Space Weather Observatory: Facilities for the Space Navigation



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Summary

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PRESENTATION

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FACILITIES FOR THE SPACE NAVIGATION



CONCLUSIONS



1. PRESENTATION











NATIONAL CENTER OF SATELLITE IMAGES OPERATIONS - CNOIS



LAUNCH VEHICLES



GEOMATICS



SCIENTIFIC INSTRUMENTATION



space weather observatory: facilities for the space navigation **1. PRESENTATION**





2. SPACE WEATHER OBSERVATORY

Facilities implemented by CONIDA to perform scientific research in Space Weather and space navigation systems. Different equipments have been installed since 2006 in order to monitor long-term and transient events from geospace anomalies and solar-terrestrial connections. CONIDA is operating a telescope (observation of artificial satellite orbits, DEBRIS) as part of projects with APSCO, and TEC monitoring instrument.



SAVNET

VLF Ionospheric Observations for Solar Flares



CALLISTO

Solar Flares Observations



LAGO Cosmic Rays Observations



SOLAR MONITORING

Solar Total Radiation, Hα Filter UV Index Sunspots Observation and Prediction Astronomical Ephemerides

CRIRP

Ionospheric ,models and GNSS data analysis

space weather observatory: facilities for the space navigation 2. SPACE WEATHER OBSERVATORY



APOSOS Debris, Satellites and Asteroids Observations



SAVNET CRAAM/EE



NAA

SM3

EACE

NIK

NPM

ND



- 8 VLF tracking receiver stations deployed in Brazil, Peru (CONIDA, UP, UNICA) and Argentina.
- 10 years of operation since 2007
- Long-term and transient solar activity (Ly- α ; solar flares)
- Mesospheric disturbances (T, NO, O3)
- Physics of the lower ionospheric (C/D) regions
- Atmos. Physics (TGFs)
- Subionospheric radio propagation modeling
- Search for seismic-EM effects
- Detection of Remote astrophysical objects

Educational, skills, human resources







SPACE WEATHER OBSERVATORY - OBJETIVE

Monitor long-term and transient events from geospace anomalies and solar-terrestial connections.



http://www.conida.gob.pe/index.php/SAVNET/savnet.html





The Large Aperture Gamma ray bursts Observatory (LAGO)

A very long baseline "array" of water Cherenkov detectors (WCD)



LAGO

- Non-centralized, collaborative network of Institutions
- Developments, expertize and data are shared across the network.
- Sites in eight countries: Argentina, Bolivia, Colombia, Ecuador, Guatemala, México, Perú & Venezuela.



• PMT(s) – Digitizer board (own desing) + GPS +

LAGO "typical" Water Cherenkov Detector

- Simple WCD detector desing: sensitivity to charged secondar particles and y (mainly trough y e+ e-) ->
- Commercial water tanks + internal reflective and diffusive coating
- FPGA based electronics: detector control and data adquisitior top sensor.



Scientific goals:

- Search for HE component of GRBs at ground level.
- Study transient and long term Solar modulation (SM) of Cosmic Rays (CR) phenomena.

Academic goals:

- Train latin-american students in H.E. and Astroparticle researchers.
- Build a latin-american grid of Astroparticle researchers.

http://www.conida.gob.pe/index.php/SAVNET/lago-perubase-conida-astrofisica



ASIA-PACIFIC GROUND-BASE OPTICAL SATELLITE OBSERVATION SYSTEM (APOSOS)



The project aims to develop a unified space observation network based on optical trackers in APSCO Member States.

The basic objective is to build the network with existing facilities, with the aim of tracking objects and space debris in Low Earth Obit (LEO).



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ASIA-PACIFIC GROUND-BASE OPTICAL SATELLITE OBSERVATION SYSTEM (APOSOS)



The optional objective is to bring new facilities, and extends the ability to track objects and space debris in MEO and GEO (peaceful purposes).

The network is for the purpose of establishing the infrastructure for a collision avoidance early warning service in the future.



http://www.conida.gob.pe/index.php/SAVNET/aposos



e-CALLISTO NETWORK





Compound

- A stronomical
- L ow-Cost
- L ow-Frequency
- I nstrument for
- S pectroscopy and
- T ransportable
- **O** bservatory



The main applications are observation of solar radio bursts and rfimonitoring for astronomical science, education and outreach

Map of current distribution of Callisto instruments in February 2015

CALLISTO spectrometer is part of e-CALLISTO network trought implementation program IHY / UNBSSI and ISWI instruments, was financially supported by SNF, SSAA, NASA, Institute for Astronomy and North-South Center of ETH Zurich and a few private sponsors and work bandwidth frequency (45 to 870 MHz).

http://www.conida.gob.pe/index.php/SAVNET/red-e-callisto



SOLAR MONITORING AND PREDICTION CENTER







SPACE WEATHER OBSERVATORY: FACILITIES FOR THE SPACE NAVIGATION 2. SPACE WEATHER OBSERVATORY

3. FACILITIES FOR THE SPACE NAVIGATION



CRIRP PROJECT



Research on Atmospheric Effects on Ionospheric Modeling through Study of Radio Wave Propagation and Solar Activity. The leading institute is China Research Institute of Radio wave Propagation (CRIRP).



The implementation of this project rise in to try to provide more accurate and suitable ionospheric models for the Asia Pacific region, with this it can be used in communication systems and satellite navigation systems by making use of data available in this area.



TEC/scintillation monitor - Antenna for GNSS signal receiving



SPACE WEATHER OBSERVATORY: FACILITIES FOR THE SPACE NAVIGATION 3. FACILITIES FOR SPACE NAVIGATION



CRIRP PROJECT



Main specification of the monitor:

- Signal receiving: GNSS carrier phase and amplitude variation
- Dynamic range: better than 30dB
- Sampling rate of original data: 20 Hz
- Resolution of output: no more than 1dB



TEC/scintillation monitor

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SPACE WEATHER – APOSOS/CRIRP



APOSOS



CRIRP



International GNSS Monitoring and Assessment iGMA Project



OVERALL OBJECTIVE:

Introduce and promote GNSS research and applications to achieve GNSS technical collaboration among the APSCO Member States.

This project is designed to:

- build GNSS data collection points among member states of APSCO
- set up a data analysis system and carry on monitoring
- assessment of GNSS service performance and modeling of ionosphere delay, etc.



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International GNSS Monitoring and Assessment iGMA Project





Software

Data Analysis







International GNSS Monitoring and Assessment iGMA Project



2018 Working Plan

- 1. Complete the site investigation of Peru (March-April)
- 2. Prepare the software training materials (March-April)
- 3. Software training and application extension (March-April, organized by APSCO)
- 4. Prepare the manual of receiver installation, operation and maintenance in May
- 5. Site construction, site operation and on-site maintenance training: April-June Peru
- 6. Site commissioning (at least one month after site installation) (finish in November)
- 7. Data analysis (finish in November)
- 8. Project acceptance (end of February-March, 2019).



4. CONCLUSIONS

CONCLUSIONS



- CONIDA has been developing the space observation center in the city of Huancayo, which will later be moved to the south of the country, in the city of Moquegua.
- CONIDA has been participating in different projects related to space navigation, such as CRIRP and iGMA.



AGENCIA ESPACIAL DEL PERU CONIDA

Ciencia y Tecnología espacial para el desarrollo

