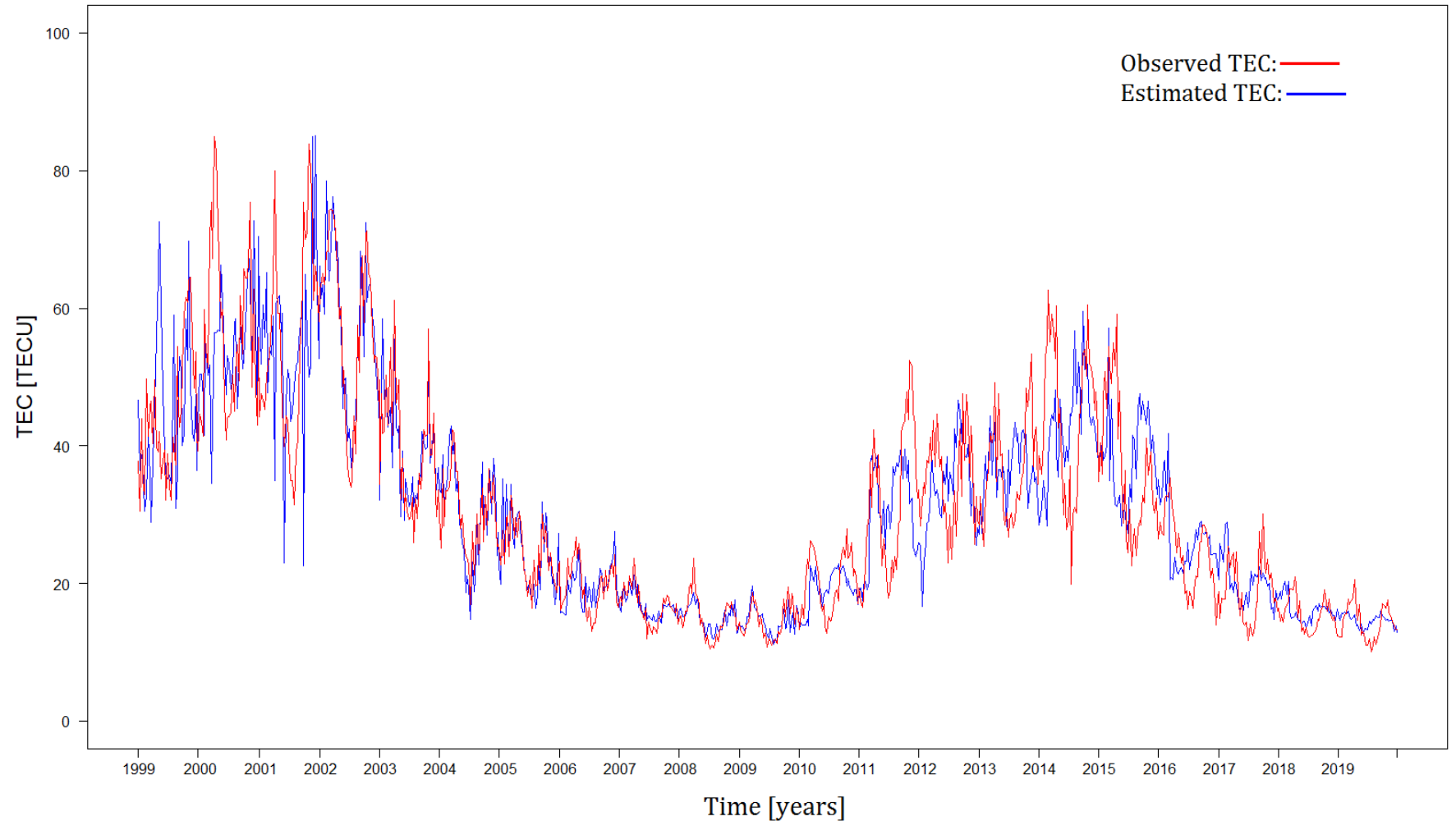
$$\chi^2 = \sum \frac{(O_i - C_i)^2}{C_i}$$

$O_i$  observed values

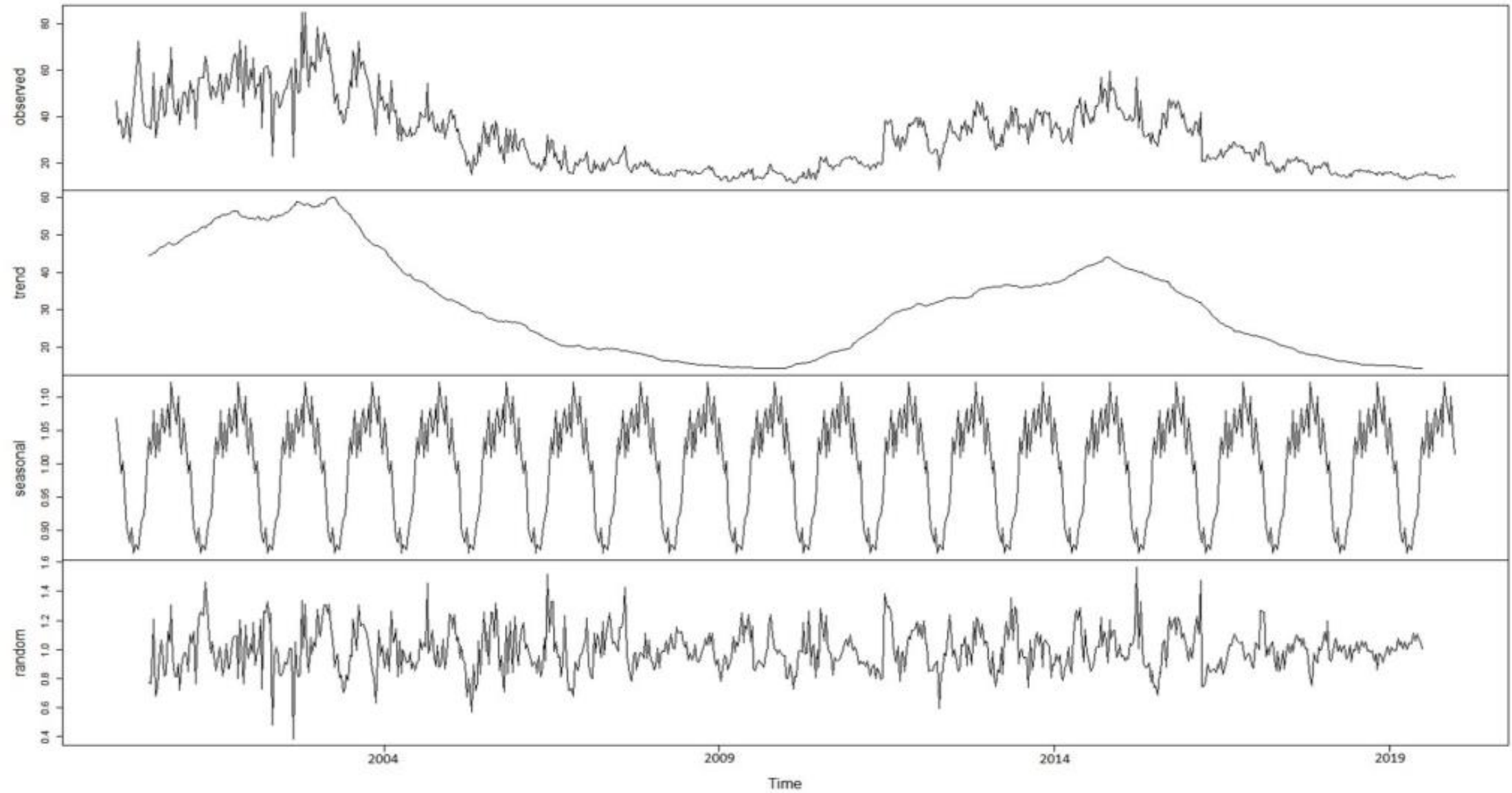
$C_i$  calculated values

goodness of the fit  
about 0.88

Observed and estimated TEC:1999-2019



### Decomposition of multiplicative time series



Time series analysis of the annual TEC from 1999 to 2019 period.

Thank  
you!